

# Search for Extra-Terrestrial Intelligence

## Geographically-spaced Synchronized Signal Detection System

West



Deep Space Exploration Society  
60 Foot Plishner Telescope  
Haswell, Colorado

Simultaneous SETI observations:  
Oct.2017 through Oct.-Nov. 2018

East



Green Bank Observatory  
Forty Foot Telescope  
West Virginia

Prepared for Deep Space Exploration Society  
by Skip Crilly  
revised Nov. 8, 2018

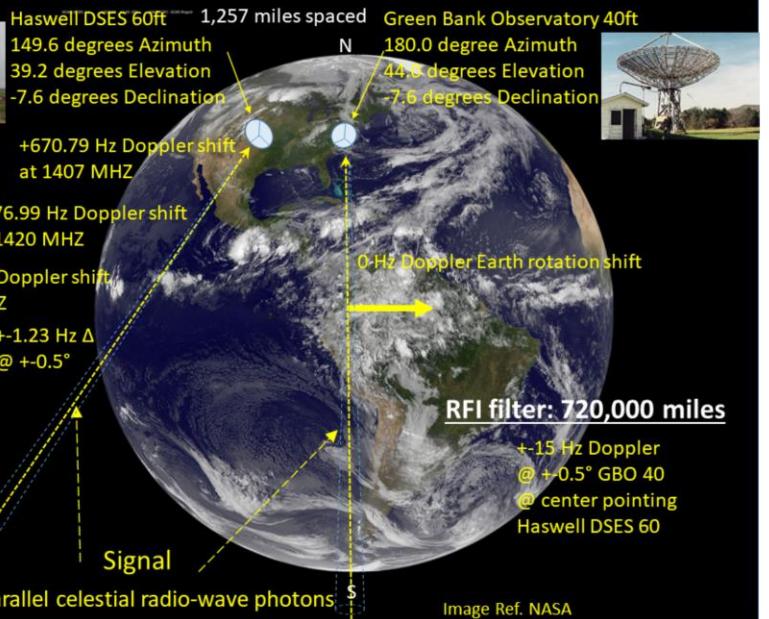
**Abstract:** A system has been designed and implemented that makes simultaneous geographically-spaced time-and-frequency-synchronized measurements of hypothetical extraterrestrial narrowband signals in the 1405-1448 MHz band. One radio telescope is the Deep Space Exploration Society sixty foot Plishner Telescope in Haswell, Colorado, and another radio telescope is the Forty Foot Telescope at the Green Bank Observatory in Green Bank, West Virginia. A GPS-signal-locked reference oscillator and a digital back-end is used at each site to permit differential Doppler measurements of radio pulses to a RF frequency resolution of 3.73 Hz. This presentation will describe signal search strategy, the receiver system, observations of simultaneous close-frequency pulses, a working hypothesis, and plans to enhance capabilities of the measurement system.

## Topics

- Description of Haswell and Green Bank Observatory telescope sites
- Search strategies
- System design
- Observations and measurement results
- Working hypothesis
- Plans

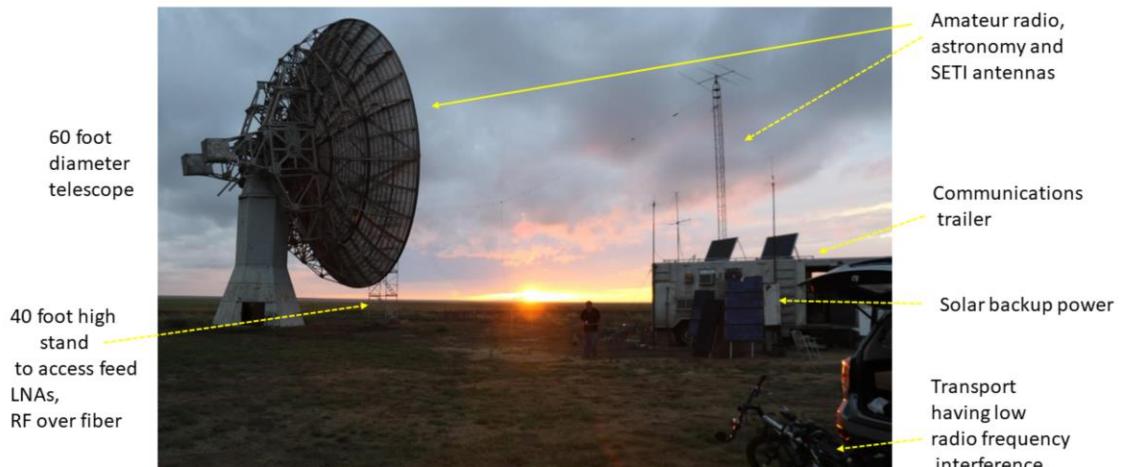
## An idea:

Use two telescopes spaced very far apart.



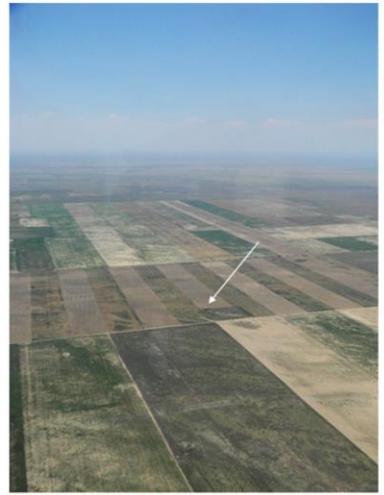
## Deep Space Exploration Society Plishner Telescope near Haswell, Colorado

Haswell simultaneous observations were made by Steve Plock, President of DSES

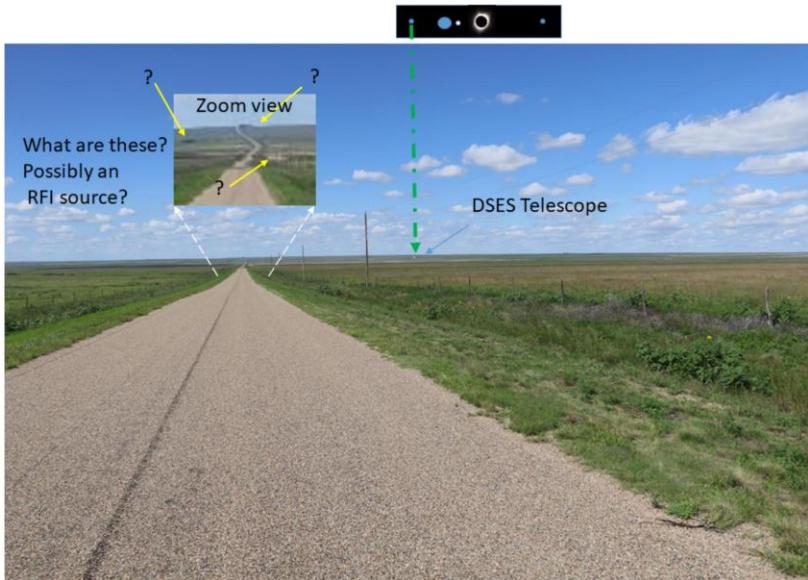


The DSES telescope is remote.

Intelligently-designed  
celestial radio transmitter here



## Looking at the 60 Foot Telescope from another direction



Lesson for SETI:  
Know your RFI sources.

## Green Bank Observatory

### Forty Foot Telescope

West Virginia

***This is where I work  
part time***

Primary Activity:

Education and Public Outreach

Thousands of students,  
educators and amateur astronomers observe per year

SETI computer surrounded by RFI absorbing foam



**Forty Foot Telescope**

Green Bank Telescope  
In background  
100m x 110m;  
~100 dBi gain @120 GHz  
***world's largest fully  
steerable telescope***

Third SETI telescope site, 28 foot diameter paraboloid under construction on New Hampshire farm, 5 pixels of dual polarization, surrounded on all sides by deep forests



Proposed prime prerequisite  
of transmitter (ETI) and receiver (human) entities

***Understand communication theory, i.e. math : e.g. Shannon's Law***

Examples:

- Energy per communicated bit limitations
- Bandwidth available for communication
- Data rate possible
- Detectability functions

## Two telescope site signal search strategy

Filtering methodology

What are the signal requirements to be significant?

SNR = Signal to Noise Ratio

1. SNR > 12 dB (15.8) at distant locations, e.g. Haswell and Green Bank, and
2. Measured close frequency (post Doppler + correction) at the same time at two distant telescope sites

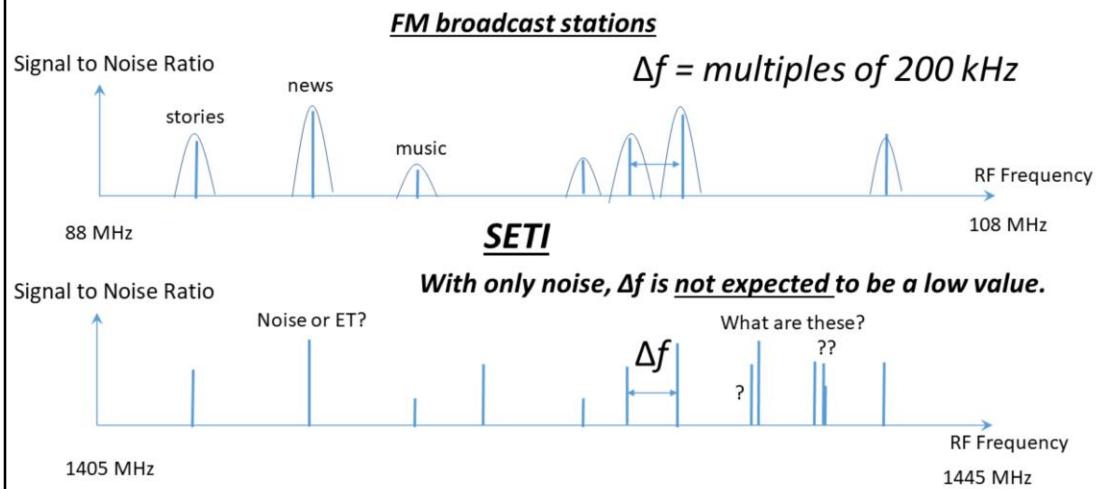
or many close frequency spaced tones at both sites

**Open squelch**

followed by

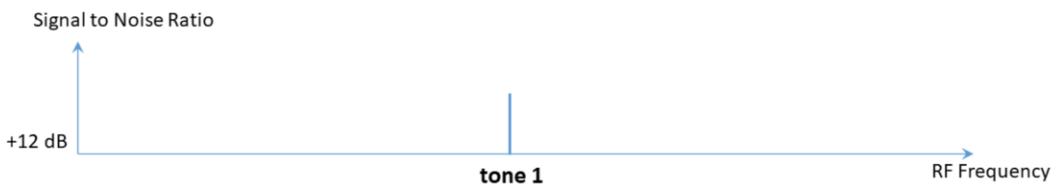
Post-squelch information-containing signal analysis

Signals carrying information use frequencies  
example: FM broadcast band



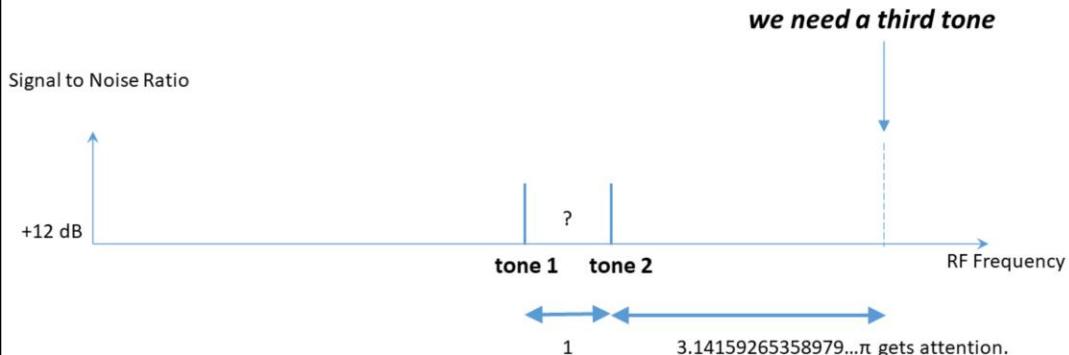
## Frequency Triplets: Why search for triplets?

