Deep Space Exploration Society

Science Meeting 11-26-2018

Agenda

- Dayton Jones Brief
- Galactic Rotation Rate Results
- Complete 60-ft antenna Source Observations

October 16-17, 2018 Observation Trip

Observation Trip 11-16-18

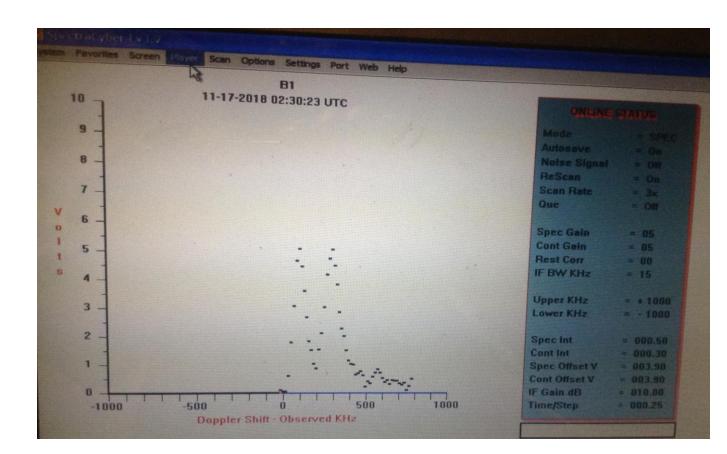
Observers: Richard Russel **Receiver:** SpectraCyber with 1420 Mhz. Cavity Filter on 60-foot dish **Pointing System:** System 1

- **Goals:** 1) Take Milky Way rotation rate data
 - 2) Observe sources with known Jansky signal levels to determine if a signal calibration level can be estimated





Equipment Setup



Dish: 60 ft. Filter: Cavity Filter Receiver: SpectraCyber

SpectraCyber Setup: (see photo) Gain: 5 SPEC Integration: 0.5 Second IF BW: 15 kHz SPEC Offset V: 3.9V* IF Gain dB: 10 Time/Step: 0.25

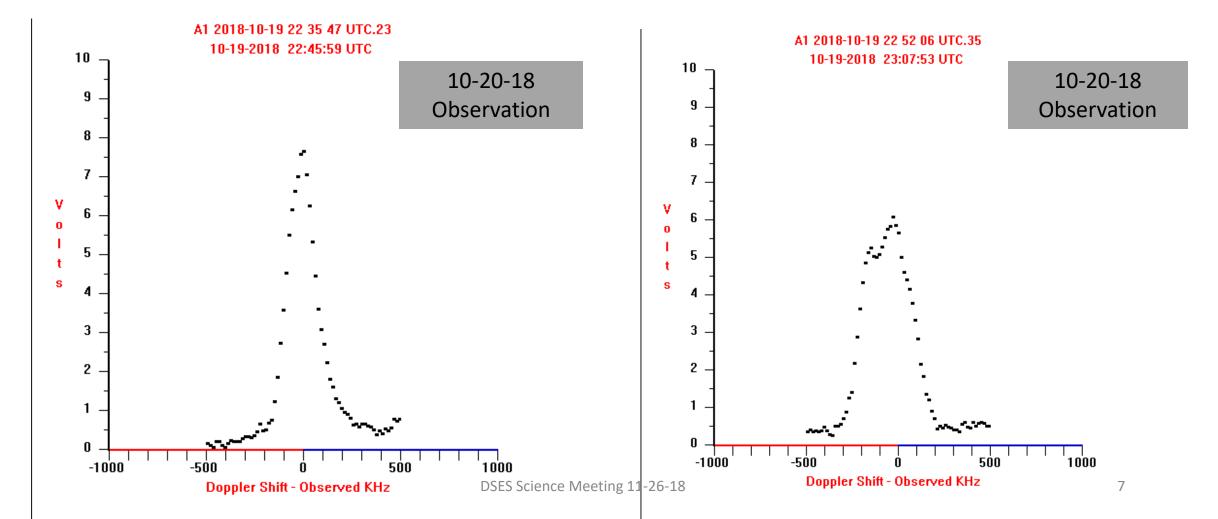
*SPEC Offset was changed at time to maintain the peak signal below 10 volts.

Galactic Rotation Rate Observations

• There was a delay in observation due to antenna and pointing system startup issues. The Galactic plane observations started with b=0, L=55 degrees.

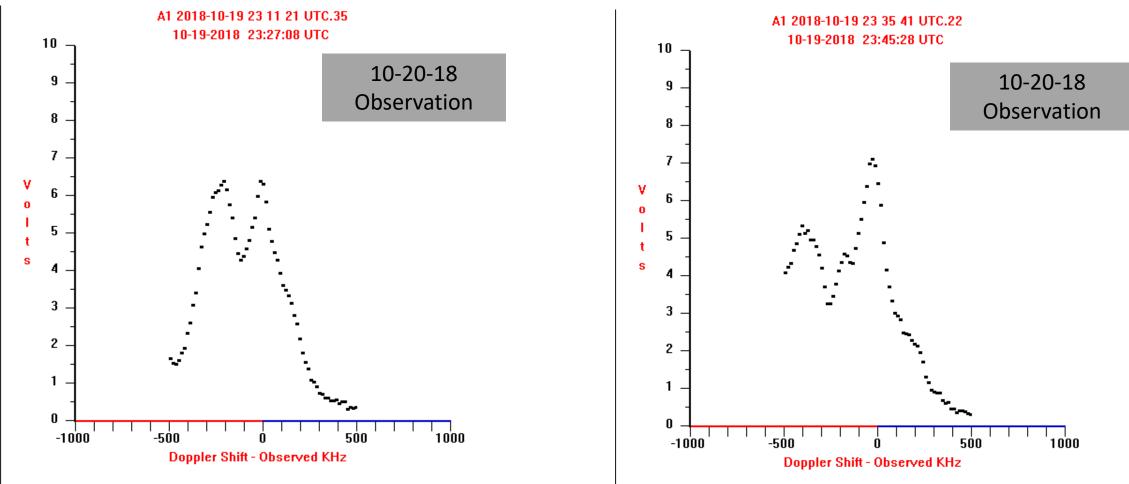
Galactic Lat 0 Long 0 RA 17h45m37s DEC -28d56m10s

Galactic Lat 0 Long 10 RA 18h07m46s DEC -20d17m24s



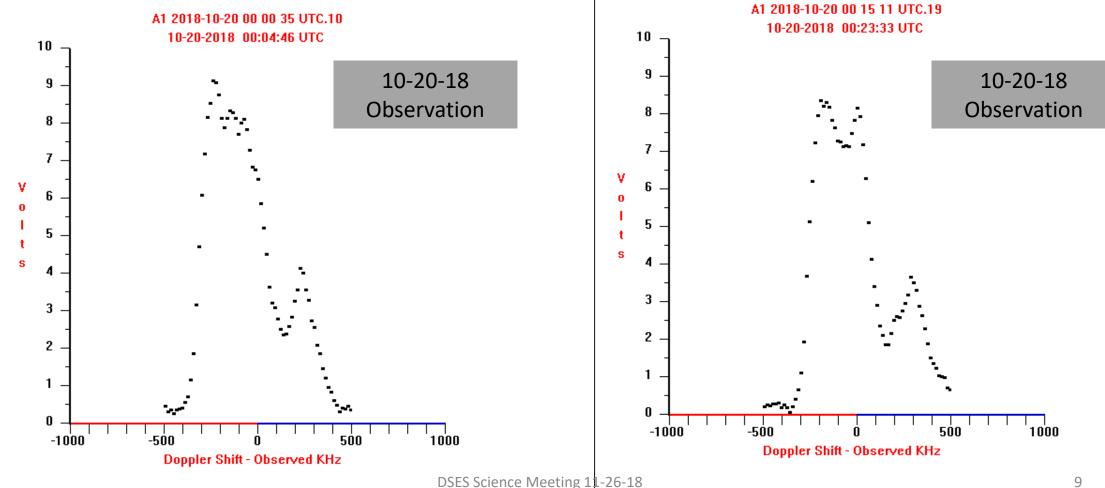
Galactic Lat 0 Long 20 RA 18h27m32s DEC -11d29m19s

Galactic Lat 0 Long 30 RA 18h46m05s DEC -02d36m33s

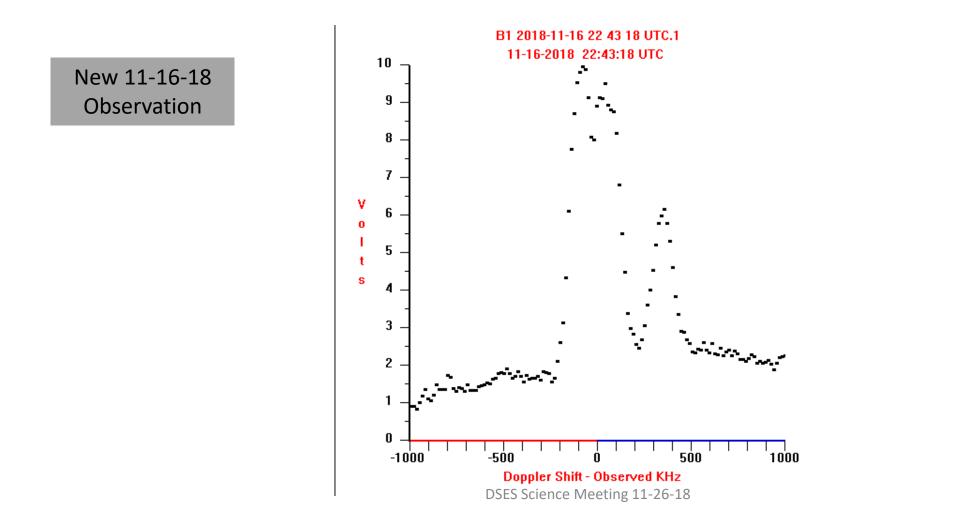


Galactic Lat 0 Long 40 RA 19h04m23s DEC 06d17m14s

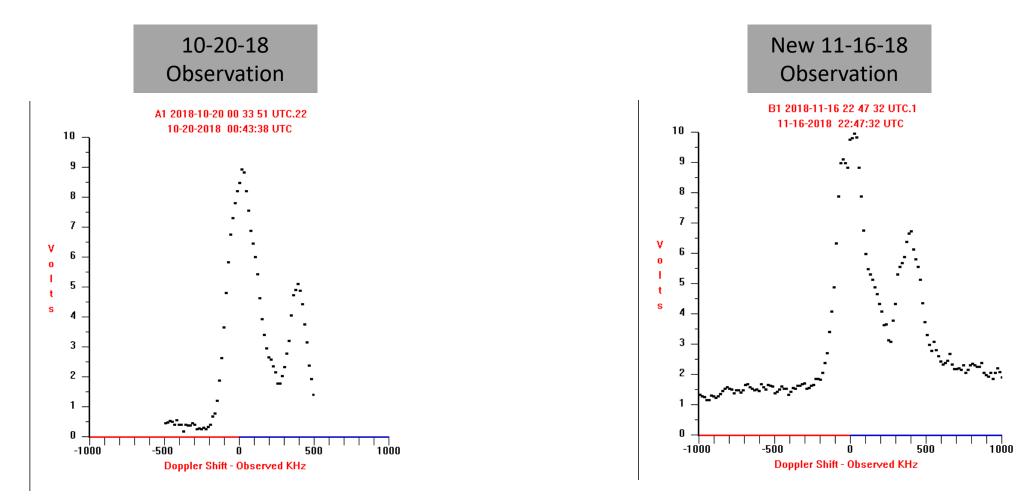
Galactic Lat 0 Long 50 RA 19h23m19s DEC 15d08m33s



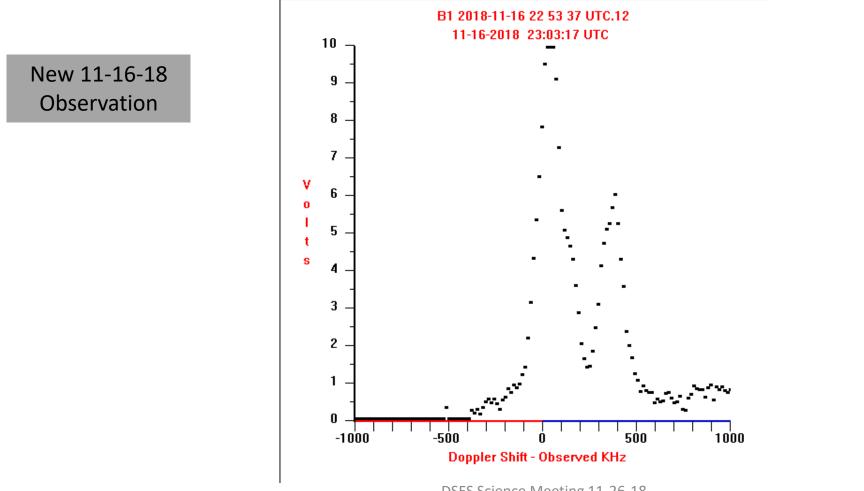
Galactic Lat 0 Long 55 RA 19h33m29s DEC 19d32m04s