***Difficult to encode / decode information in one tone***



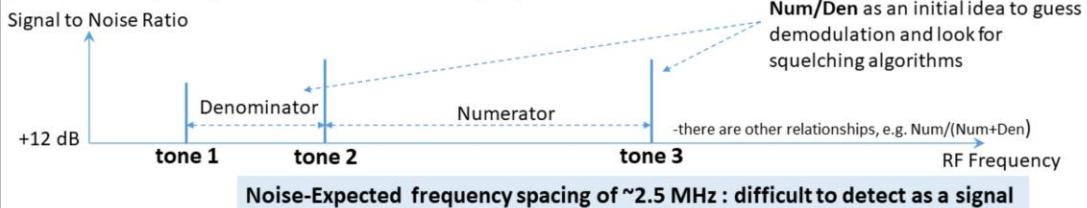
Idea: It seems preferable to search for information content in one dimension, e.g. frequency, at a time.

***Difficult to encode / decode information in two tones***

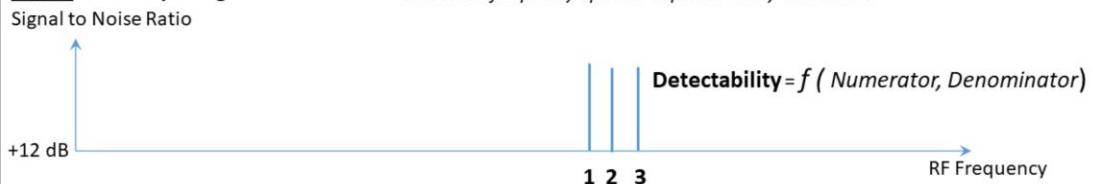


## Frequency Triplets: Why search for triplets?

**Noise:** ~wide spacing between tones is normally expected

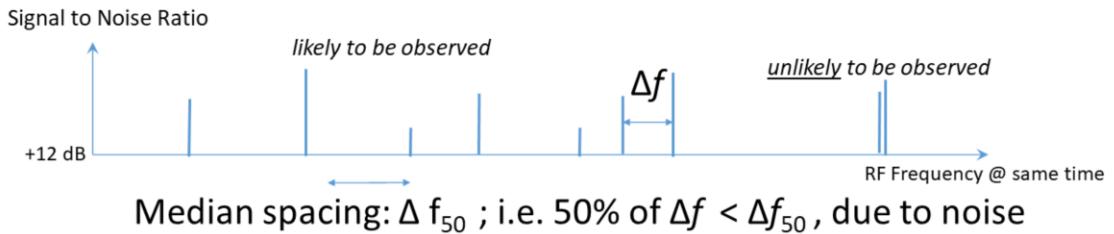


**Signal:** Close spacing between tones This close-frequency-spaced triplet is easily detectable



## Likelihood of close-spaced tones due to noise

$$\text{Detectability} = f(\text{Numerator}, \text{Denominator})$$



$$\text{Likelihood}(\Delta f) = 1 - e^{-\ln 2 (\Delta f / \Delta f_{50})}$$

$$\text{Likelihood}(\Delta f) \approx \ln 2 (\Delta f / \Delta f_{50}) \quad \text{if } (\Delta f / \Delta f_{50}) \ll 1$$

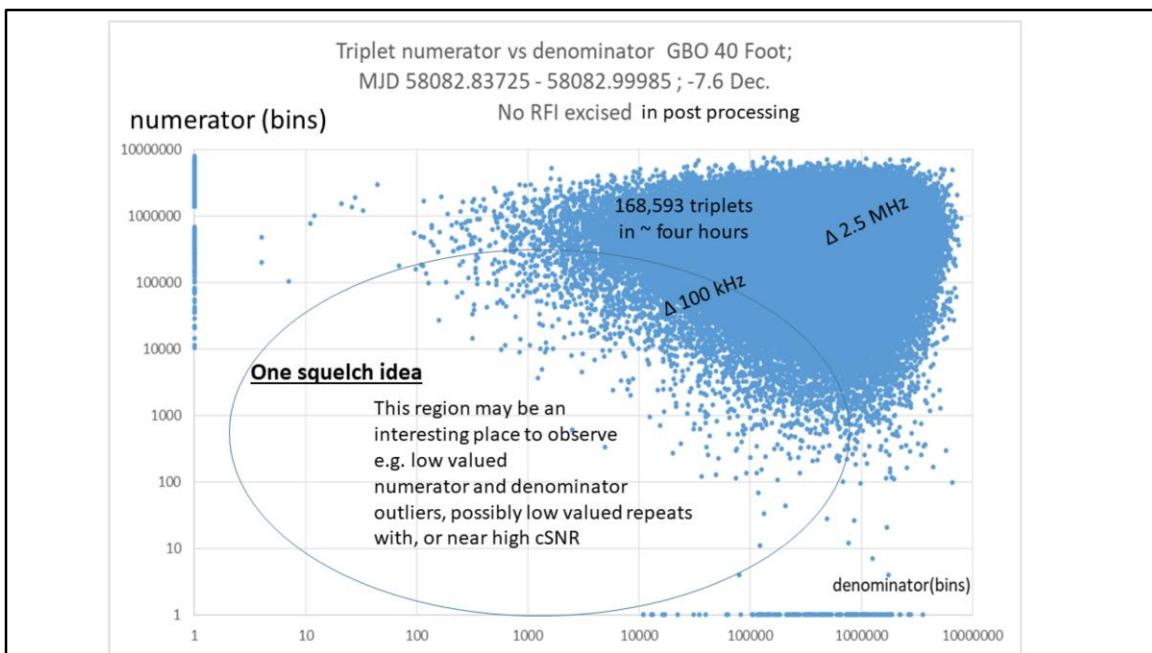
derived from Poisson Point Process model; notebook #16 page 68; tested in simulation

The likelihood of tone spacing was derived from the concept that tones, expected to be due to noise, occur at frequencies uniformly distributed within a given interval of frequency. A first order differential equation results from the analysis of multiple trials having various tone spacings, resulting in an exponential solution to a probability density function. The cumulative probability is then obtained from an integral of the probability density function.

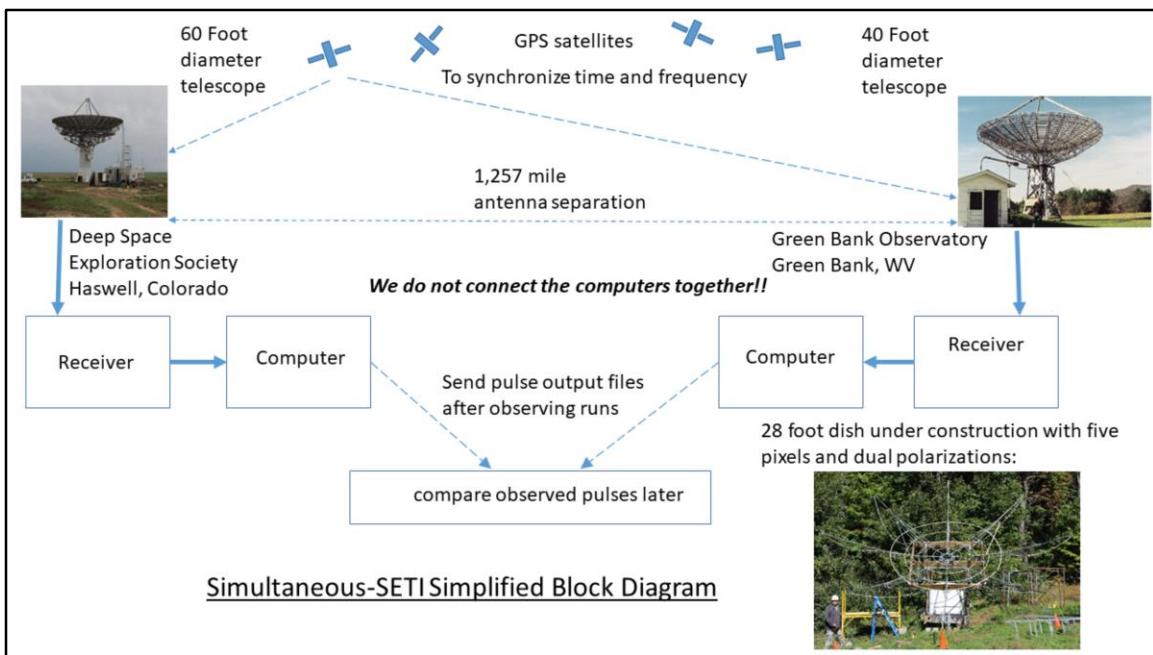
The calculation above is referred to as a likelihood, as the calculation is most commonly used in the observations, in this work, to calculate the probability of an event a posteriori.

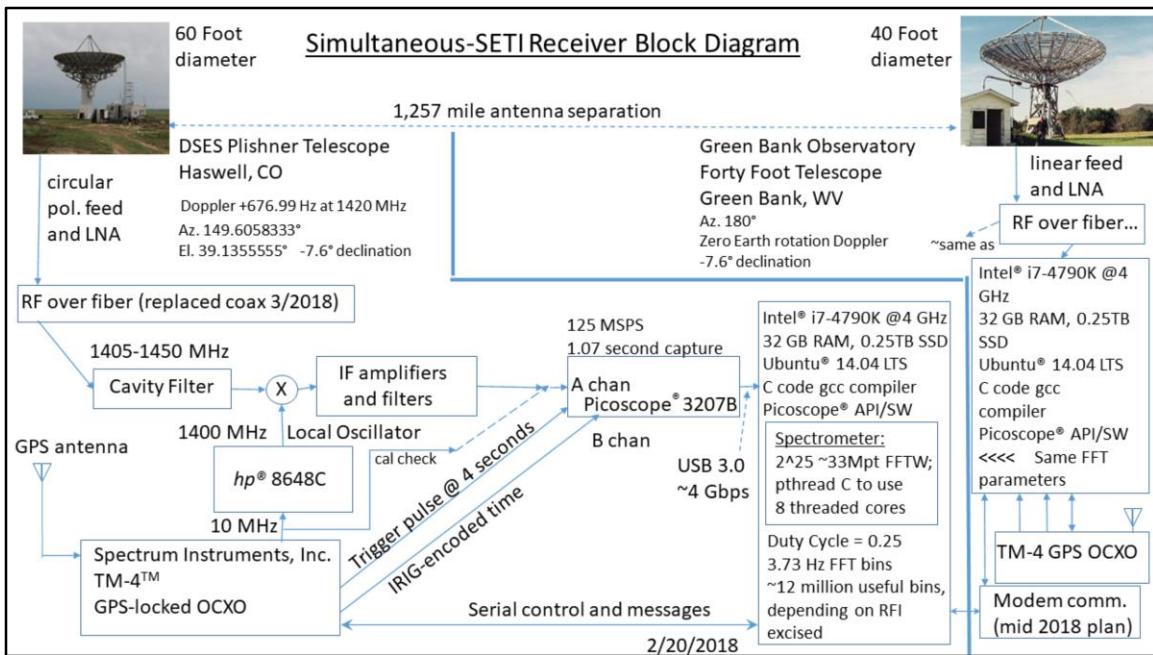
The calculation above is an *estimated* likelihood due to noise, as RFI can affect the determination of an estimate of  $\Delta f_{50}$ .  $\Delta f_{50}$  is estimated from an observation of tone spacings, expected to be due to Rayleigh IQ amplitude distributed noise, that exceed the 12.0 dB composite one second signal to noise ratio measurement. Composite SNR is implemented to reduce noise-caused false positives that exceed a lower 11.2 dB threshold in a single 0.26 s interval of the time samples that are applied to the uniform windowed FFT. The 11.2 dB threshold and the 12.0 dB threshold must both be exceeded to record a pulse to the output file. The combination of the two SNR

thresholds has not been used in a stochastic process to rigorously derive  $\Delta f_{50}$ .