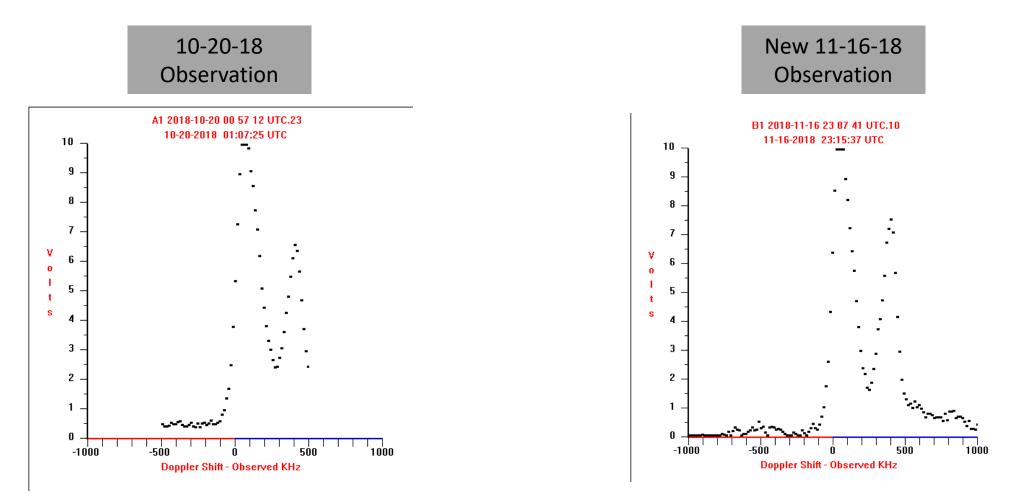
Galactic Lat 0 Long 60 RA 19h43m54s DEC 23d53m25s



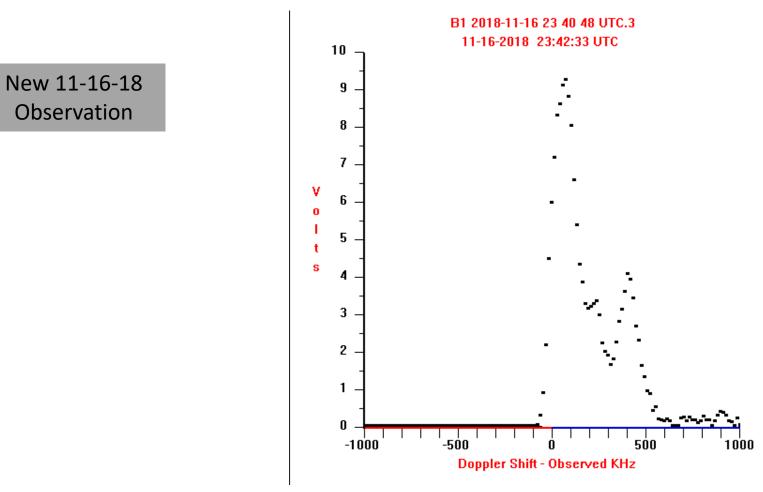
Galactic Lat 0 Long 65 RA 19h55m13s DEC 28d11m52s



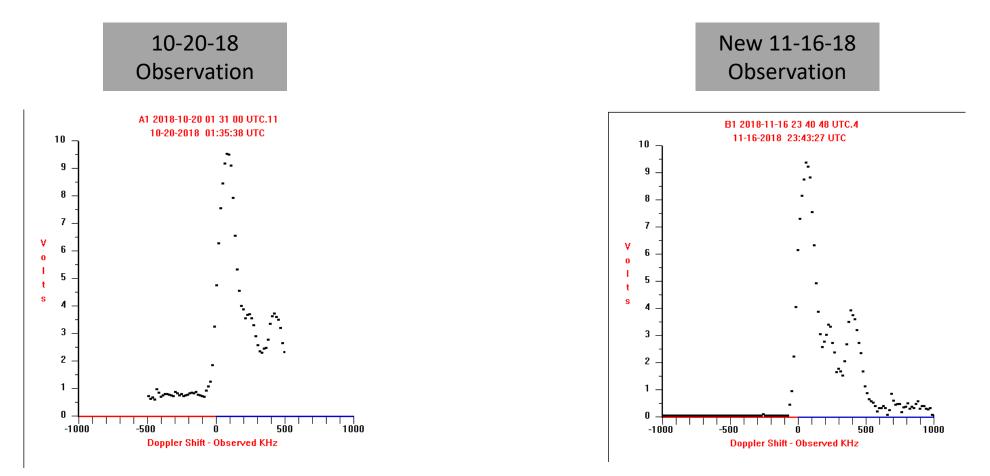
Galactic Lat 0 Long 70 RA 20h07m28s DEC 32d26m33s



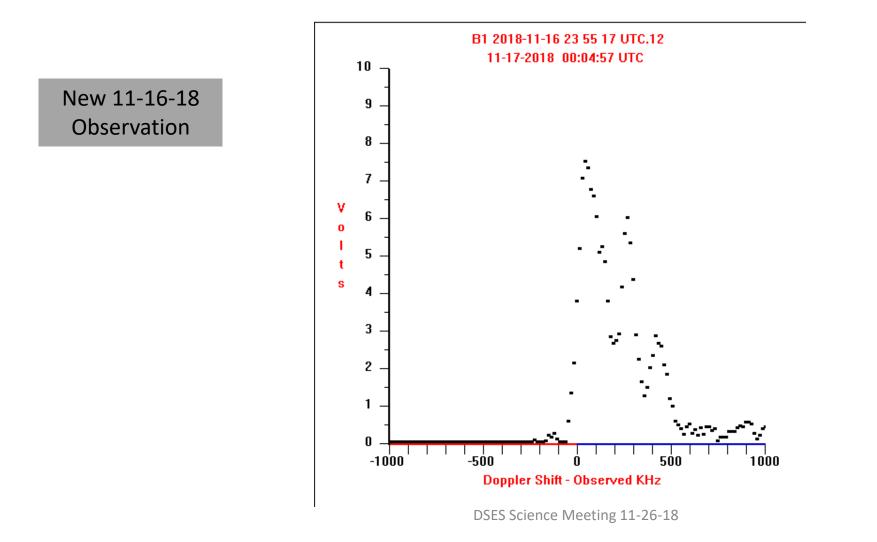
Galactic Lat 0 Long 75 RA 20h20m55s DEC 36d36m20s



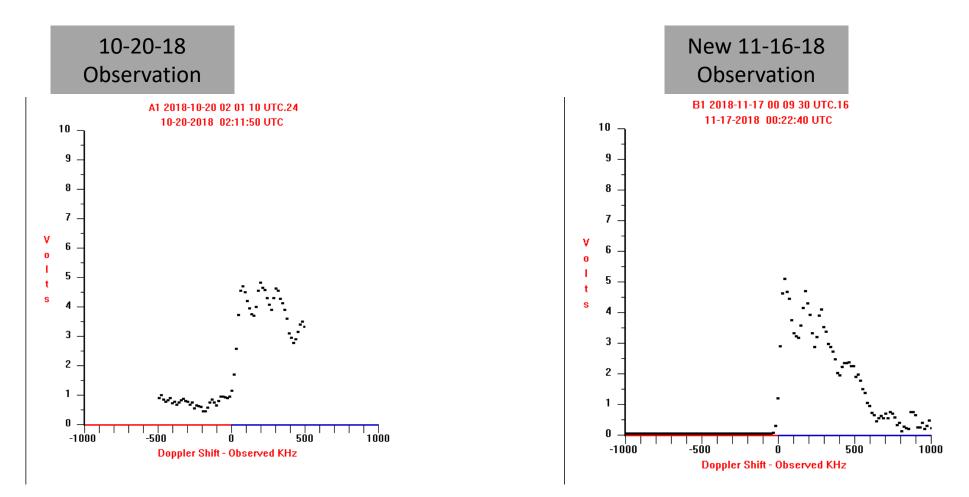
Galactic Lat 0 Long 80 RA 20h35m53s DEC 40d39m49s



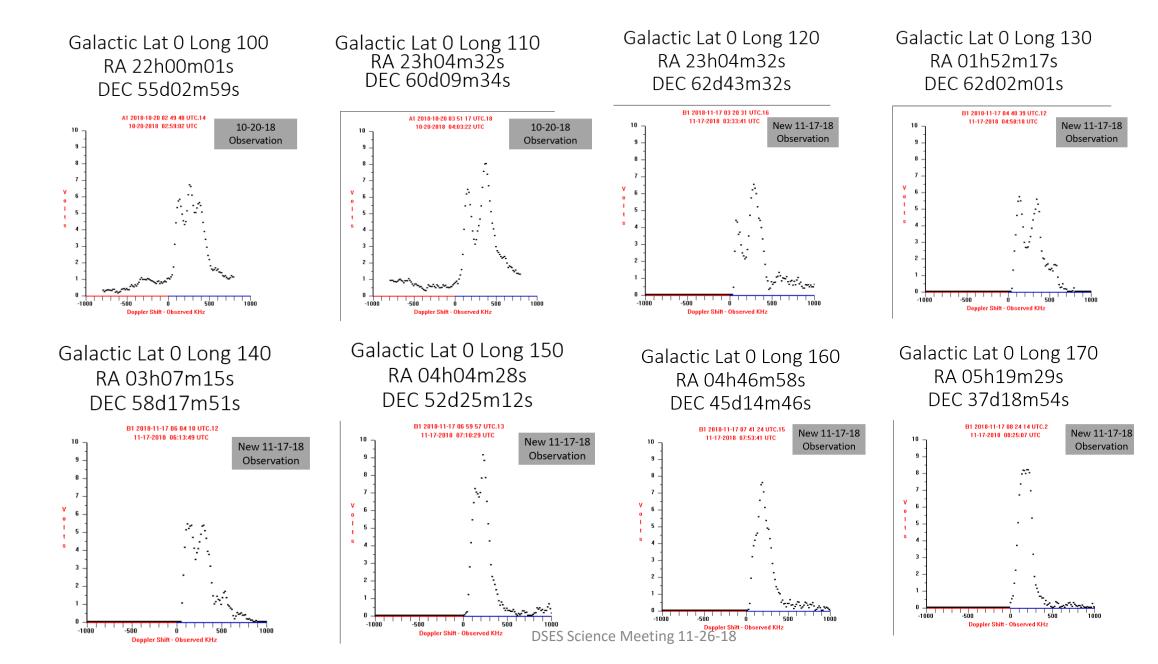
Galactic Lat 0 Long 85 RA 20h52m46s DEC 44d35m07s



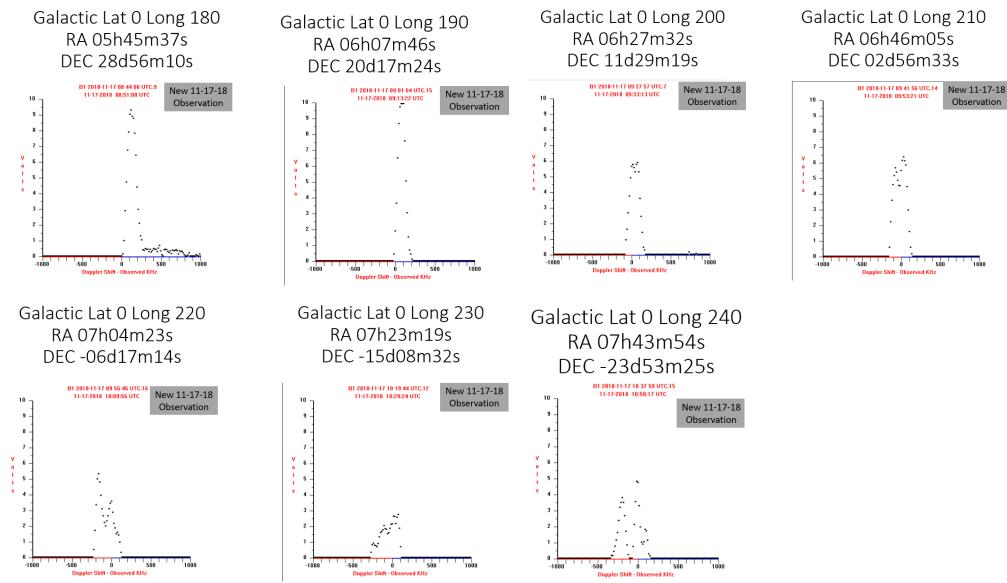
Galactic Lat 0 Long 90 RA 21h12m01s DEC 48d19m46s



Quadrant II & III Observations

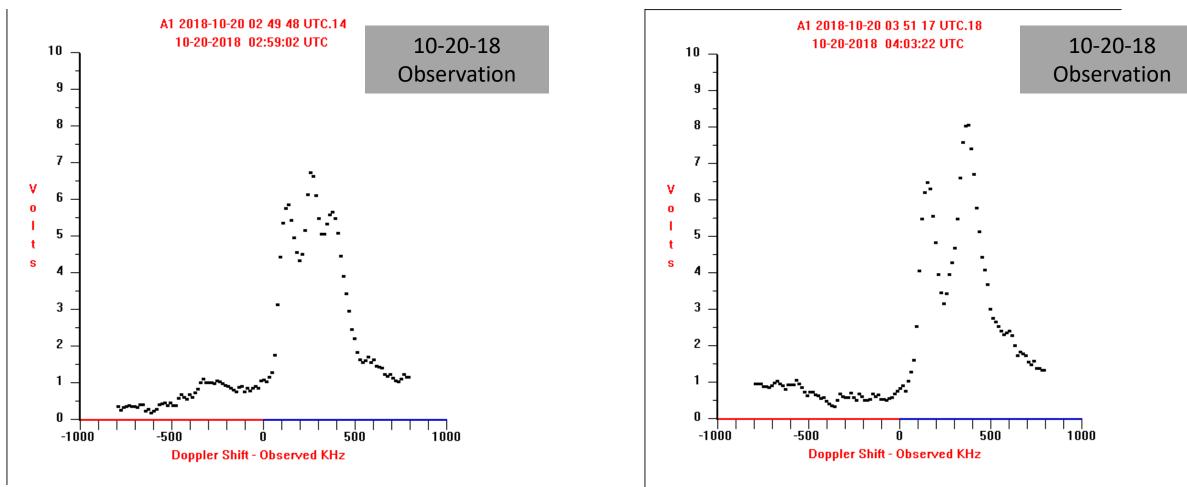


Quadrant III Observations



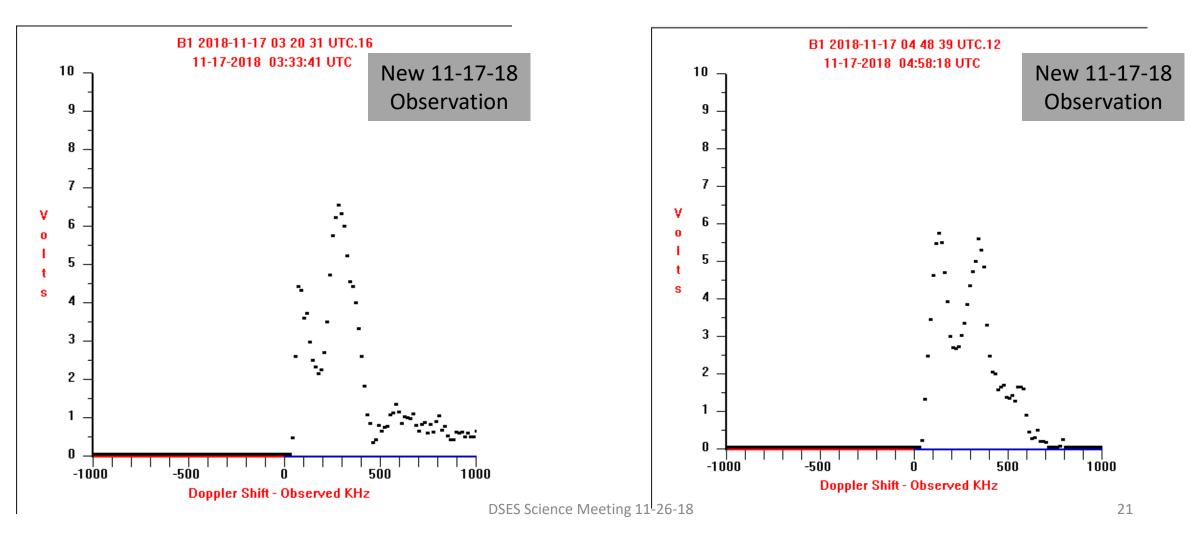
Galactic Lat 0 Long 100 RA 22h00m01s DEC 55d02m59s

Galactic Lat 0 Long 110 RA 23h04m32s DEC 60d09m34s



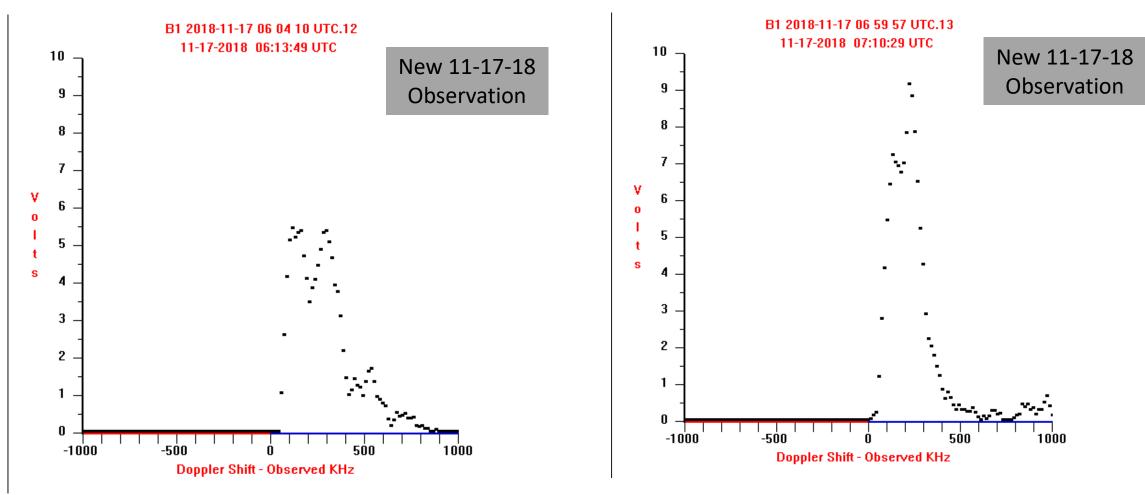
Galactic Lat 0 Long 120 RA 23h04m32s DEC 62d43m32s

Galactic Lat 0 Long 130 RA 01h52m17s DEC 62d02m01s



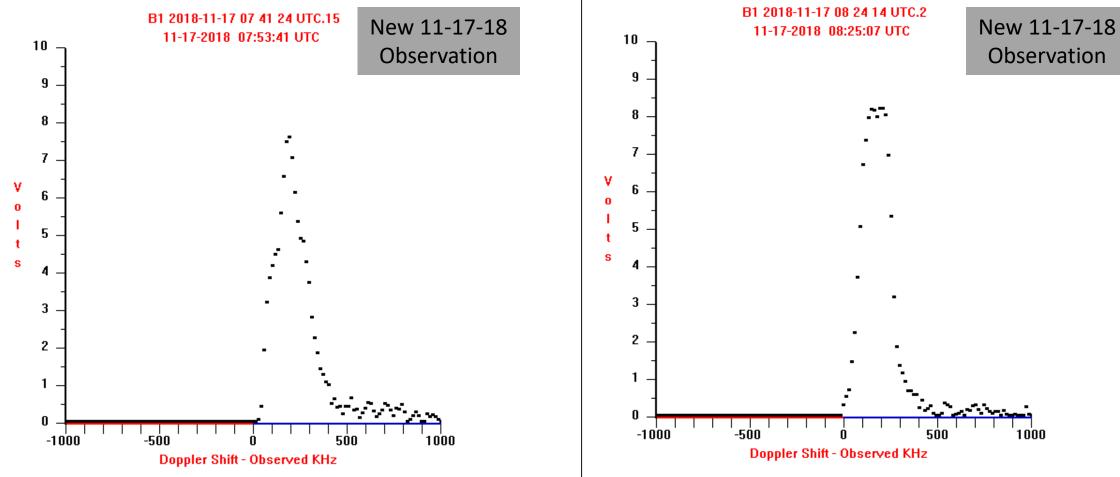
Galactic Lat 0 Long 140 RA 03h07m15s DEC 58d17m51s

Galactic Lat 0 Long 150 RA 04h04m28s DEC 52d25m12s



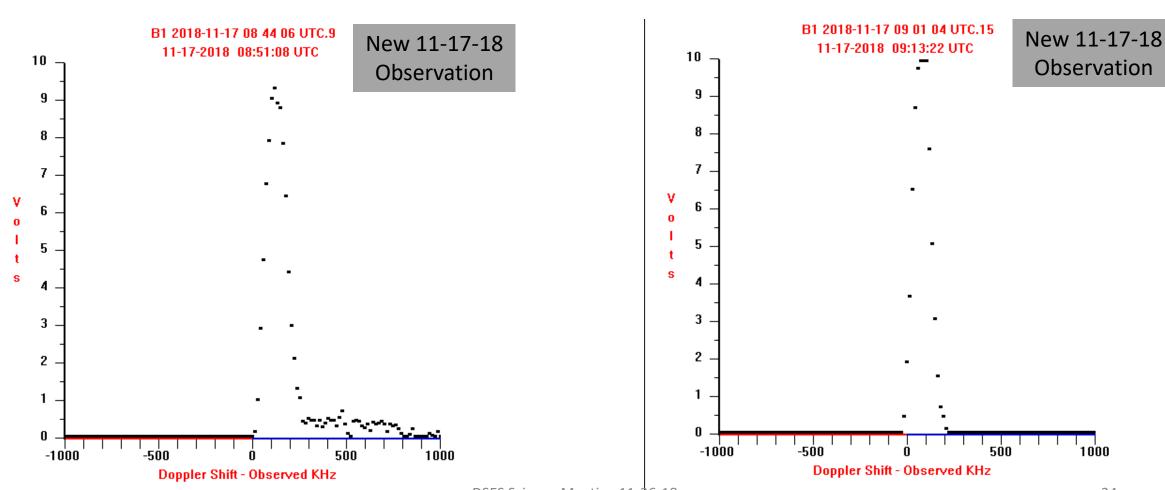
Galactic Lat 0 Long 160 RA 04h46m58s DEC 45d14m46s

Galactic Lat 0 Long 170 RA 05h19m29s DEC 37d18m54s



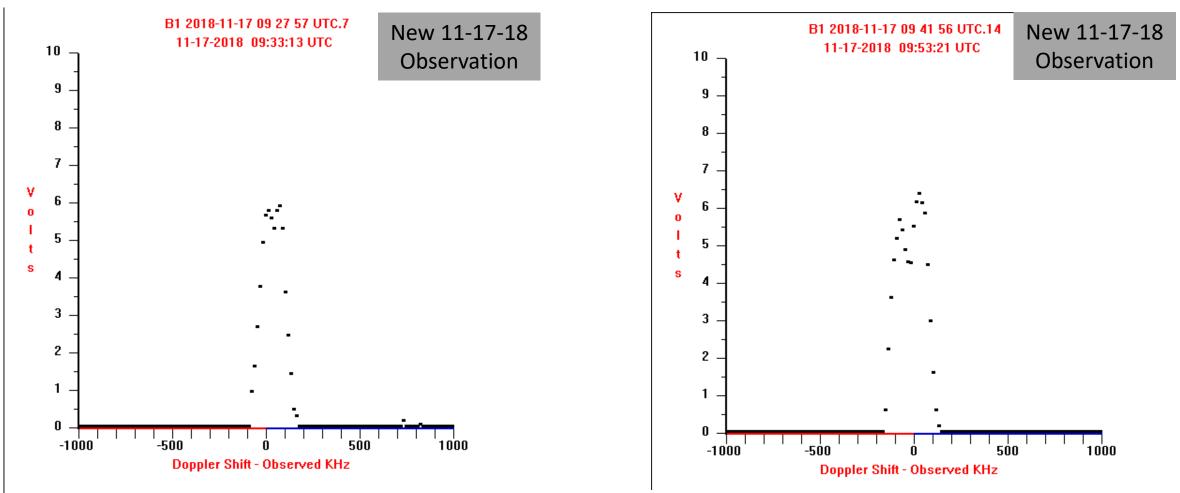
Galactic Lat 0 Long 180 RA 05h45m37s DEC 28d56m10s