## System Design

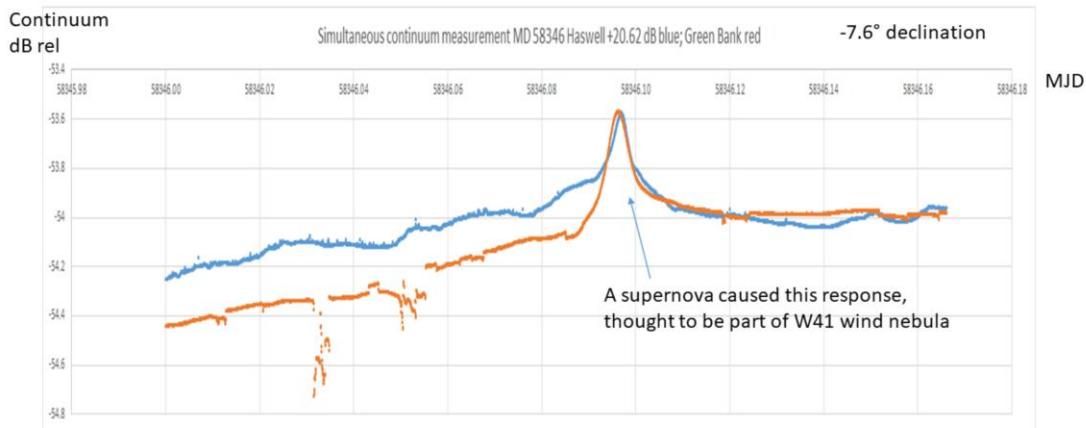




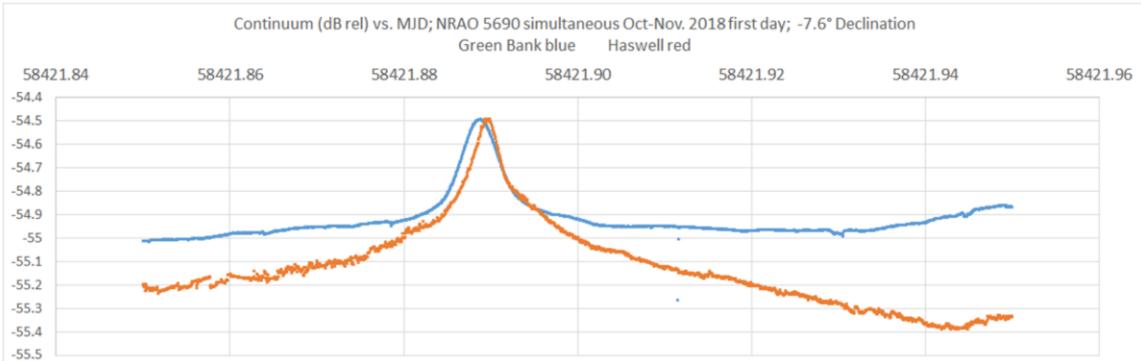
## System Pointing Validation

Are the Haswell and Green Bank telescopes  
pointing at the same Right Ascension & Declination?

## August 16, 2018 NRAO 5690 simultaneous Green Bank & Haswell observation



## Oct. - Nov. 2018 NRAO 5690 transit Haswell and Green Bank pointing verification



## Geo-multiplexing....the same signal may not be received at all locations on Earth

Is there a message in a message?

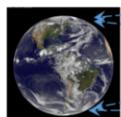
To open squelch :

*One, or a few, tones of many simultaneously transmitted tones need(s) to be at close frequency at two telescope sites. After squelch opens, associated pulsed tones observed at each receive site contain a subset of a message.*

### Why send geo-multiplexed information?

- For higher channel capacity?
- To prevent nuclear-deterring adversaries from secretly receiving "useful" information?

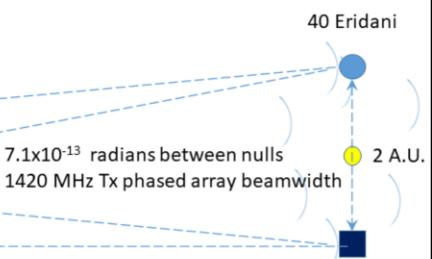
### Message decoding requires international cooperation



~100 beamwidths in 1D, ≈7k in 2D

8.3x10<sup>-11</sup> radians angle subtended

16.3 light years = distance to 40 Eridani

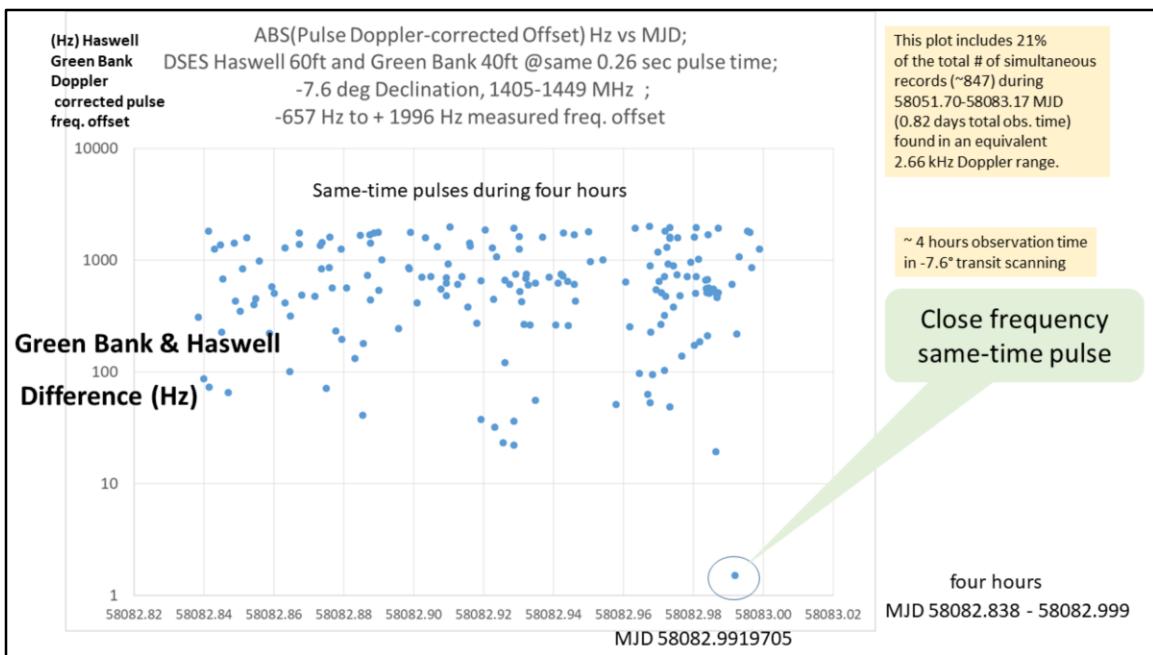


## Observations, measurements and calculations

Observations that are described here occurred during Haswell and Green Bank simultaneous transit scan pointing to -7.6° declination on November 25, 2017, (MJD 58082) April 10-11, 2018, (MJD 58219) and Green Bank standalone June 7-9, 2018, (MJD 58276-8), and August 15-16, 2018 (MJD 58345-6).

### Notes:

- 1. Some pulse observations, described in this presentation, calculate to be highly unlikely to be noise-caused.**
- 2. The following is not a *claim* of detection of an extra-terrestrial intelligent signal.**



Simultaneous pulse observed at time 58082.9919705  
≈1.5 Hz difference after Doppler at center of beam and sample clock offset correction

### Triplet tone frequencies and cSNR un-postprocessed

#### Haswell

Tone 1 1429.535342 MHz 12.186 dB

Δ ≈73 kHz measured (expected ≈2.5 MHz)

Tone 2 1429.461417 MHz 12.474 dB

Tone 3 1414.049824 MHz 12.819 dB

#### Green Bank

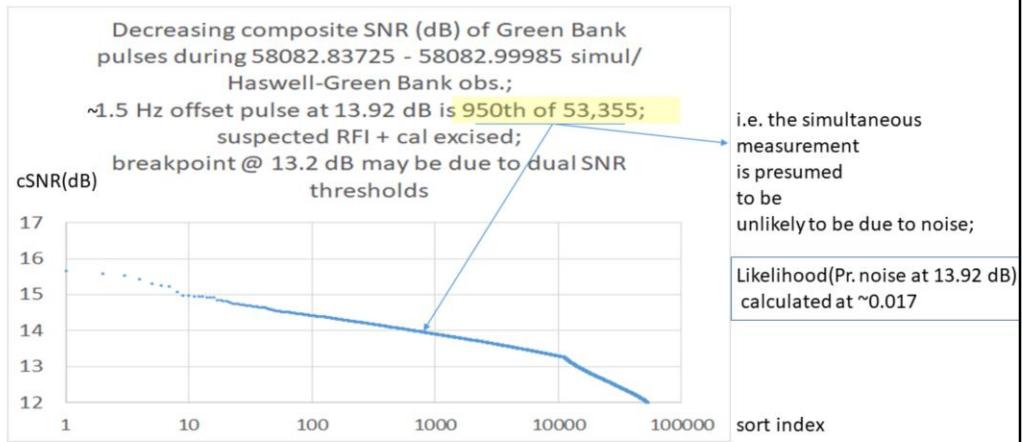
1429.534683 MHz 13.92 dB

Δ ≈80 kHz measured (expected ≈2.5 MHz)

1429.454120 MHz 12.675 dB

1428.671384 MHz 13.261 dB

## 13.92 dB composite 1 s SNR of Green Bank & Haswell simultaneous ~1 Hz offset pulse



## Estimated Likelihood of simultaneous ~1 Hz offset Haswell & Green Bank pulsed tones, due to noise

Likelihood function uses **an idea** that “un-squelching” the receiver requires:

1. 12.0 dB threshold-crossing composite 1 s signal-to-noise ratio at Haswell & Green Bank
2. 13.92 dB threshold at Haswell **or** Green Bank
3. adjacent tone bin spacing  $\leq 21626$  at Haswell and Green Bank (i.e.  $< 80$  kHz)

- ~1 Hz offset during the same 0.26 s interval (expected ~once per day)  
**AND (**

1. 13.92 dB composite 1s SNR Green Bank likelihood  $\sim 0.017 * 2$  (GB or Haswell)
2. 21626 bins ( $\Delta 80$  kHz) Green Bank numerator tone spacing likelihood  $\sim 0.03$
3. 19844 bins ( $\Delta 73$  kHz) Haswell numerator tone spacing likelihood  $\sim 0.03$

**Combining three AND factors in Likelihood function:**

- $0.017 * 2 * 0.03 * 0.03 \approx 0.000031 = 31$  ppm

i.e. event is expected 31 times per million simultaneous Haswell-Green Bank pulses,

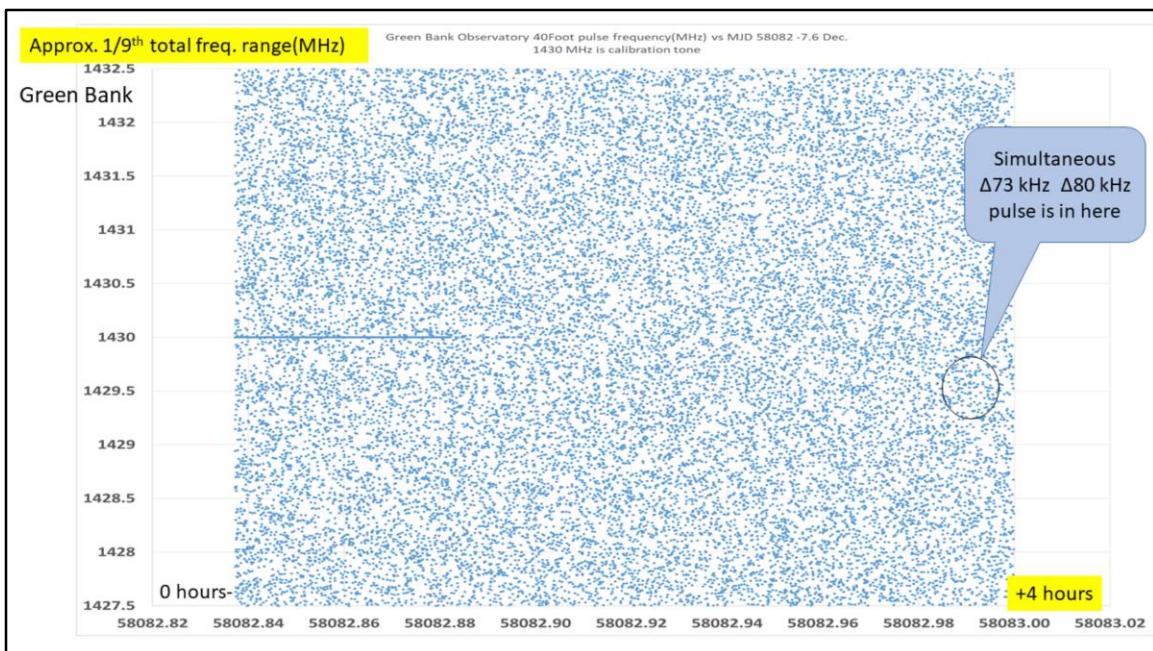
*if squelch conditions are independent, and noise is the anomaly's suspected cause.*