Galactic Lat 0 Long 190 RA 06h07m46s DEC 20d17m24s



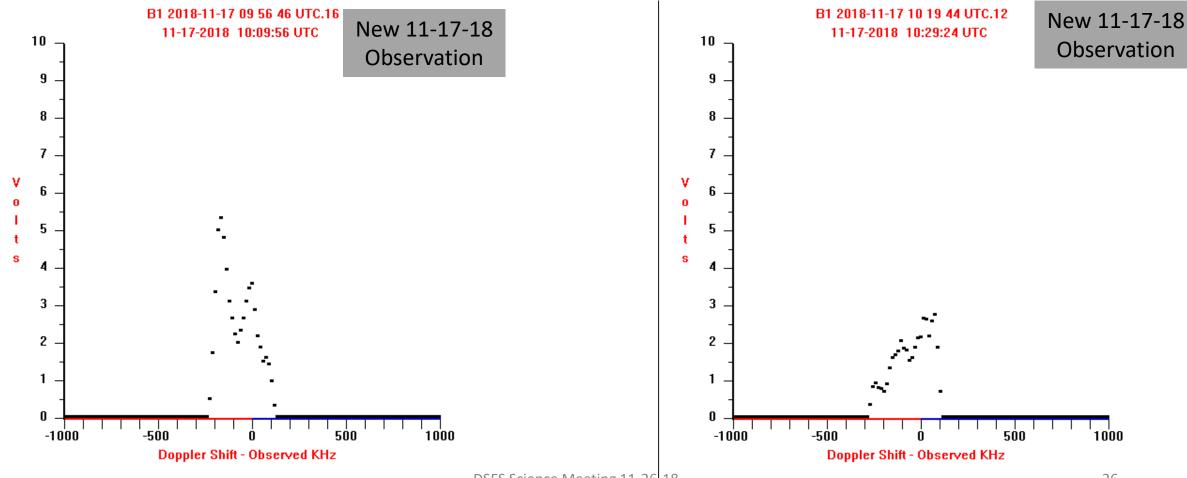
Galactic Lat 0 Long 200 RA 06h27m32s DEC 11d29m19s

Galactic Lat 0 Long 210 RA 06h46m05s DEC 02d56m33s

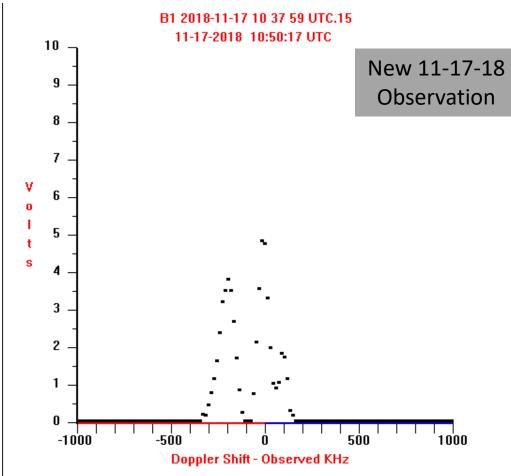


Galactic Lat 0 Long 220 RA 07h04m23s DEC -06d17m14s

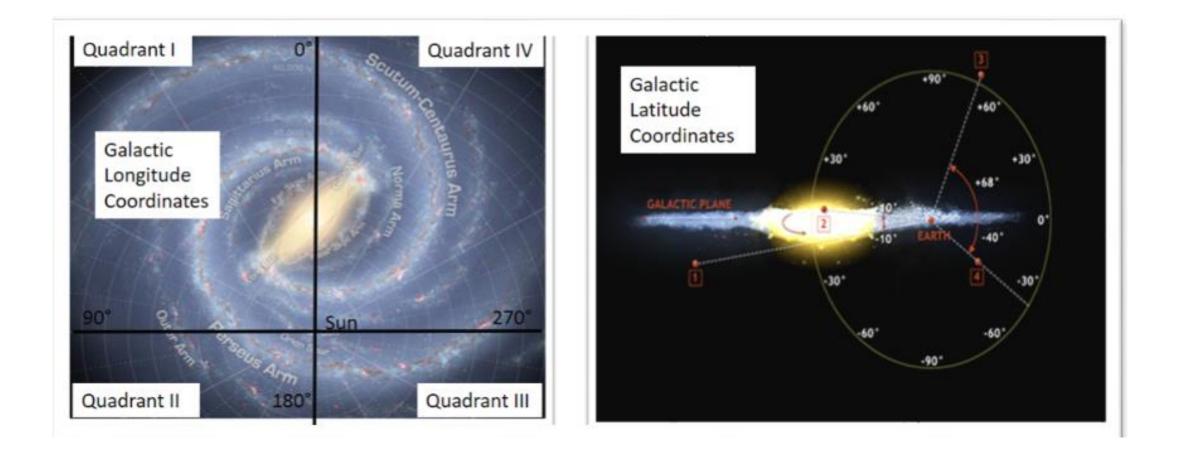
Galactic Lat 0 Long 230 RA 07h23m19s DEC -15d08m32s



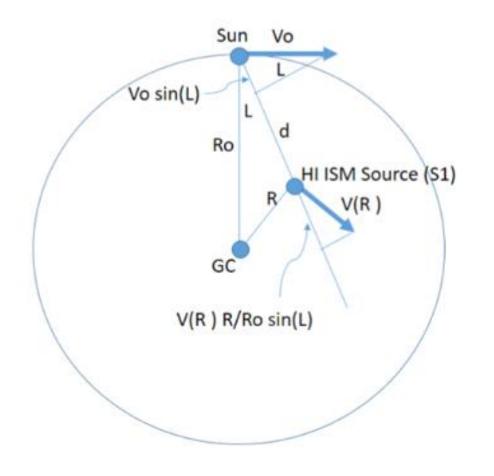
Galactic Lat 0 Long 240 RA 07h43m54s DEC -23d53m25s



Milky Way Longitude and Latitude Coordinate System

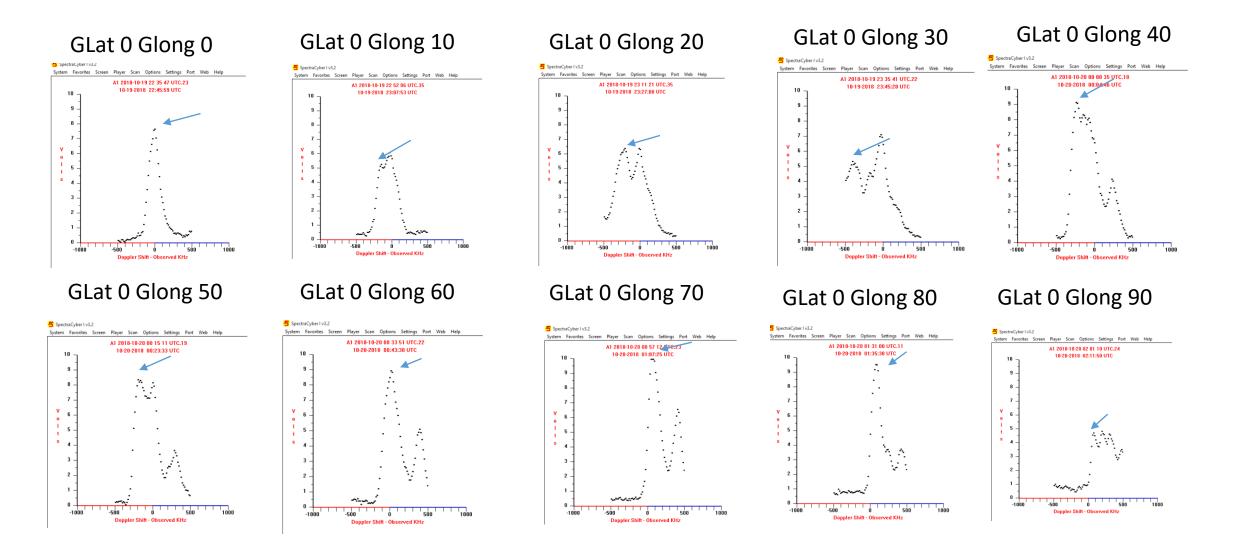


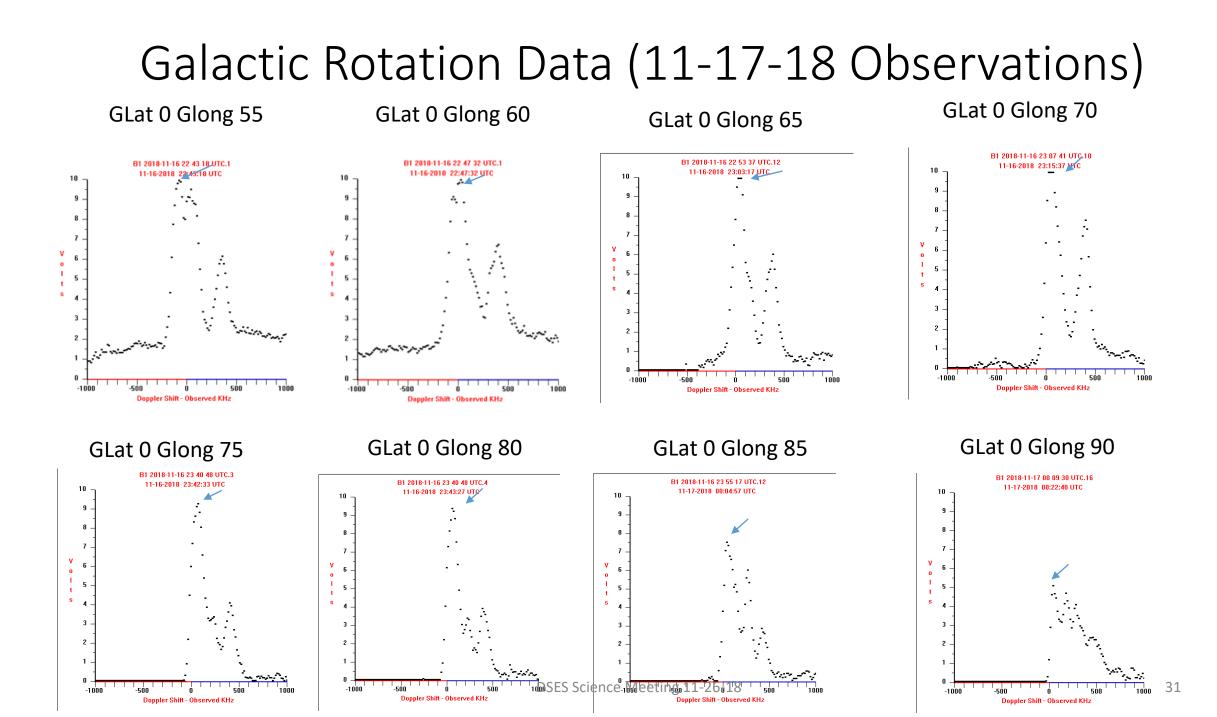
Rotation Rate Geometry (R< Ro)



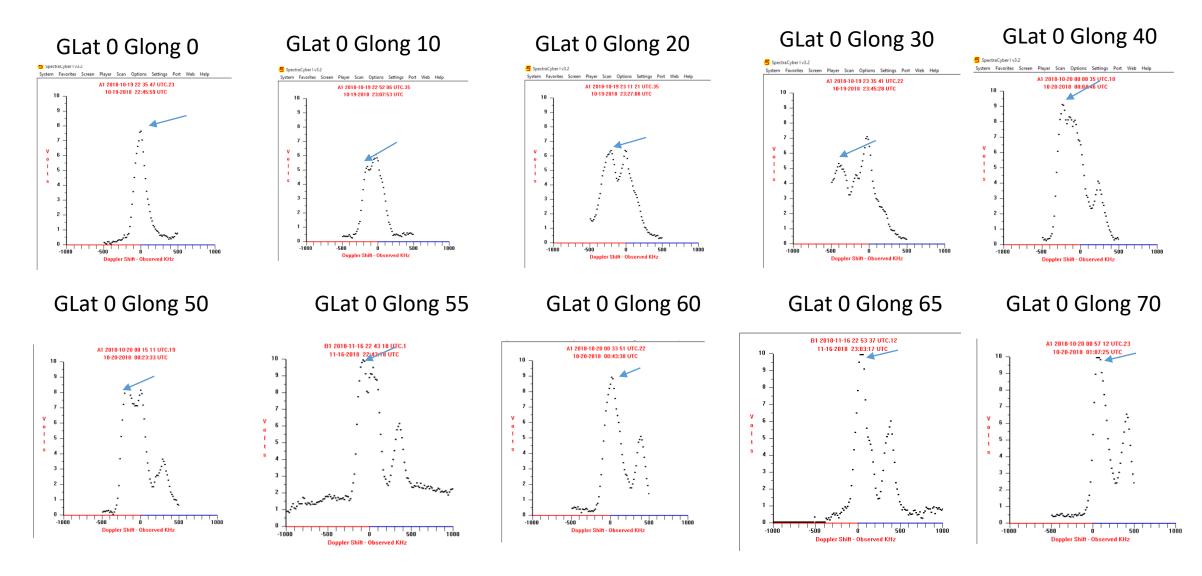
$$Vr = V(R)\frac{Ro}{R}sin(L) + Vo sin(L)$$

Galactic Rotation Data (10-20-18 Observations)

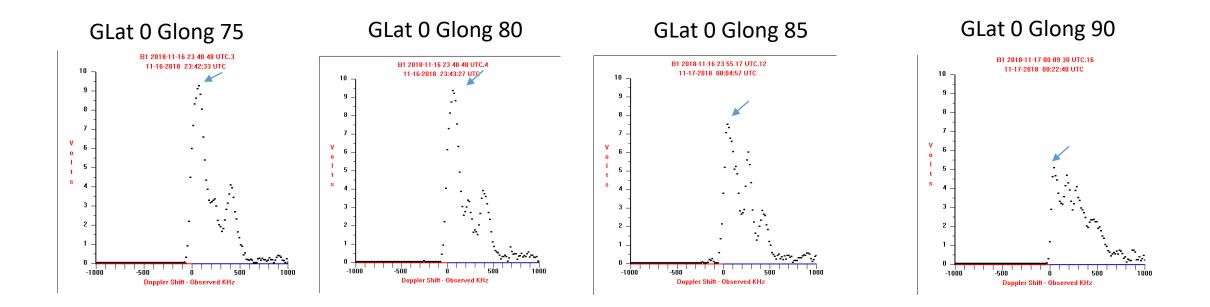




Galactic Rotation Data



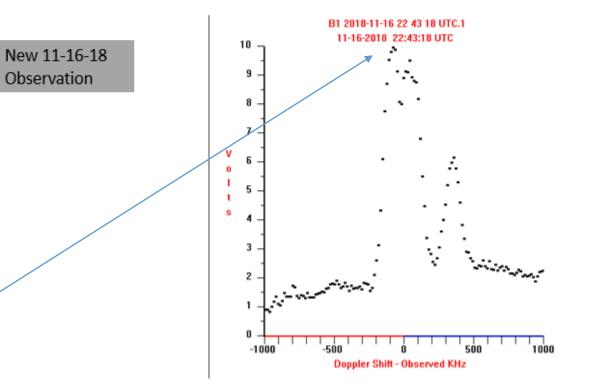
Galactic Rotation Data



Calculating Velocities

	А	В	С	D	E	
1	HZ	HZ	s 11 16 2018 22 43	18 0.250000 -1000 1	s 10 20 2018 0 23 33	0
2	5.0000	1,420,404,751	0.9399390	Offset	Delta	C
3	Nov 17 2018	1,420,404,756	0.9399390	5.02512540000	0.0000000000	
4		1,420,404,761	0.9399390	Base Frequency	1,420,405,751	
5		1,420,404,766	0.9350560	Min Peak	1,420,405,671	
6		1,420,404,771	0.9350560	Doppler (km/Sec)	-16.88	
7		1,420,404,776	0.9350560			
8		1,420,404,781	0.8520490	-204.73		
9		1,420,404,786	0.8520490	-203.67		
10		1,420,404,791	0.8520490	-202.62		
11		1,420,404,796	1.0351540	-201.56		
12		1,420,404,801	1.0351540	-200.51		
			·		·	
	А	В	С	D	E	
185		1,420,405,666	9.9902090	-17.94		
186		1,420,405,671	9.9902090	-16.88	Min Peak	
187		1,420,405,676	9.9902090	-15.83		
188		1,420,405,681	9.9120840	-14.77		

Galactic Lat 0 Long 55 RA 19h33m29s DEC 19d32m04s



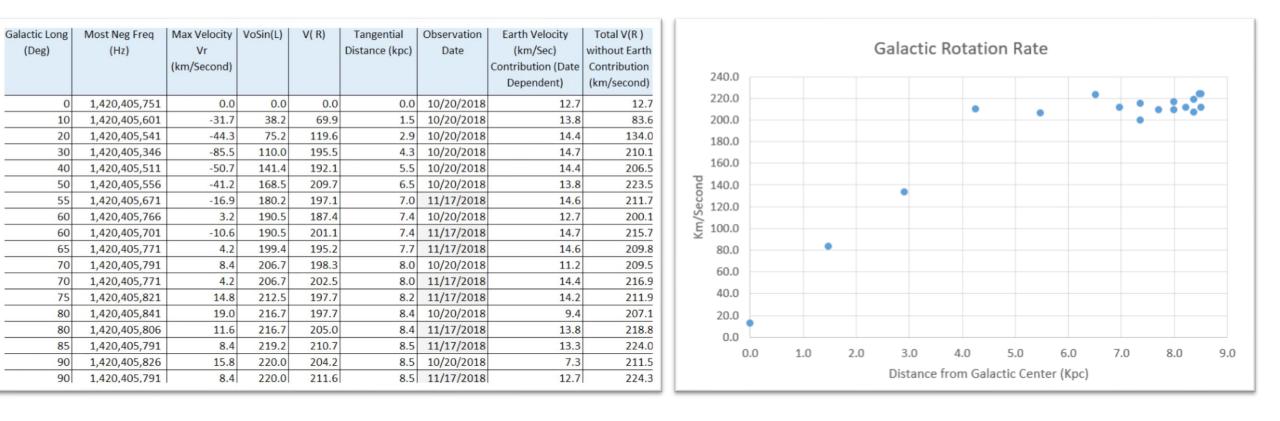
Frequency Used

Observation

Quadrant I: Earth Orbital Velocity Corrections

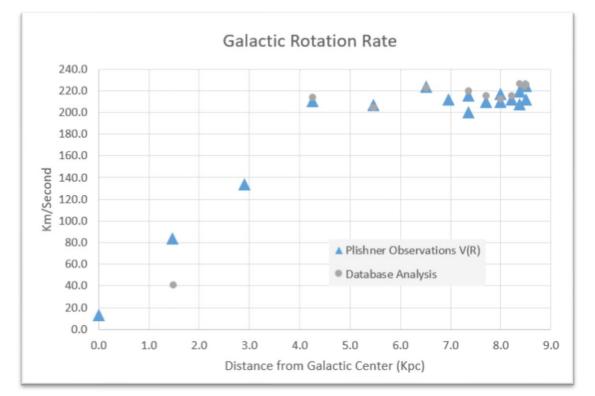
Parameter	Value	Units	Date	Days from Summer Solstice (22 June)	Earth Deg from GC	Velo Lon						Velocity Long 25													
Velocity (Vr)) km/s	22-Jun	0.0	o		0.0	-1.3	-2.5	-3.8	-5.0	-7.3	-7.3	-8.4	-9.4	-10.4	-11.2	-12.0	-12.7	-13.3	-13.8	-14.2	-14.4	-14.6	-14.7
Galactic Long	(Degrees	2-Jul	10.1	10		2.5	1.3	0.0	-1.3	-2.5	-5.0	-5.0		-7.3	-8.4	-9.4		-11.2		-12.7	-13.3	-13.8	-14.2	-14.4
Circular Velocity	29.5	km/s	12-Jul	20.3	20		5.0	3.8	2.5	1.3	0.0	-2.5	-2.5			-6.2	-7.3	-8.4	-9.4	-10.4	-11.2	-12.0	-12.7	-13.3	-13.8
Earth Orbit Degrees from GC	() degrees	22-Jul	30.4	30		7.3	6.2	5.0	3.8	2.5	0.0	0.0		-2.5	-3.8	-5.0	-6.2	-7.3	-8.4	-9.4	-10.4	-11.2	-12.0	-12.1
_			30-Jul	40.6	40		9.4	8.4	7.3	6.2	5.0	2.5	2.5	1.3	0.0	-1.3	-2.5	-3.8	-5.0	-6.2	-7.3	-8.4	-9.4	-10.4	-11.2
Ecliptic Vs Gal Plane	60.2	degrees	9-Aug	50.7	50		11.2	10.4	9.4	8.4	7.3	5.0	5.0	3.8	2.5	1.3	0.0	-1.3	-2.5	-3.8	-5.0	-6.2	-7.3	-8.4	-9.4
			19-Aug	60.8	60		12.7	12.0	11.2	10.4	9.4	7.3	7.3	6.2	5.0	3.8	2.5	1.3			-2.5	-3.8		-6.2	-7.3
			30-Aug	71.0	70		13.8	13.3	12.7	12.0	11.2	9.4	9.4	8.4	7.3	6.2	5.0	3.8	2.5	1.3	0.0	-1.3	-2.5	-3.8	
			9-Sep	81.1	80		14.4	14.2	13.8	13.3	12.7	11.2	11.2	10.4	9.4	8.4			5.0	3.8	2.5	1.3	0.0	-1.3	-2.5
			19-Sep	91.3				14.6	14.4	14.2	13.8	12.7	12.7		11.2	10.4	9.4	8.4			5.0		2.5		
			29-Sep	101.4	100			14.6	14.7	14.6	14.4	13.8	13.8	13.3	12.7	12.0	11.2		9.4		7.3	6.2	5.0	3.8	
			9-Oct	111.5				14.2	14.4	14.6	14.7	14.4	14.4		13.8	13.3					9.4		7.3		5.
			19-Oct	121.7	120			13.3	13.8	14.2	14.4		14.7		14.4	14.2			12.7		11.2	10.4			7.
			29-Oct	131.8	130			12.0	12.7	13.3	13.8	14.4	14.4		14.7	14.6			13.8		12.7	12.0	11.2		9.
			8-Nov	141.9			9.4	10.4	11.2	12.0	12.7	13.8	13.8		14.4	14.6	14.7	14.6	14.4			13.3		12.0	11.
			19-Nov	152.1	150		7.3	8.4	9.4	10.4	11.2	12.7	12.7		13.8	14.2	14.4	14.6	14.7			14.2	13.8	13.3	12.
			29-Nov	162.2	160		5.0	6.2	7.3	8.4	9.4		11.2		12.7	13.3			14.4		14.7	14.6	14.4	14.2	13.
			9-Dec	172.4	170		2.5	3.8	5.0	6.2	7.3		9.4	10.4	11.2	12.0	12.7		13.8		14.4	14.6	14.7	14.6	14.
			19-Dec	182.5	180		0.0	1.3	2.5	3.8	5.0	7.3	7.3	8.4	9.4	10.4	11.2	12.0	12.7	13.3	13.8	14.2	14.4	14.6	14.
			29-Dec	192.6	190			-1.3	0.0	1.3	2.5	5.0	5.0		7.3	8.4	9.4	10.4	11.2	12.0	12.7	13.3	13.8	14.2	14.
			8-Jan	202.8	200			-3.8	-2.5	-1.3	0.0	2.5	2.5	3.8	5.0			8.4	9.4	10.4	11.2	12.0	12.7	13.3	13.
			18-Jan	212.9	210			-6.2	-5.0	-3.8		0.0	0.0			3.8			7.3		9.4		11.2	12.0	12.
			29-Jan	223.1	220			-8.4	-7.3	-6.2	-5.0		-2.5			1.3					7.3	8.4	9.4	10.4	11.
			8-Feb	233.2	230		-11.2 -	-10.4	-9.4	-8.4	-7.3	-5.0	-5.0		-2.5	-1.3				3.8	5.0	6.2	7.3	8.4	9.
			18-Feb	243.3	240			-12.0	-11.2	-10.4	-9.4		-7.3		-5.0	-3.8						3.8	5.0		7.
			28-Feb	253.5	250	-	13.8 -	-13.3	-12.7	-12.0	-11.2	-9.4	-9.4		-7.3	-6.2				-1.3				3.8	
			10-Mar	263.6	260			-14.2	-13.8	-13.3	-12.7	-11.2	-11.2		-9.4	-8.4	-7.3		-5.0		-2.5			1.3	
			20-Mar	273.8	270		14.7 -	-14.6	-14.4	-14.2	-13.8	-12.7	-12.7		-11.2	-10.4	-9.4	-8.4	-7.3	-6.2	-5.0	-3.8		-1.3	
			30-Mar	283.9	280		14.4 -	-14.6	-14.7	-14.6	-14.4	-13.8	-13.8	-13.3	-12.7	-12.0	-11.2	-10.4	-9.4	-8.4	-7.3	-6.2	-5.0	-3.8	-2.
			10-Apr	294.0	290			-14.2	-14.4	-14.6	-14.7	-14.4	-14.4		-13.8	-13.3		-12.0	-11.2		-9.4	-8.4			-5.
			20-Apr	304.2	300			-13.3	-13.8	-14.2	-14.4		-14.7			-14.2			-12.7		-11.2	-10.4		-8.4	-7.
			30-Apr	314.3	310			-12.0	-12.7	-13.3	-13.8	-14.4	-14.4		-14.7	-14.6			-13.8		-12.7	-12.0	-11.2	-10.4	-9.
			10-May	324.4	320		-9.4 -	-10.4	-11.2	-12.0	-12.7	-13.8	-13.8	-14.2	-14.4	-14.6	-14.7	-14.6	-14.4	-14.2	-13.8	-13.3	-12.7	-12.0	-11.
			20-May	334.6	330		-7.3	-8.4	-9.4	-10.4	-11.2	-12.7	-12.7	-13.3	-13.8	-14.2	-14.4	-14.6	-14.7	-14.6	-14.4	-14.2	-13.8	-13.3	-12.
			30-May	344.7	340			-6.2	-7.3	-8.4	-9.4	-11.2	-11.2	-12.0	-12.7	-13.3	-13.8	-14.2	-14.4	-14.6	-14.7	-14.6	-14.4	-14.2	-13.
			9-Jun	354.9	350		-2.5	-3.8	-5.0	-6.2	-7.3	-9.4	-9.4	-10.4	-11.2	-12.0	-12.7	-13.3	-13.8	-14.2	-14.4	-14.6	-14.7	-14.6	-14.
			19-Jun	365.0	360			-1.3	-2.5	-3.8	-5.0	-7.3	-7.3		-9.4	-10.4	-11.2	-12.0	-12.7	-13.3	-13.8	-14.2	∧ -14.4		r:14.