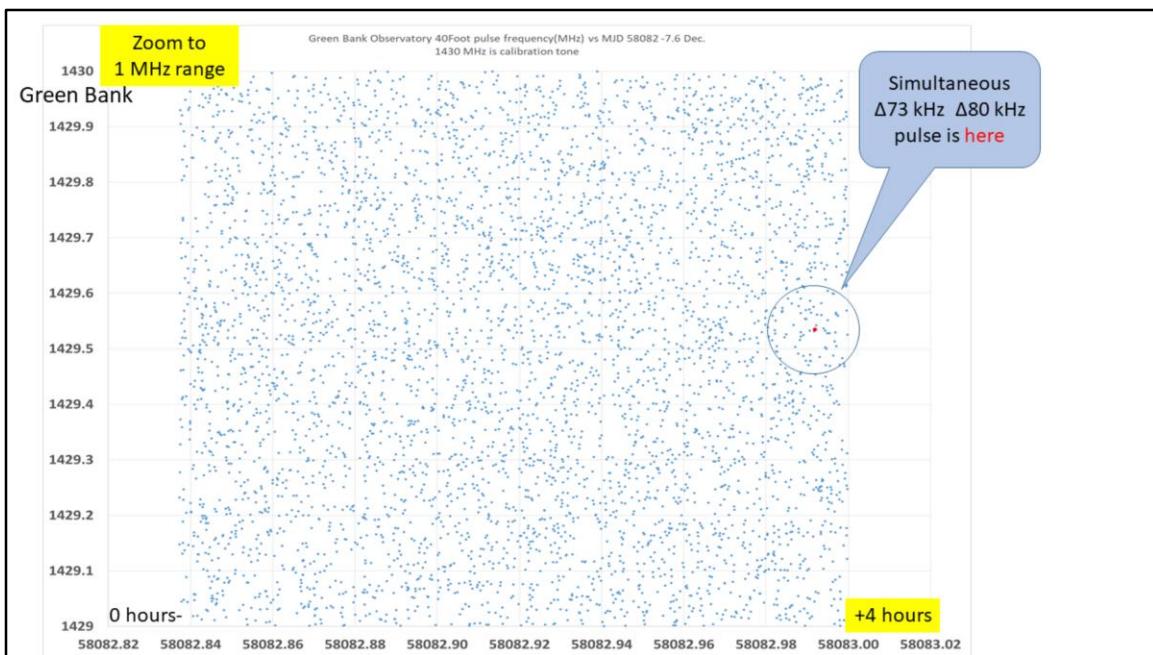
Squelch has opened

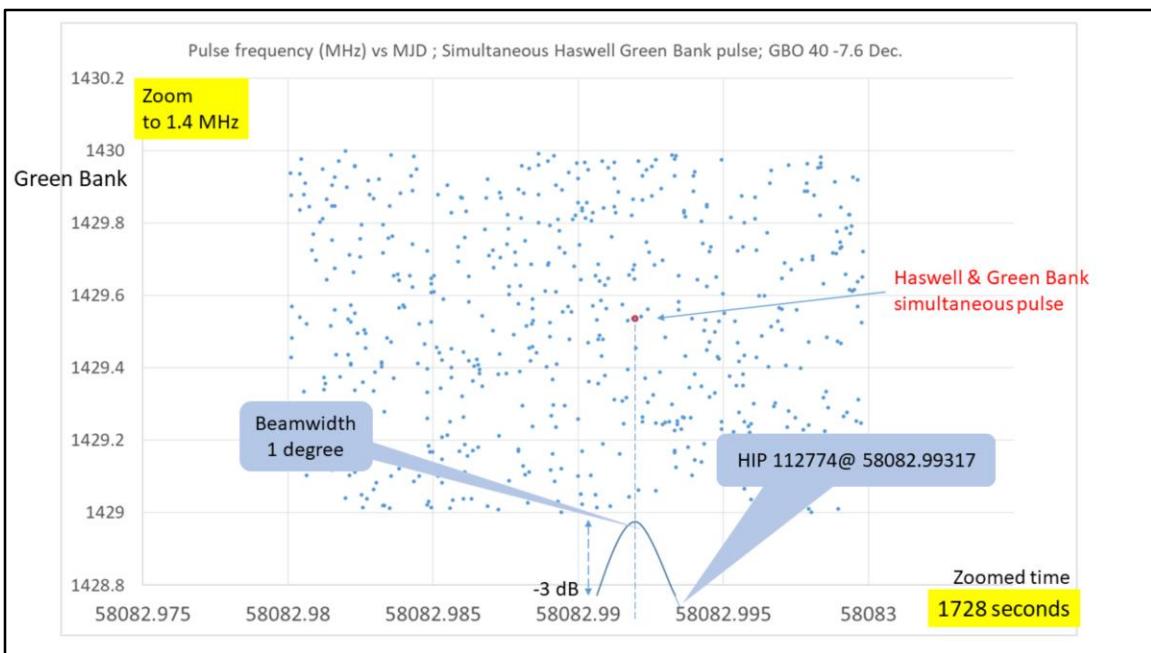


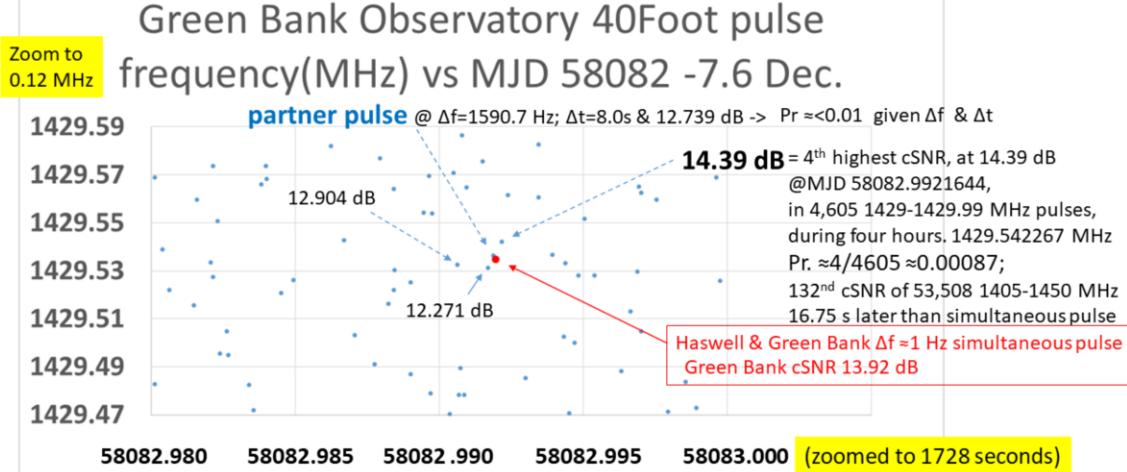
## Post-squelch information-containing signal analysis

Did other anomalous pulses appear  
near the time & frequency of the simultaneous pulse?









An unusual partner pulse was observed. A high composite SNR pulse was observed.

1.1	58082.992	<b>58082.991878</b>	-53.32536	30971	7928583	<b>1429.536273 MHz</b>	
-54.18	11.925	-3.058	-2.118	3.231	<b>12.739 dB</b>	17.127	292847 1815684 2108531 2.64559
	2.75343	2.82148					
6.200111320928676066341809887074e+00			11				
<b>4.928038974273873767492193268345e+02</b>			2				
1.051994121804501204928973163956e+214			5				
S							
1.1	58082.992	<b>58082.991878</b>	-53.32536	31112	7964918	<b>1429.671632 MHz</b>	
-54.27	11.38	-7.632	-0.897	3.564	<b>12.304 dB</b>	17.173	1815684 36335 1852019
<i>ratio Shannon Entropy = 2.37878</i>		2.64559	2.89629				
<b>2.001174213134003494000057278689e-02</b>	20	<b>0123479 late 5 6 8</b>					
<b>1.020213319435215440085210805420e+00</b>	10	<b>0123459 digits late 6 7 8</b>					
2.773786403406941519449065953088e+00		14					
S							
1.1	58082.992	<b>58082.991878</b>	-53.32536	31317	8017156	<b>1429.866234 MHz</b>	
-54.23	11.538	-1.471	-0.267	-1.115	<b>12.221 dB</b>	17.199	<b>36335 52238 <math>\Sigma = 88573</math></b>
<i>ratio Shannon Entropy = 2.35988</i>		2.69951	2.69951				
<b>1.4376771707719829365</b>	62543002615e+00	<b>01346789 late 2 5</b>		14			
<b>4.210903239036432115141323885143e+00</b>		5 01234569 late 7 8					
6.741740635061165008466862357335e+01		37 late 239					
S							
<b>Bin <math>\Sigma</math> event rank 1652 / 216711 = 0.0076</b>							

*Green Bank partner pulse measured data  
associated triplet of the partner of the  
simultaneous Haswell-Green Bank pulse*

*Close tone spacing  
Expect: ~2.5 MHz  
Measure: 135, 195 kHz*

*tone spacings in FFT bins*

It was surprising to observe that the 1590.7 Hz offset pulse partnered to the simultaneous pulse, is itself a triplet with a low predicted likelihood due to noise. The rank of the combined tone spacing was calculated at 1,652 of 216,711 combined triplet tone spacings.

This tones comprising the triplet, and the high cSNR pulse described in the previous slide are referred to as **associated pulses**, as the pulses appear to be associated with a simultaneous pulse.

If a transmitting entity is attempting to reduce the noise-liability of transmitted pulses measured at the receiver, then it seems natural that the transmitting entity will transmit low noise-liability pulses, at close to the same time and frequency as a simultaneous pulse.