Galactic Rotation Rate Results



Galactic Rotation Rate Results Comparison with Database Analysis

1	A	В	C	F	G	Н	L.	J	K	L	М	N
	Galactic Long	Most Neg Freq	Max Velocity	VoSin(L)	V(R)	Tangential	Observation	Earth Velocity	Total V(R)	Database	Database	Delta
	(Deg)	(Hz)	Vr			Distance (kpc)	Date	(km/Sec)	without Earth	Analysis	V(R)	(km/sec)
			(km/Second)					Contribution	Contribution	Range	(km/s)	
1								(Date Dependent)	(km/second)	(kpc)		
2	0	1,420,405,751	0.0	0.0	0.0	0.0	10/20/2018	12.7	12.7			
3	10	1,420,405,601	-31.7	38.2	69.9	1.5	10/20/2018	13.8	83.6	1.5	41.2	
4	20	1,420,405,541	-44.3	75.2	119.6	2.9	10/20/2018	14.4	134.0			
5	30	1,420,405,346	-85.5	110.0	195.5	4.3	10/20/2018	14.7	210.1	4.3	214.0	-3.
5	40	1,420,405,511	-50.7	141.4	192.1	5.5	10/20/2018	14.4	206.5	5.5	204.4	2
7	50	1,420,405,556	-41.2	168.5	209.7	6.5	10/20/2018	13.8	223.5	6.5	222.5	0.
8	55	1,420,405,671	-16.9	180.2	197.1	7.0	11/17/2018	14.6	211.7			
9	60	1,420,405,766	3.2	190.5	187.4	7.4	10/20/2018	12.7	200.1	7.4	219.5	-19
10	60	1,420,405,701	-10.6	190.5	201.1	7.4	11/17/2018	14.7	215.7			
11	65	1,420,405,771	4.2	199.4	195.2	7.7	11/17/2018	14.6	209.8	7.7	215.4	-5.
2	70	1,420,405,791	8.4	206.7	198.3	8.0	10/20/2018	11.2	209.5	8.0	212.7	-3.
13	70	1,420,405,771	4.2	206.7	202.5	8.0	11/17/2018	14.4	216.9			
4	75	1,420,405,821	14.8	212.5	197.7	8.2	11/17/2018	14.2	211.9	8.2	215.5	-3.
15	80	1,420,405,841	19.0	216.7	197.7	8.4	10/20/2018	9.4	207.1	8.4	226.7	-19.
16	80	1,420,405,806	11.6	216.7	205.0	8.4	11/17/2018	13.8	218.8			
17	85	1,420,405,791	8.4	219.2	210.7	8.5	11/17/2018	13.3	224.0	8.5	224.2	-0.
18	90	1,420,405,826	15.8	220.0	204.2	8.5	10/20/2018	7.3	211.5			1.4
19	90	1,420,405,791	8.4	220.0	211.6	8.5	11/17/2018	12.7	224.3	8.5	226.0	-1.
20											STDev	7.



Mass of Milky Way inside R=8.5kpc

$$M = \frac{V^2 R}{G}$$

$$M = \frac{\left(224.3\frac{km}{s}x\frac{10^3m}{km}\right)^2 (8.5kpc)(\frac{3.09x10^{19}m}{1kpc})}{(6.67x10^{-11}m^3kg^{-1}s^{-2})} = 1.98x10^{41}kg$$

$$M = (1.98x10^{41}kg) \left(\frac{1Msun}{2x10^{30}kg}\right) = 9.9x10^{10}Msun \approx 1.0x10^{11}Msun$$

Based On Galactic Radio Astronomy

$$M \approx 1.0 x 10^{11} M sun \frac{R}{Ro}$$

Complete 60-Foot Antenna Spectracyber Observations

10-19-18 Observation

10

9

8

7

6

5

3

2

1

0

-1000

-500

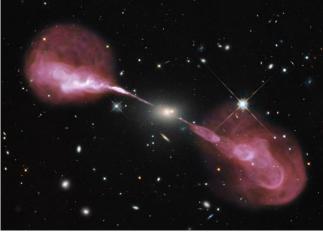
10-19-2018 21:58:15 UTC

Doppler Shift - Observed KHz

500

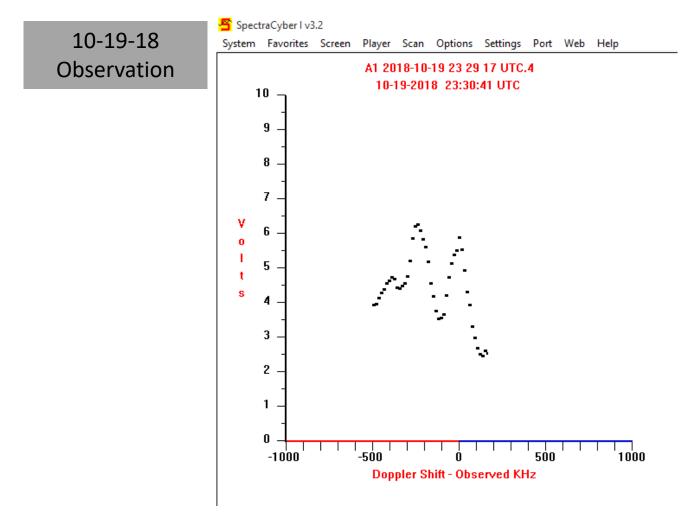
1000

Source: 3C348 (Hercules A) RA 16h48m DEC 05d04m 46.73 JY



<u>APOD: 2012 December 5 - Plasma Jets from Radio Galaxy Hercules A</u> **Explanation:** Why does this galaxy emit such spectacular jets? No one is sure, but it is likely related to an active supermassive black hole at its center. The galaxy at the <u>image center</u>, Hercules A, appears to be a relatively normal <u>elliptical galaxy</u> in visible light. When <u>imaged</u> in <u>radio waves</u>, however, tremendous <u>plasma</u> jets over one million light years long appear. Detailed analyses indicate that the central galaxy, also known as <u>3C 348</u>, is actually over 1,000 times more massive than our Milky Way Galaxy, and the central black hole is nearly 1,000 times more massive than the <u>black hole</u> at our <u>Milky Way's</u> center. <u>Pictured above</u> is a visible light image obtained by the Earth-orbiting <u>Hubble Space Telescope</u> superposed with a radio image taken by the recently upgraded <u>Very Large Array</u> (VLA) of radio telescopes in <u>New Mexico</u>, <u>USA</u>. The physics that creates <u>the jets</u> remains a topic of research with a likely energy source being infalling matter <u>swirling toward</u> the central <u>black hole</u>.

Source: Unknown RA 18h30m DEC -10d10m

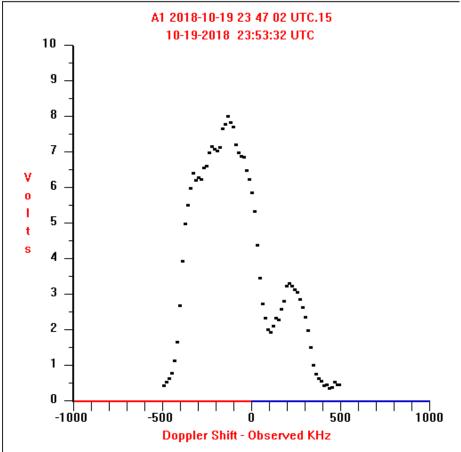


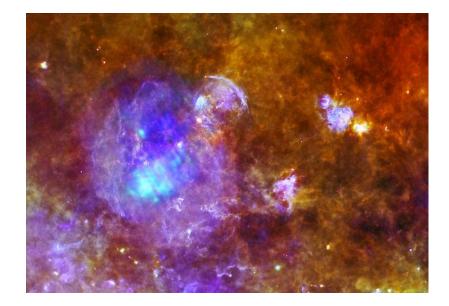
Source: W44 (Supernova Remnant) RA 18h53m30s DEC 1d18m 230 JY

10-19-18 Observation

SpectraCyber I v3.2

System Favorites Screen Player Scan Options Settings Port Web Help





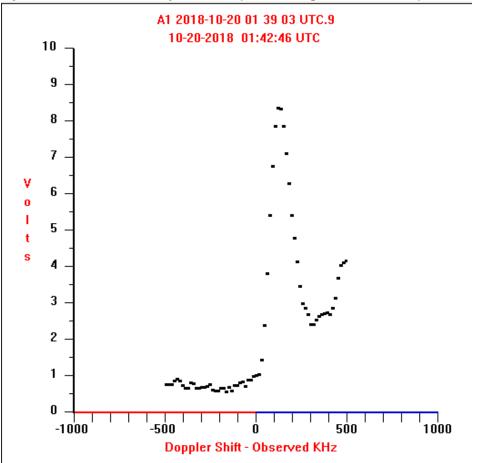
•The aftershock of a stellar explosion rippling through space is captured in this new view of supernova remnant W44, which combines far-infrared and X-ray data from ESA's Herschel and XMM-Newton space observatories.

•W44, located around 10 000 light-years away within a forest of dense star-forming clouds in the constellation of Aquila, the Eagle, is one of the best examples of a supernova remnant interacting with its parent molecular cloud.

Source: HB21 (Supernova remnant) RA 20h43m DEC 50d25m 220 JY

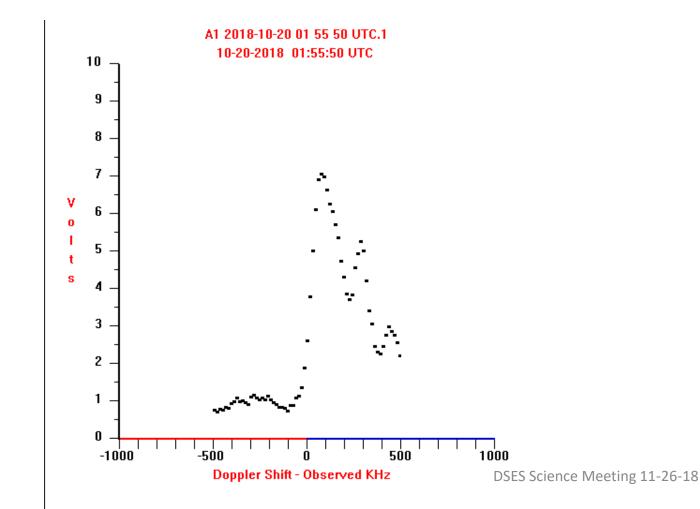
SpectraCyber I v3.2

System Favorites Screen Player Scan Options Settings Port Web Help

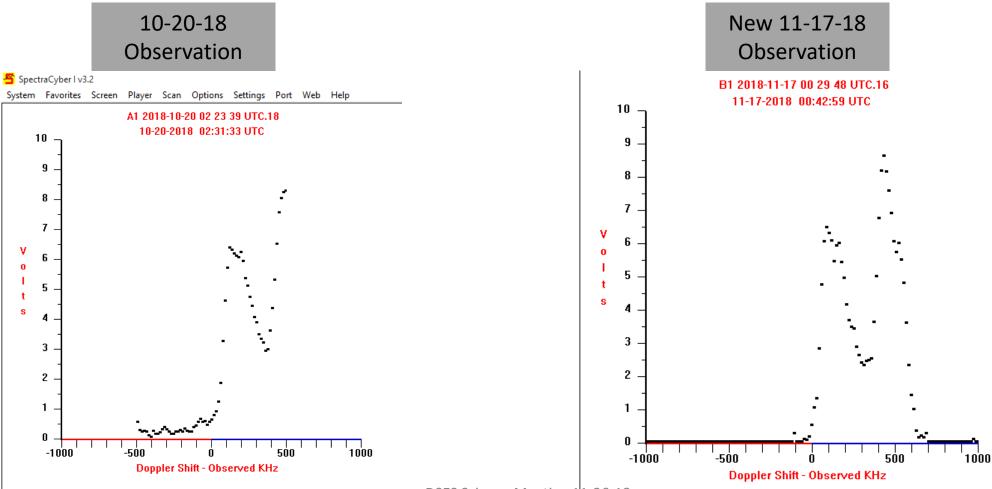


Source: W80 (HII Ionized region) RA 20h55m DEC 44d03m 550 JY

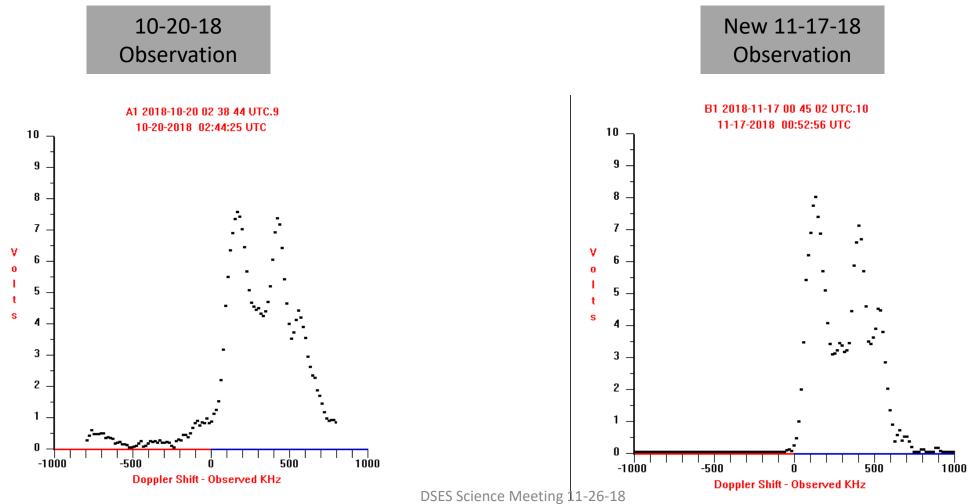
10-20-18 Observation



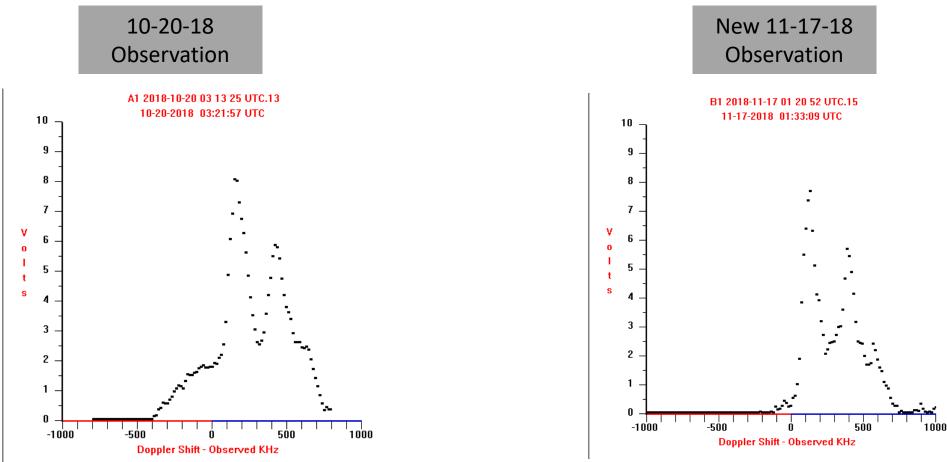
Source: DA554 (Part of HB22) RA 21h35m DEC 57d39m (106 JY @1420Mhz)



Source: DA 560 RA 21h44m DEC 57d33m 64 JY



Source: CTB 107 RA 22h24m DEC 63d32m 70 JY



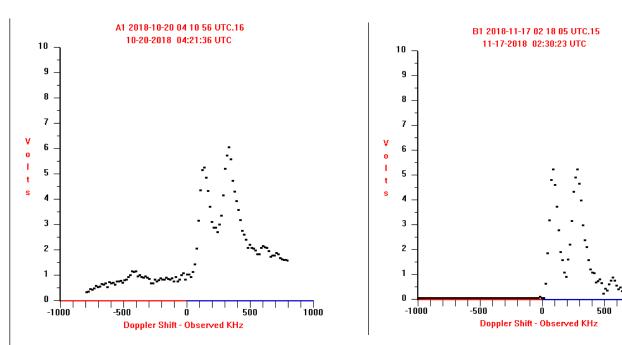
Source: Cassiopeia A RA 23h27m DEC 58d49m 2720 JY

New 11-17-18

Observation

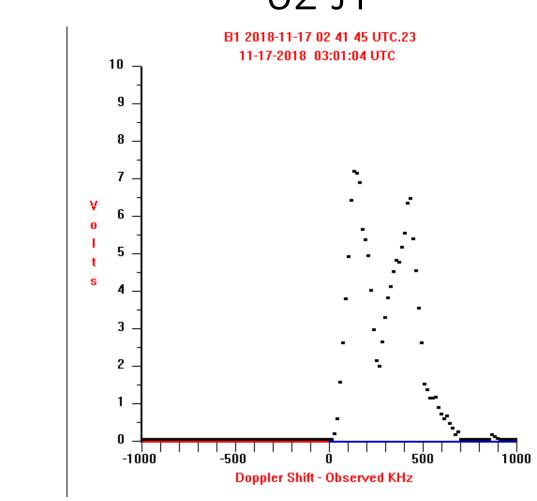


10-20-18 Observation



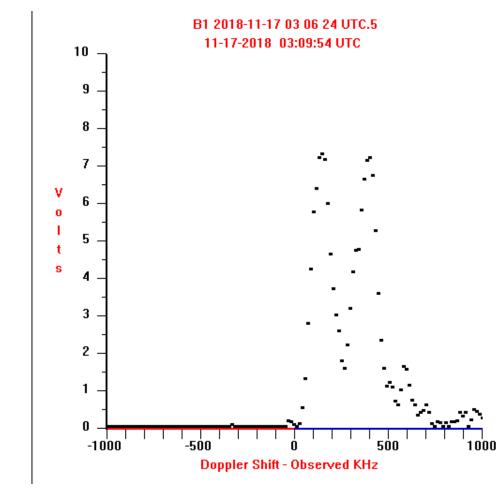
Explanation: Massive stars in our Milky Way Galaxy live spectacular lives. Collapsing from vast cosmic clouds, their nuclear furnaces ignite and create heavy elements in their cores. After a few million years, the enriched material is blasted back into interstellar space where star formation can begin anew. The expanding debris cloud known as Cassiopeia A is an example of this final phase of the stellar life cycle. Light from the explosion which created this supernova remnant would have been first seen in planet Earth's sky about 350 years ago, although it took that light about 11,000 years to reach us. This false-color Chandra X-ray Observatory image shows the still hot filaments and knots in the Cassiopeia A remnant. High-energy emission from specific elements has been color coded, silicon in red, sulfur in yellow, calcium in green and iron in purple, to help astronomers explore the recycling of our galaxy's star stuff - Still expanding, the blast wave is seen as the blue outer ring. The sharp X-ray image, spans about 30 light-years at the estimated distance of Cassiopeia A. The bright speck near the center is a neutron star, the incredibly dense, collapsed remains of the massive stellar core.

Source: DA612 RA 23hrs 55m DEC 67d56m 62 JY



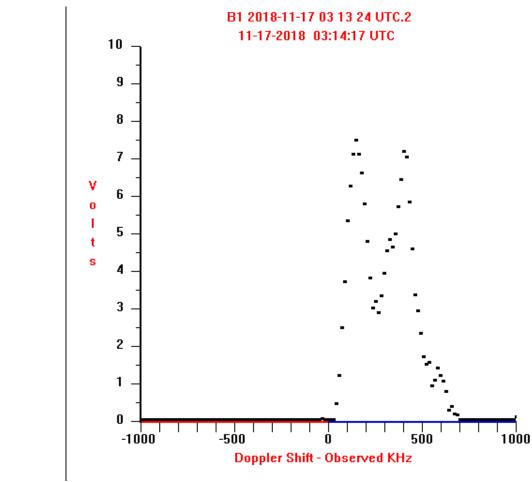
New 11-17-18 Observation

Source: DA001 RA 00hrs 02m DEC 67d19m 150 JY



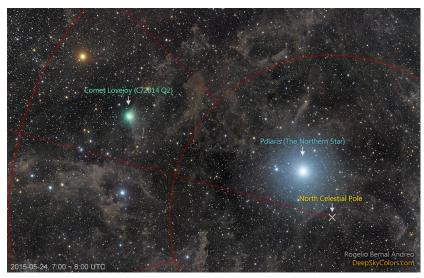
New 11-17-18 Observation

Source: DA003 RA 00hrs 05m DEC 68d06m 66 JY



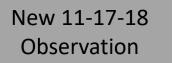
New 11-17-18 Observation

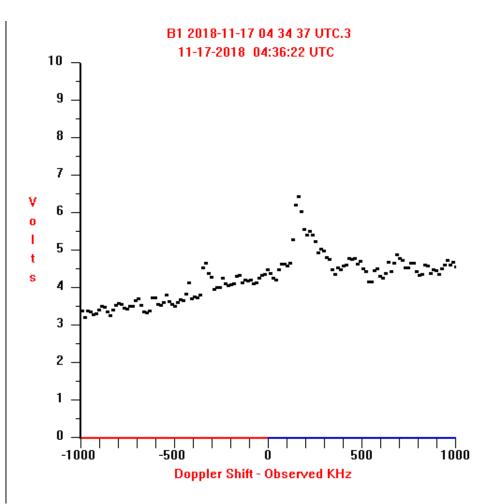
Source: Polaris RA 01hrs 28m DEC 90d00m



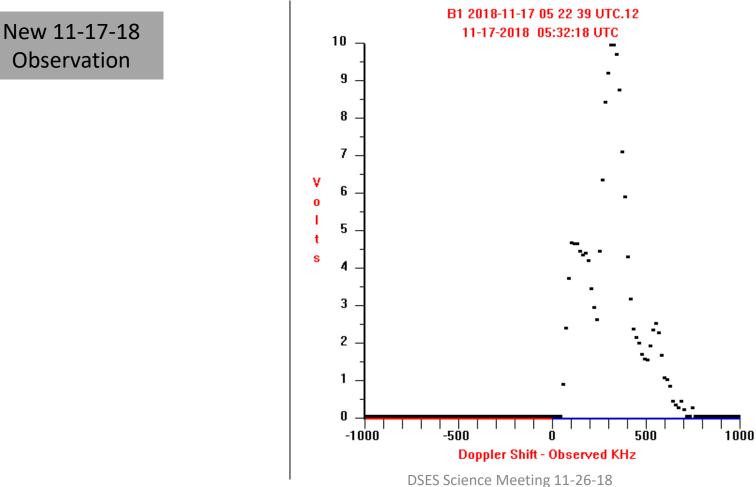
Explanation: (2015) One of these two bright sky objects is moving. On the right is the famous star Polaris. Although only the 45th <u>brightest star</u> in the sky, <u>Polaris</u> is famous for appearing stationary. Once you find it, it will <u>always appear</u> in the same direction -- all night and all day -- for the rest of your life. This is because the northern spin pole of the Earth -- called the <u>North</u> <u>Celestial Pole</u> -- points near <u>Polaris</u>.