### Likelihood analysis of three triplets containing seven tones close in time and frequency

#### close partner to GB Haswell pulse

Tphase.site	MJD time	Freq. before corrections	composite1sSNR	denominator(bins)	numerator(bins)
1.1 GB	58082.99187790	<b>1429.536273 MHz</b>	<b>12.739 dB</b>	292847	1815684
1.1 GB	58082.99187790	<b>1429.671632 MHz</b>	12.304 dB	1815684	36335
1.1 GB	58082.99187790	<b>1429.866234 MHz</b>	12.221 dB	<b>36335</b>	<b>52238 C.</b> 0.3299601 MHz total
		<u>8.0 second time difference; seven 0.26 s intervals separating .9918779 and .9919705</u>			
1.2 HA	58082.99197050	1414.049824 MHz	<b>12.819 dB</b>	360460	14389
1.2 HA	58082.99197050	<b>1429.461417 MHz</b>	12.474 dB	14389	4137018
1.2 HA	58082.99197050	<b>1429.535342 MHz</b>	12.186 dB	4137018	<b>19844 A.</b> $\Delta$ 73.3 kHz

#### Simultaneous Haswell&Green Bank pulse ~1 Hz difference with Haswell Doppler corrected

1.1 GB	58082.99197050	<b>1428.671384 MHz</b>	<b>13.261 dB</b>	251301	760778
1.1 GB	58082.99197050	<b>1429.454120 MHz</b>	12.675 dB	760778	210114
1.1 GB	58082.99197050	<b>1429.534683 MHz</b>	<b>13.920 dB</b>	210114	<b>21626 A.</b> $\Delta$ 80.5 kHz

#### Likelihood Function: (receiver squelch function)

- A. Meas. noise average frequency difference between adjacent tones = 2.50 MHz;  $80.5 \text{ kHz} / 2.5 \text{ MHz} = 0.032$ ;  $73.3 \text{ kHz} / 2.5 \text{ MHz} = 0.029$
- B.  $<1.6 \text{ kHz} / 2.50 \text{ MHz} = 0.00064$  @ 8 tries = 0.005
- C. Likelihood of  $<(36335+52338) = 0.0076$

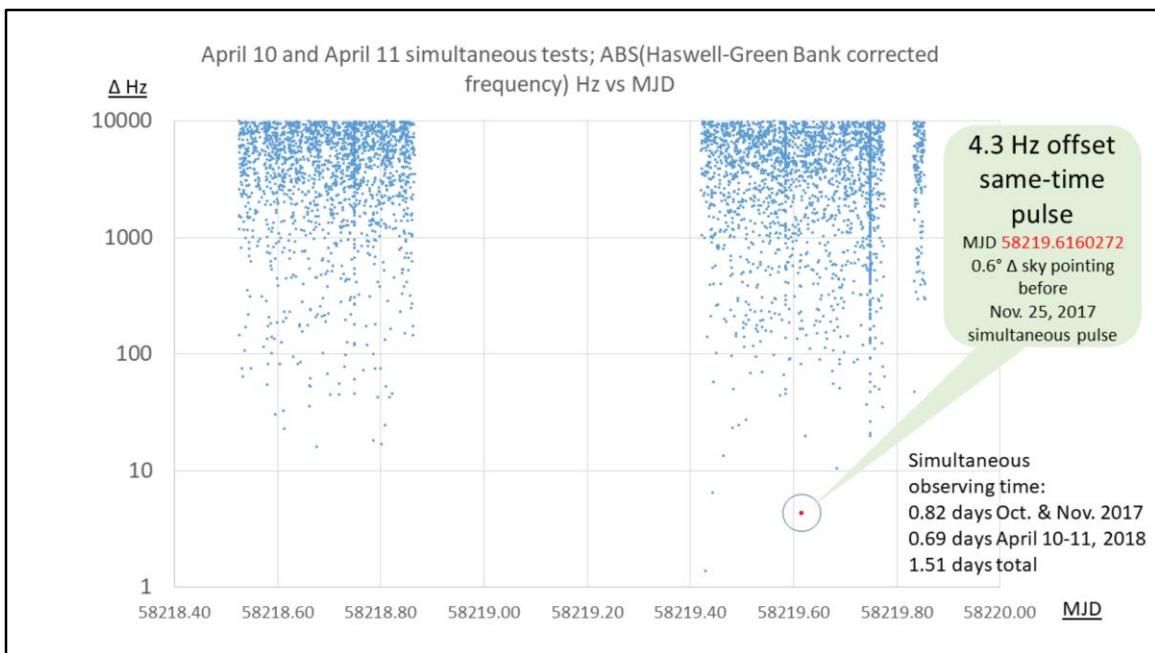
**Estimated Pr.(A. AND B. AND C.) =  $3.6 \times 10^{-8}$**

Three pairs of Haswell close frequency spaced pulses (MHz) at time of [simultaneous Green Bank & Haswell 11/25/2017 pulse event](#) @ 58082.9919705; expect ~ 2.5 MHz spacing

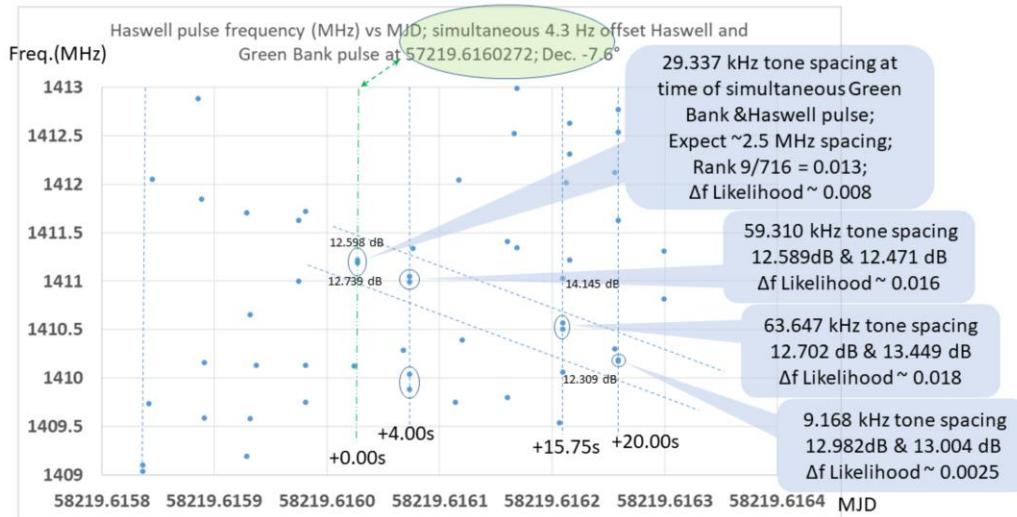
**Pr. (A&B&C) =0.0003**

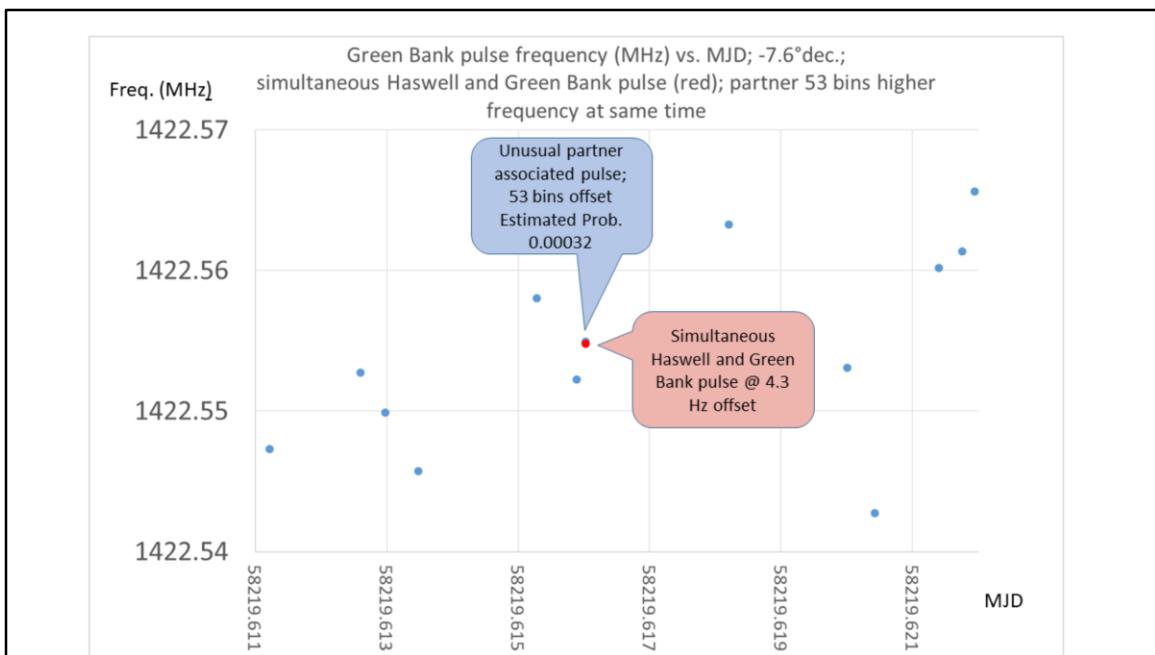
1.2 58082.9919705	<b>1429.461417 MHz</b>	11.238	2.82	1.657	0.004	12.474	19.16	14389	4137018
3.03915	3.16825 0	2.875125443046771839599694210856e+02				19			
7.330126498690833109980296566619e+124	9	2.098578716467387692404358116884e+323228496				3	S	<b>Pr.A= 0.029</b>	
1.2 58082.9919705	<b>1429.535342 MHz</b>	11.224	-8.965	0.571	3.05	12.186	17.272	4137018	<b>19844</b>
2.89629	2.69951 2.52164	4.796691723362093179193322339908e-03				6			
1.004808214265100155122801535450e+00	31	2.73138338222090805838810539540e+00				14	S		
1.2 58082.9919705	<b>1430.617684 MHz</b>	11.582	-2.303	-1.345	-1.11	12.173	17.051	19844	290539
0		1.464115097762547873412618423705e+019				2.283342594678110828968988737034e+069			2.89629
							1.22750101080819012905236783302e+991643		3.03915
									5
1.2 58082.9919705	<b>1434.062058 MHz</b>	11.237	4.155	-9.143	1.912	12.449	17.042	290539	924592
2.4956	2.64559 2.69951	3.182333524931248472666320184209e+00				11			
2.410293277330761589368091428148e+01	6	2.936099040627757048940012881633e+10				2	S	<b>Pr.B= 0.084 x2</b>	
1.2 58082.9919705	<b>1434.27273 MHz</b>	11.556	6.051	-1.538	-2.675	12.917	17.295	924592	<b>56552</b>
2.6645	2.64559	6.116427570214754183466869711181e-02				27			2.69951
11	2.895256004502940247926126942976e+00	3	S						
1.2 58082.9919705	<b>1439.466836 MHz</b>	12.093	-19.73	1.911	-0.514	12.706	17.466	56552	139432
0		2.465486631772527936883102984863e+017				5.09885185433663173236524917035e+03			2.89629
							2.098578716467387692404358116884e+323228496		3
1.2 58082.9919705	<b>RF 1440.219452 MHz</b>	18.374	17.343	16.86	18.83	23.943	26.471	143126	58903
		4.115464695443175942875508299392e-0110				1.509149835684187435566507129451e+004			2.54952
							4.5228896477593833381031639699e+008		S
1.2 58082.9919705	<b>1441.339032 MHz</b>	12.238	-2.56	1.098	-4.519	12.773	26.473	58903	300535
2.30596	2.75343 2.73967	5.102201925199056075242347588408e+00				8			
1.643834691908166578075053153057e+02	9	2.459425013413591978985376347436e+71				4	S	<b>Pr.C= 0.030 x2</b>	
1.2 58082.9919705	<b>1441.414321 MHz</b>	11.542	-3.914	-4.162	1.319	12.15	24.522	300535	<b>20210</b>
2.46772	2.69951 2.89629	6.724674330776781406491756367811e-02				23			
1.069559352231106167055650327124e+00	4	2.914095127503567680263818833628e+00				2	S		