Source: <u>https://apod.nasa.gov/apod/ap150602.html</u>



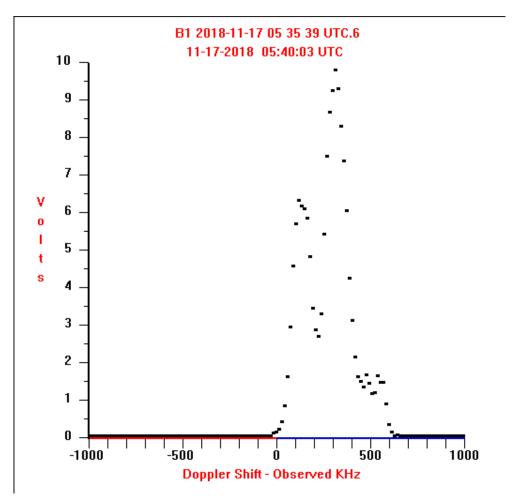


Source: DA076 RA 02hrs 25m DEC 62d08m 150 JY



Source: W04 RA 02hrs 33m DEC 61d26m 90 JY

New 11-17-18 Observation



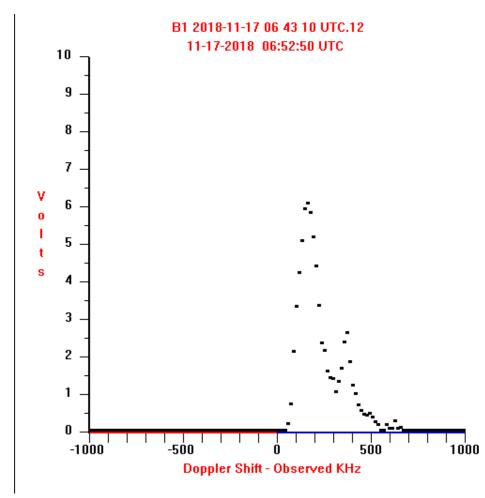


Explanation: A <u>huge chimney</u> venting hot clouds of gas out from the <u>plane of our Milky Way</u> <u>Galaxy</u> has recently been imaged in <u>radio waves</u>. The <u>Canadian Galactic Plane Survey</u> team used an array of <u>radio telescopes</u> to survey an <u>ionized gas</u> region known as W4. At the bottom of W4 and in the center of the <u>above image</u> is a very young <u>open cluster</u> of stars known as OCI 352. <u>Research continues</u> into how these stars created the <u>W4 superbubble</u>. Possible explanations include <u>supernova explosions</u> or strong <u>stellar winds</u> from these stars. It does appear clear, however, that hot gas is <u>expanding outwards</u>, being funneled by relatively cool and dense gas in a <u>chimney</u>-like fashion. The <u>W4 chimney</u>, which lies 6500 <u>light-years</u> from Earth and spans 250 light-years across, is visible as the comparatively dark area extending toward the top of the above image.

Source: <u>https://apod.nasa.gov/apod/ap000207.html</u>

Source: IC342 (Galaxy) RA 03hrs 46m DEC 68d05m 3124 JY

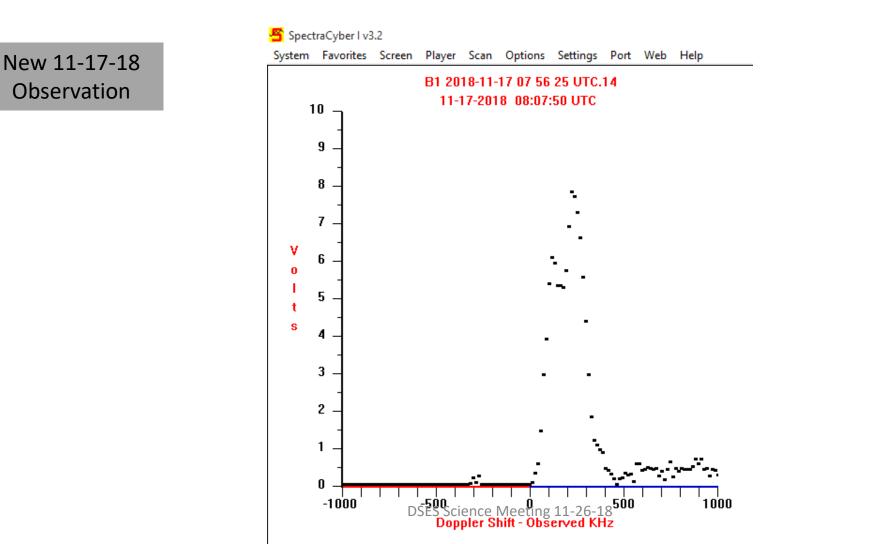
New 11-17-18 Observation



Explanation: <u>Similar</u> in size to large, bright spiral galaxies in our neighborhood, <u>IC 342</u> is a mere 10 million light-years <u>distant</u> in the long-necked, northern constellation <u>Camelopardalis</u>. A sprawling <u>island universe</u>, IC 342 would otherwise be a prominent galaxy in our night sky, but it is hidden from clear view and only glimpsed through the veil of stars, gas and dust clouds along the plane of our own <u>Milky Way galaxy</u>. Even though IC 342's light is dimmed by intervening <u>cosmic clouds</u>, this <u>sharp telescopic image</u> traces the galaxy's own obscuring dust, blue star clusters, and glowing pink star forming regions along spiral arms that wind far from <u>the galaxy's core</u>. IC 342 may have undergone a recent burst of <u>star formation</u> activity and is close enough to have gravitationally influenced the evolution of the <u>local group</u> of galaxies and the Milky Way.

Source: https://apod.nasa.gov/apod/ap170708.html

Source: DA156 RA 05hrs 01m DEC 46d28m 150 JY



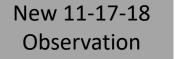
56

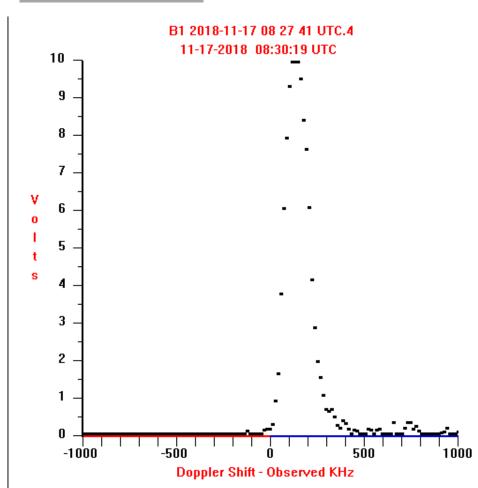
Source: W08 (IC410) RA 05hrs 22m DEC 33d22m 77 JY



Explanation: This telescopic close-up shows off the otherwise faint emission nebula IC 410. It also features two remarkable inhabitants of the cosmic pond of gas and dust below and left of center, the tadpoles of IC 410. Partly obscured by foreground dust, the nebula itself surrounds NGC 1893, a young galactic cluster of stars. Formed in the interstellar cloud a mere 4 million years ago, the intensely hot, bright cluster stars energize the glowing gas. Composed of denser cooler gas and dust, the tadpoles are around 10 light-years long and are likely sites of ongoing star formation. Sculpted by winds and radiation from the cluster stars, their heads are outlined by bright ridges of ionized gas while their tails trail away from the cluster's central region. IC 410 lies some 10,000 light-years away, toward the nebula-rich constellation Auriga.

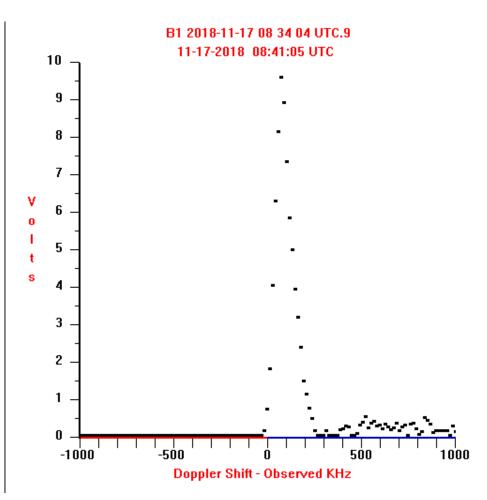
Source: https://apod.nasa.gov/apod/ap180124.html

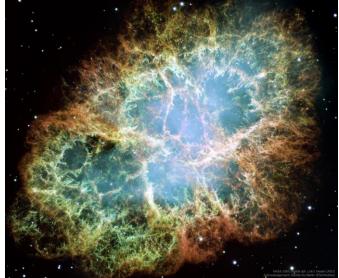




Source: W09 (M1) (Crab Nebula) RA 05hrs 34m DEC 22d00m 1120 JY

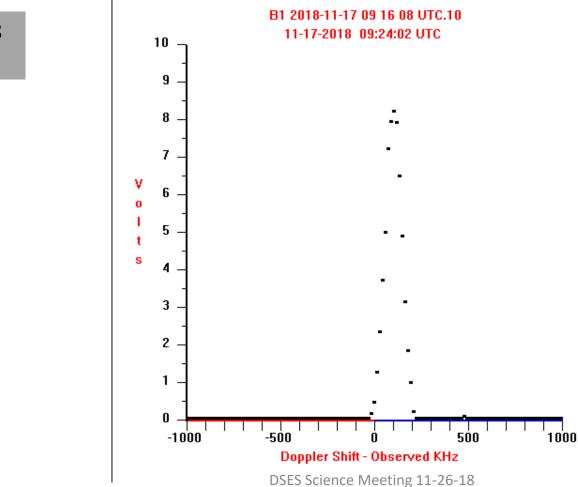
New 11-17-18 Observation





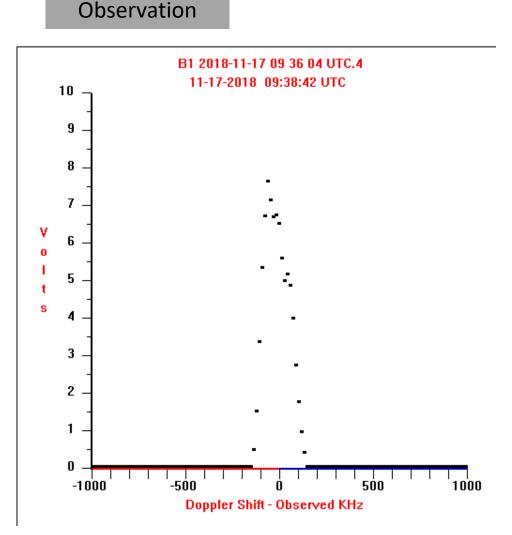
Explanation: This is the mess that is left when a star explodes. The <u>Crab Nebula</u>, the result of a <u>supernova</u> seen in <u>1054 AD</u>, is filled with <u>mysterious</u> filaments. The <u>filaments</u> are not only tremendously complex, but appear to have <u>less mass than expelled</u> in the original supernova and a <u>higher speed than expected</u> from a free explosion. The <u>featured image</u>, taken by the <u>Hubble Space Telescope</u>, is presented in three colors chosen for <u>scientific interest</u>. The <u>Crab Nebula</u> spans about 10 <u>light-years</u>. In <u>the nebula</u>'s very center lies a <u>pulsar</u>: a <u>neutron star</u> as massive as the <u>Sun</u> but with only the size of a <u>small town</u>. The <u>Crab Pulsar</u> rotates about 30 times each second.

Source: W14(CTA41) RA 06hrs 17m DEC 22d35m 215 JY



New 11-17-18 Observation

Source: W16 (Rosetta Nebula HII Ionized Region) RA 06hrs 31m DEC 04d54m New 11-17-18 170 JY



Explanation: Would the <u>Rosette Nebula</u> by any other <u>name</u> *look* as sweet? The bland <u>New</u> <u>General Catalog</u> designation of <u>NGC 2237</u> doesn't appear to diminish the appearance of <u>this</u> flowery <u>emission nebula</u>. Inside the nebula lies an <u>open cluster</u> of bright young stars designated <u>NGC 2244</u>. These stars <u>formed about four million years ago</u> from the nebular material and their <u>stellar winds</u> are clearing a hole in the nebula's center, insulated by a layer of <u>dust</u> and hot gas. <u>Ultraviolet light</u> from the hot cluster stars causes the surrounding nebula to glow. The <u>Rosette</u> <u>Nebula</u> spans about 100 <u>light-years</u> across, lies <u>about 5000 light-years away</u>, and can be seen with a small telescope towards the <u>constellation</u> of the Unicorn (<u>Monoceros</u>).

Source: <u>https://apod.nasa.gov/apod/ap170214.html</u> DSES Science Meeting 11-26-18

Summary

- Galactic Rotation Rate Observations show excellent Results
 - One more round of observations needed
- Radio Source observations increasing –
- Plan to use the new spectracyber observing PI system
- Next Observing session: Friday December 14 Start about 12AM Local