April 11, 2018 simultaneous Green Bank & Haswell  
close frequency pulse

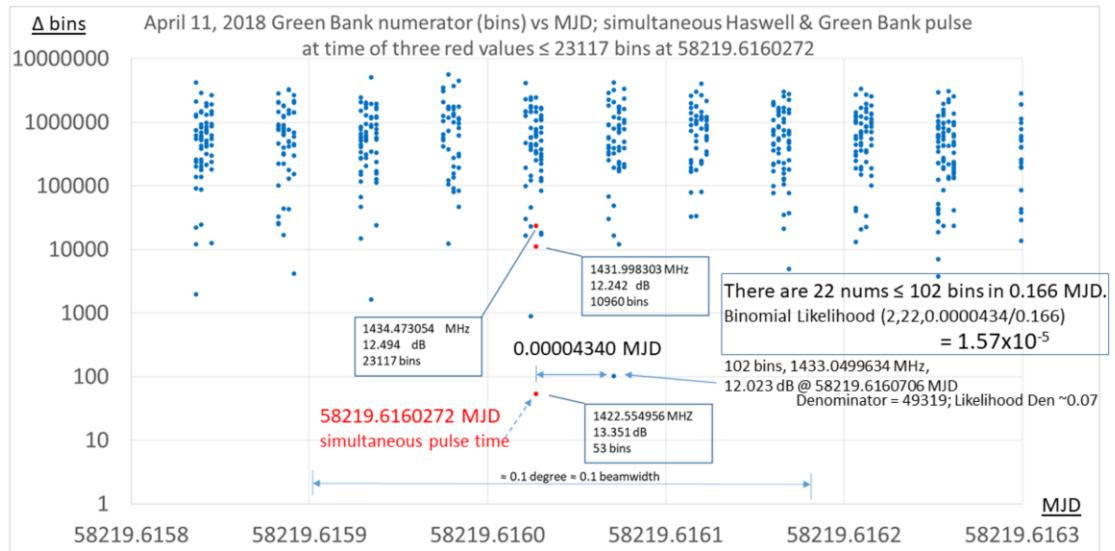


## Haswell close-frequency spaced pulses at & near time of simultaneous April 11, 2018 Haswell & Green Bank pulse





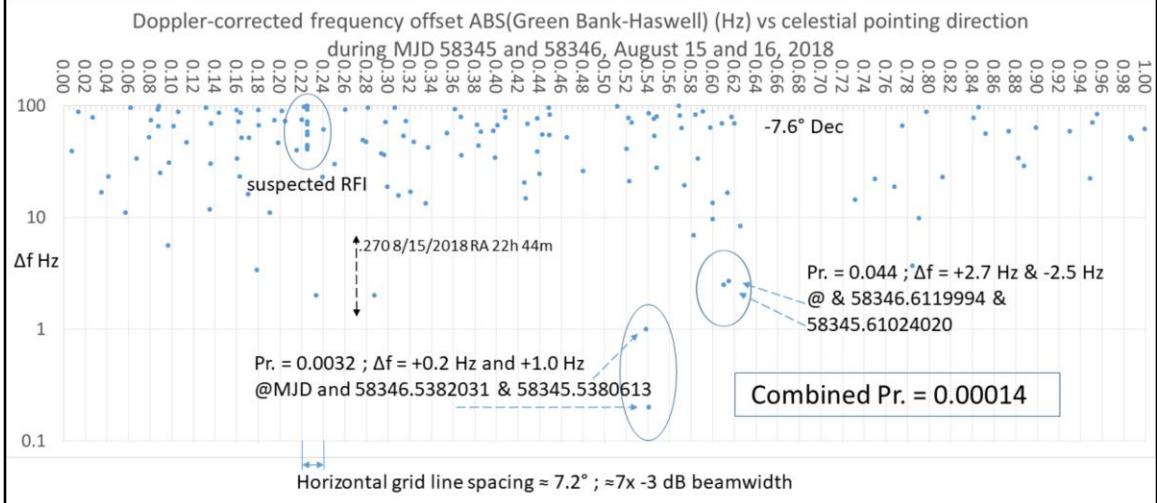
Other Green Bank closely spaced pulses at or near time of simultaneous Haswell and Green Bank pulse?

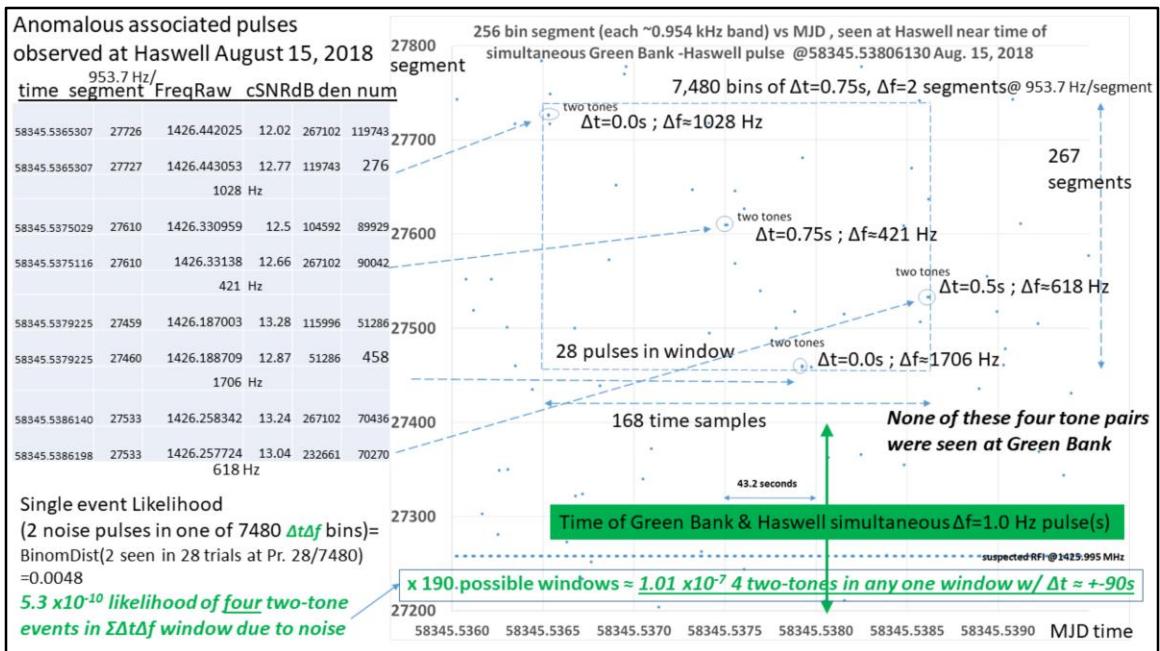


August 15 and 16, 2018  
simultaneous Haswell and Green Bank pulses  
and associated pulses

## August 15 and 16, 2018 simultaneous Haswell and Green Bank observations

Search for simultaneous anomalous pulses with RFI excision, close to the time of these simultaneous pulses.





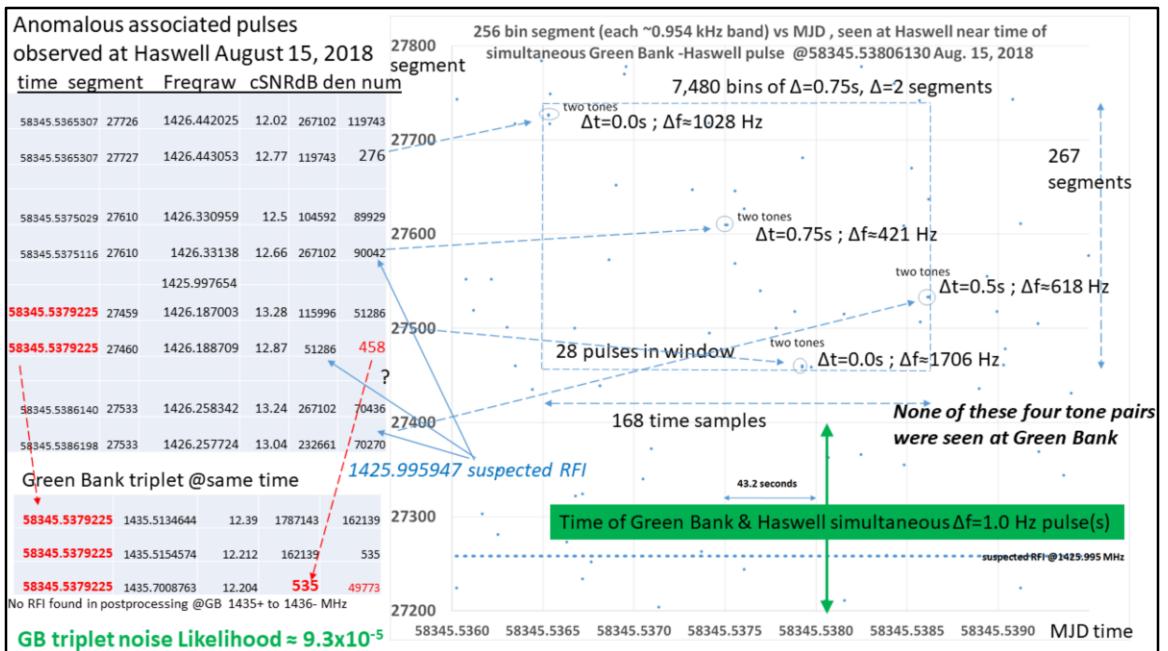
One glaring very difficult problem is raised in the data and diagram in **Anomalous associated pulses observed at Haswell August 15, 2018**.

It seems difficult to explain why the time-frequency window's event, comprising four close-frequency tone pairs, a calculated  $5.3 \times 10^{-10}$  noise-likelihood event, should be captured at Haswell within -140 s and +50 s of a simultaneous same FFT bin, 1.0 Hz offset frequency pulse, Doppler-corrected, observed in the same 0.26 s interval at Haswell and Green Bank. The Doppler-corrected one FFT bin overlap of a simultaneous Haswell and Green Bank pulse is expected to result from a source distance four times the Earth-Moon distance. If the signal's source is human RFI, it seems that the RFI source needs to be quite distant.

If the suspected RFI source transmits common signals to Green Bank and Haswell, then it would seem likely that some of the anomalous associated tones seen at Haswell should appear at Green Bank. However, none of the eight signals of the four two tone pairs exceeded the SNR threshold at Green Bank. There is a possibility that the suspected signal source has a design that results in one pulsed signal appearing at both sites, and many pulsed signals not appearing at both sites. This possibility leads to the idea that the transmitter uses an antenna with multiple beams in its pattern.

The Haswell and Green Bank computers running the FFTs do not communicate with each

other, and are not connected to the Internet. Standalone computers, connected to the Green Bank Forty Foot Telescope and Haswell Sixty Foot Telescope, observe the same location in the sky, have local oscillators and have ADC trigger time synchronized by GPS satellites, yet the systems otherwise measure telescope signals separately. However, the system behavior appears to be unusually correlated between Green Bank and Haswell. To make the problem more inexplicable, in the same sky pointing direction, a simultaneous Green Bank&Haswell pulse 0.2 Hz offset from each other, occurred the next day, shown in the plot of August 15 and 16 pulses on the previous slide. The problem appearing in these slides, and other observed anomalies, lead to the hypothesis stated below in ***Conclusion from pre-hypothesis observations.***



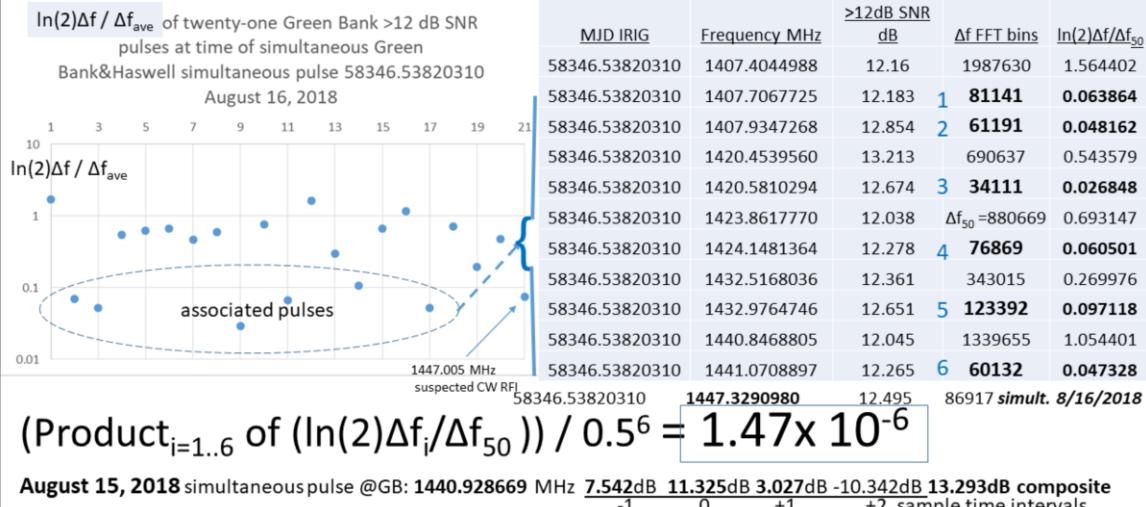
At the time 58345.5379225, a triplet was captured at Haswell having three tones spaced by 51,286 and 458 bins, from 1425.997654 to 1426.188709. These are raw data, not Doppler corrected frequencies. The noise likelihood of this and the other three close spaced tones is described on the previous slide. The driving interest in these two tones is that the tones' capture time was within the pointing beamwidth of a simultaneous Green Bank Haswell pulse at a Doppler corrected Green Bank to Haswell measured frequency difference of 1.0 Hz.

Upon examining the Green Bank file, a close-tone-spaced triplet was found to occur at the same time, i.e. 58345.5379225. Statistical analysis indicates that the likelihood of seeing this Green Bank triplet in the presence of only noise is  $9.3 \times 10^{-5}$ .

There was suspected RFI observed at Haswell at 1425.995947 MHz. No RFI was seen at Green Bank between 1435 and 1436 MHz. There is a possibility that the first tone in the Haswell triplet, at 1425.997654 MHz, is due to the suspected RFI approximately 2 kHz lower in frequency. The other two tones in the Haswell triplet do not correlate with known or suspected RFI.

The existence of the two triplets further seem to correlate Haswell and Green Bank measurements, countering attempts to falsify the working hypothesis.

Normalized Noise-Likelihood calculations of Green Bank associated pulses observed **at the same 0.25 s interval time**, 58346.5382031 MJD, i.e. the time of the **August 16, 2018 Green Bank & Haswell Δf=0.2 Hz simultaneous pulse**



August 15, 2018 simultaneous pulse @GB: 1440.928669 MHz  $\frac{7.542}{-1}$  dB  $\frac{11.325}{0}$  dB  $\frac{3.027}{+1}$  dB  $\frac{-10.342}{+2}$  dB  $\frac{13.293}{+2}$  dB composite sample time intervals

In the 0.25 s interval time of the 0.2 Hz offset simultaneous Green Bank and Haswell pulse(s), a high rate of doublets and triplets were observed at Green Bank. The median  $\Delta f$  in likelihood calculations was obtained from the frequency offsets thought to be due to noise. The average  $\Delta f$  during the four hour run was used in the plot above, after suspected RFI was excised. Almost all of the RFI that was excised in post-processing were suspected RFI tones close in frequency to a subset of harmonics of 1 MHz. The table above lists tones that are thought to comprise the doublets and triplets for which likelihood calculations have been performed.

Haswell doublet and triplet in August 16, 2018  $\Delta f = 2.7$  Hz simultaneous pulses

**Noise likelihood  $\approx 0.000398$**  = Binom.dist(1 in 9 trials @ 0.0036) \* 0.0417 \* 0.1531 \* 2 combinations

time MJD	frequency MHz	t-2 SNR dB	t-1 SNR dB	SNR dB	t+1 sNR dB	comp. SNR dB	denom. bins	num. bins	Pr. = ln(2) Δf/Δf <sub>50</sub>
58346.61199940	1407.80719150	2.395	3.438	11.257	-3.249	12.498	1880533	215194	0.4531
58346.61199940	1407.81349840	0.51	-2.982	11.482	-1.977	12.13	215194	<b>1693</b>	<b>0.0036</b>
58346.61199940	1409.83315330	-3.362	6.469	11.332	4.528	<b>13.289</b>	1693	542147	1.1416
58346.61199940	1411.45493240	-7.382	-3.789	12.534	-0.44	12.884	542147	435343	0.9167
58346.61199940	1419.76942640	3.468	1.3	11.895	-4.385	12.88	1073232	476058	1.0025
58346.61199940	1436.13574800	-0.326	-8.894	11.222	2.754	12.093	1073726	303503	0.6391
<b>58346.61199940</b>	<b>1436.40655060</b>	-4.564	6.034	11.423	2.756	<b>13.038</b>	303503	<b>72693</b>	<b>0.1531</b>
58346.61199940	1436.48034860	3.032	-11.591	12.323	0.835	<b>13.089</b>	72693	<b>19810</b>	<b>0.0417</b>
58346.61199940	1442.14771460	-2.098	1.874	11.486	-9.713	12.133	19810	1521322	3.2035
58346.61199940	1445.09125280	-3.125	4.204	11.258	1.347	12.515	1521322	790150	1.6639
58346.61199940	1447.40744830	6.028	-6.921	11.254	-6.148	12.504	513699	108050	0.2275
Green Bank	58346.61199940	58083.33375170	0.00450	1.62 ° RA		12.606	132642	2.7	
Haswell	58346.61199940	58083.33375170				<b>13.038</b>	72693	2.7	
Green Bank	58345.78468750	58083.50417475				12.225	166709	3.7	
Haswell	58345.78468750	58083.50417475				<b>13.096</b>	1046784	3.7	

*center tone of Haswell triplet is  
simultaneous 2.7 Hz off at Green Bank*

## November 25-August 16 summary of simultaneous pulses

Radio Telescope Location	Recorded IRIG MJD	IRIG Ref. to 58082.9919705	MJD time diff.	RA Angle Diff.	Likelihood simulated	SNR/bin /s	Num bins	$\Delta f$ Hz
<u>November 25 &amp; 26, 2017</u>								
Green Bank	58082.99197050	58082.99197050		0 ref ° RA		13.92	21626	1.5
Haswell	58082.99197050	58082.99197050				12.186	19844	1.5
Green Bank	58083.04392	58082.04406584				12.516	673691	3.7
Haswell	58083.04392	58082.04406584				12.449	1879417	3.7
<u>April 10 &amp; 11, 2018</u>								
Green Bank	58219.42857350	58082.80211656				13.003	547800	1.4
Haswell	58219.42857350	58082.80211656	.99197050	1st try to match		12.716	650423	1.4
Green Bank	58219.61602720	58082.99008348	-0.00189	-0.68 ° RA	1 in 93+4 trials	12.469	1715899	4.3
Haswell	58219.61602720	58082.99008348	.9919705	2nd try to match	to match any RA twice in April 11 range	12.941	1104425	4.3
<u>August 15 &amp; 16, 2018</u>								
Green Bank	58345.17852430	58082.89635196				12.352	40796	-3.4
Haswell	58345.17852430	58082.89635196				13.116	215602	-3.4
Green Bank	58345.61024020	58083.32924984		0 ref ° RA		12.305	573351	-2.5
Haswell	58345.61024020	58083.32924984				12.001	183726	-2.5
Green Bank	58346.23070310	58082.95141147				12.312	781038	-2.0
Haswell	58346.23070310	58082.95141147				12.239	1145777	-2.0
Green Bank	58345.28731770	58083.00544322				12.245	80565	-2.0
Haswell	58345.28731770	58083.00544322				12.425	56068	-2.0
Green Bank	58346.53820310	58083.25975336		0 ref ° RA		12.495	86917	0.2
Haswell	58346.53820310	58083.25975336				12.335	87197	0.2
Green Bank	58345.53806130	58083.25687332	-0.00288	-1.04 ° RA	1 in 107+5 trials	13.293	2131096	1.0
Haswell	58345.53806130	58083.25687332				12.75	340430	1.0
Green Bank	58346.61199940	58083.33375170	0.00450	1.62 ° RA		12.606	1328642	2.7
Haswell	58346.61199940	58083.33375170				13.038	72693	2.7
Green Bank	58345.78468750	58083.50417475				12.225	166709	3.7
Haswell	58345.78468750	58083.50417475				13.096	1046784	3.7

In **November 25-August 16 summary of simultaneous pulses**, the post-processing results are summarized to post the apparent received signals that matched at Green Bank and Haswell. Three sky pointing directions each have a pair of simultaneous pulses, at 0.68, -1.04 and 1.62 degree difference in pointing Right Ascension between the repetitions. Frequency offset ranges from -4.3 Hz to +2.7 Hz. The 1.62 degree RA pointing difference is significantly outside the approximate 1.0 degree antenna beamwidths, and is not considered as significant as the other two repetitions of simultaneous pulses.

**9,951 11/25-8/16 simulated trials required**

**HIP 24472 73 ly red dwarf 0.3° offset**

Summary of simultaneous pulses observed,  
including October-November 2018 observations

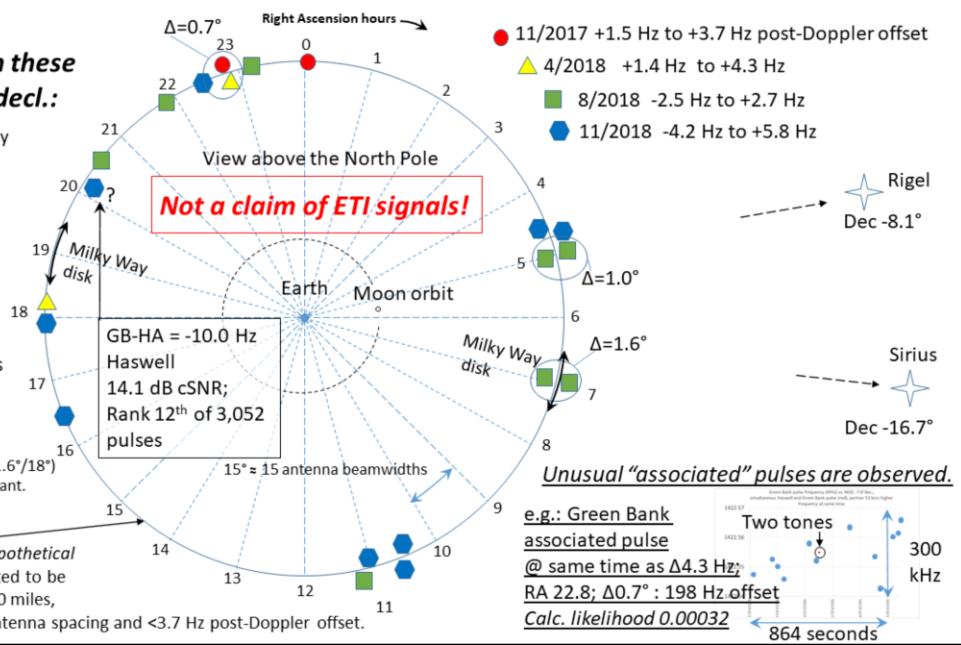
**So far,  
we have seen these  
pulses, -7.6° decl.:**

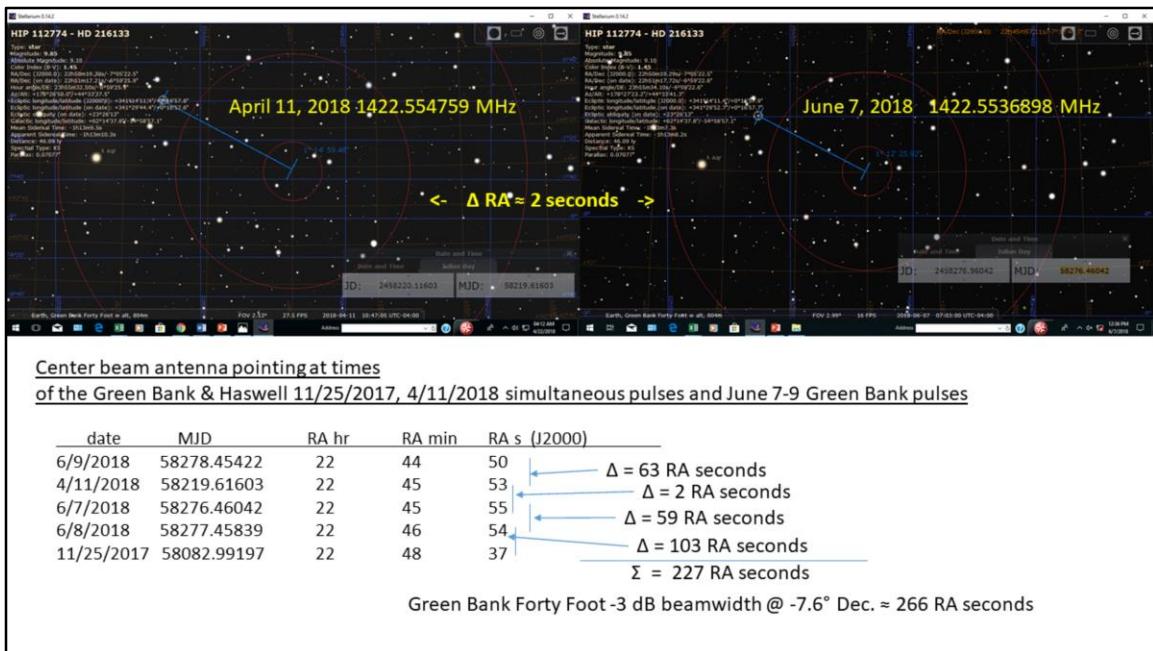
Twenty close-frequency same-time pulses have been observed during 102 hours of observation at -7.6° Declination, 11/2017 to the present.

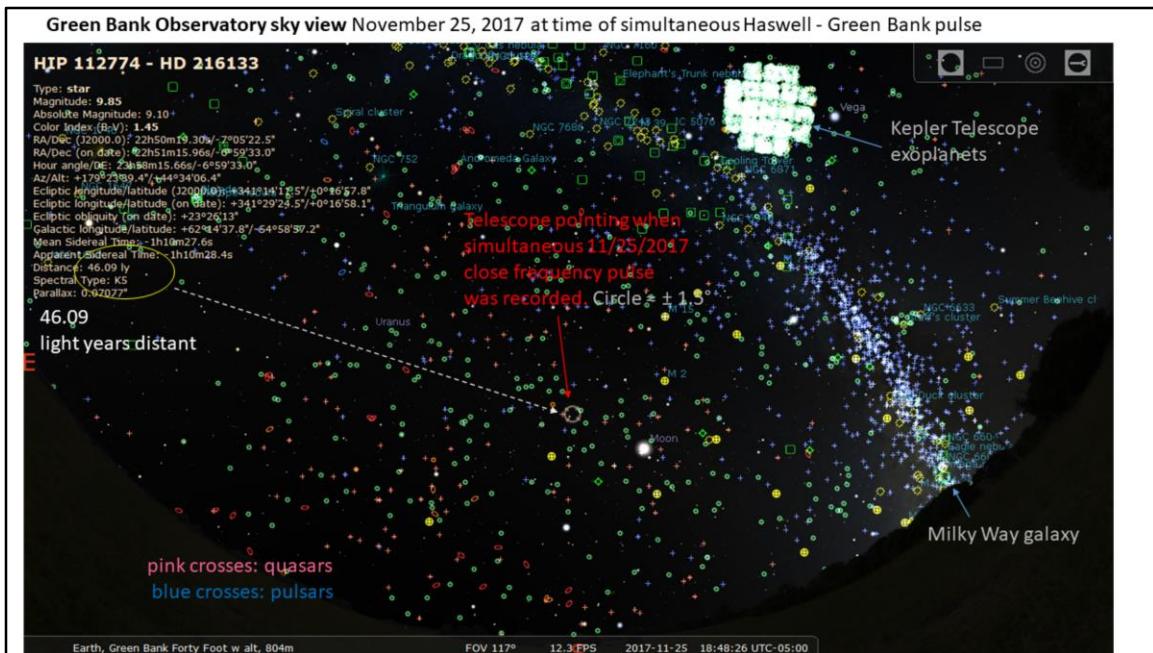
Three pairs of pulses measured having ΔRA pointing  $\leq 1.6^\circ$  offset.

Binom.dist(3,19 tries, pr.  $1.6^\circ/18^\circ$ ) =0.15, not by itself significant.

The distance to a *hypothetical* transmitter is expected to be greater than 720,000 miles, due to 1,257 mile antenna spacing and <3.7 Hz post-Doppler offset.







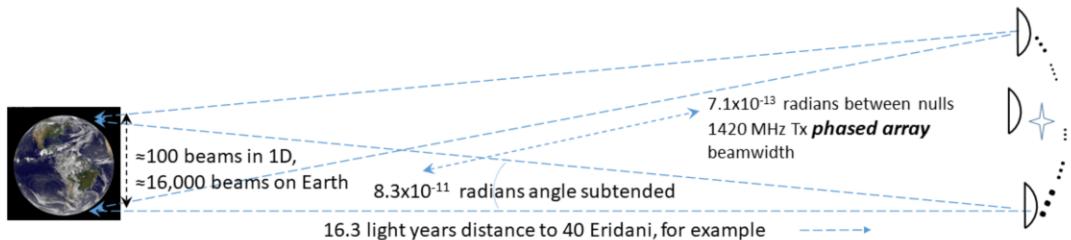
Unexpected observations compel a hypothesis,  
leading to an experiment to falsify, not “prove”.

Summary of hypothesis:

Anomalous **telescope-to-telescope similar** “simultaneous” pulses are predicted.

Anomalous **telescope-to-telescope different** “associated” pulses are predicted.

**Idea:** ETI may transmit certain types of high capacity signals to Earth  
**to compel human cooperation and readily indicate transmission motive.**



# Conclusion from pre-hypothesis observations

*The observed anomalies compel a hypothesis with tests.*

## **Working Hypothesis:**

1. Close-time ( $\pm 0.1\text{s}$ ) close-frequency ( $\pm 5 \text{ Hz}$ ) 1404-1445 MHz narrowband ( $<5 \text{ Hz}$ ) pulses are predicted to be observed at distant-separated ( $>1000 \text{ km}$ ) radio telescope sites.
2. Anomalous pulses, associated with pulses 1., are predicted to be observed close in frequency to each other, and close in time ( $\pm 200 \text{ s}$ ) to the observation of the simultaneous close frequency pulses 1.
3. The simultaneous pulses 1., and associated pulses 2., are predicted to be observed at an indication rate at least two orders of magnitude higher than the rate due to the likelihood of pulses caused by known noise processes.
4. The simultaneous pulses 1. are predicted to have multi-site measurement attributes explained by a distant ( $>10^6 \text{ km}$ ) source.
5. The associated pulses 2. are predicted to have multi-site measurement attributes explained by a distant ( $> 10 \text{ ly}$ ) large ( $>1 \text{ AU}$ ) phased array.
6. The pulses are predicted to be observed to repeat when pointing in fixed ( $\pm 1 \text{ degree RA Dec.}$ ) celestial pointing directions.

**Known alternate hypotheses:** Noise sources, natural objects, terrestrial RFI and reflections, clandestine satellites, equipment characteristics, calculation errors, hoaxes.

## SETI simultaneous telescope improvements underway

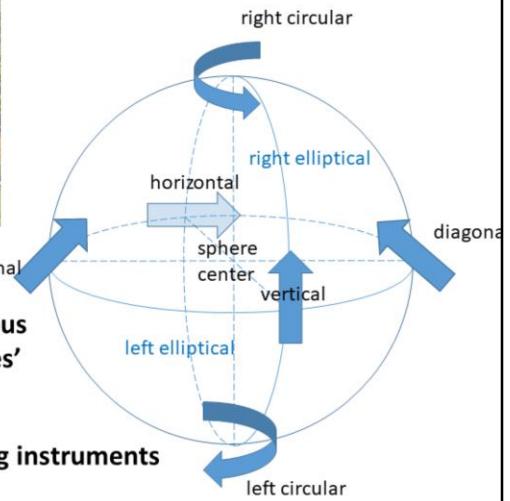
### More telescopes:

A third  
simultaneous  
telescope under  
construction  
in New Hampshire



### More pixels per antenna

Measure simultaneous  
and associated pulses'  
wave polarization



### More computers, faster computers, better measuring instruments

## Plans for system improvement

- Add computer communication between sites? *Risk: corrupting data*
  - Store raw data at both sites if SNR threshold is simultaneously crossed in Doppler corrected frequency window and/or close tones seen – increases post-trigger sensitivity by ~10x.
- Follow up HIP 112774 and 40 Eridani with low  $\Delta f$  & low SNR squelch algorithms
- Verify Doppler calculations
- Post-process raw data to search at lower SNR; cross-correlate; measure time delay
- Calculate likelihoods using apriori and aposteriori scenarios
- Non-contiguous triplet search
- Long term unattended simultaneous searches
- Adjust declination during observation to “hop” to nearby stars
- Telescope metrology, pointing accuracy, beamwidth, sensitivity
- Polarization measurements and Poincare Sphere search
- Two sites: ~one same-T same-f per day → **Three sites  $\approx 10^6$  days @ 12 dB threshold**
  - **Third site plan: ~continuously observing, eight meter paraboloid antenna under construction in NH**

**Thank you** (It is difficult to acknowledge all who helped me and gave me advice)

Steve Plock, Ed Corn, Rich Russel, Pasha Roberts, Paul Berge, Glenn Davis  
Adam Beardsley of ASU, Richard Prestage of GBO,  
Jay Lockman of GBO, Ron Maddalena of GBO, Ken Kellerman of NRAO, Duncan Lorimer of WVU,  
Scott Ransom of UVA and NRAO, Dan Reichart of UNC, Sandy Weinreb of Caltech,  
Jon Richards of SETI Institute, Jill Tarter of SETI Institute, Robert Gray,  
and many other very helpful friends.  
Online Arbitrary Precision Calculator :Thanks Mikko Tommila  
Equipment vendors and SW and HW engineers  
Thanks to folks of Green Bank Observatory, Breakthrough Listen,  
Berkeley SETI Research Center, SETI Institute, Deep Space Exploration Society  
and Society of Amateur Radio Astronomers folks, hams, friends and family.

**Thank you for listening and your advice during preparation of ideas.**

Reference: "Getting Ready for SETI", SARA Journal, December 2017

**Thank You!**

**Questions?**

Other topics, further detail, artifacts

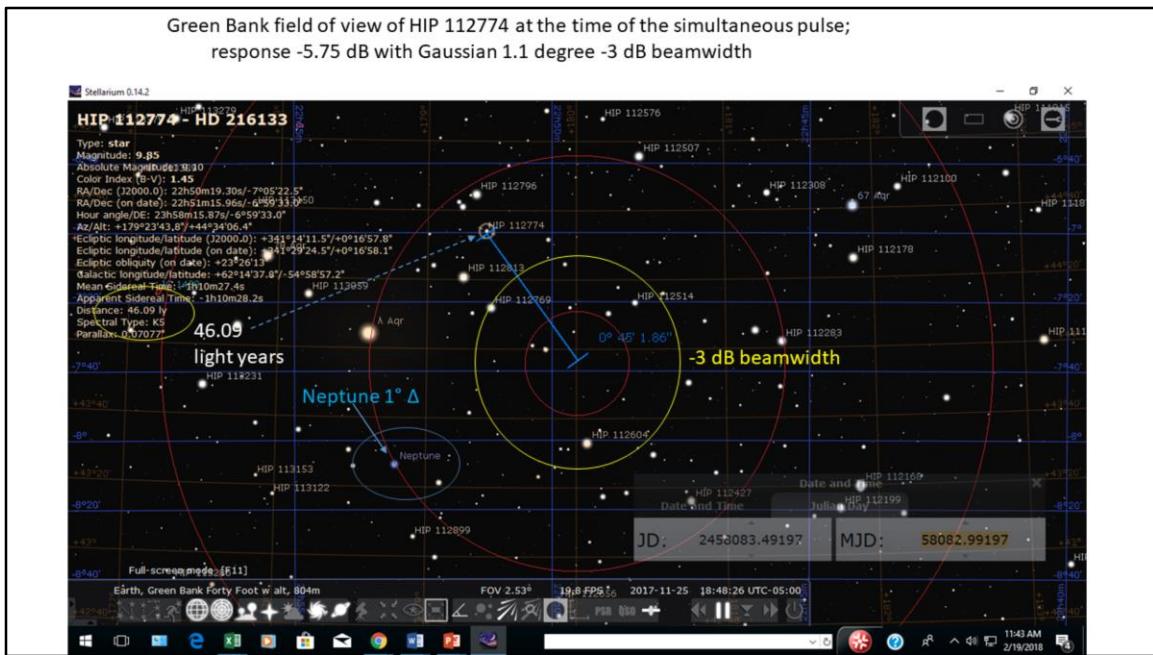
**8.2 meter diameter paraboloid, on Az-El mount, under construction in New Hampshire**

Dual polarized  
five pixels  
Az-El mount



Where were the telescopes pointing  
when the simultaneous pulse occurred?

Are there nearby stars or objects in the direction?  
(< 100ly)



Green Bank field of view of HIP 112774 at the time of the April 11, 2018 simultaneous pulse; response -15.8 dB with Gaussian 1.1 degree -3 dB beamwidth



# April 11, 2018 Green Bank sky view at time of simultaneous 4.3 Hz offset pulse



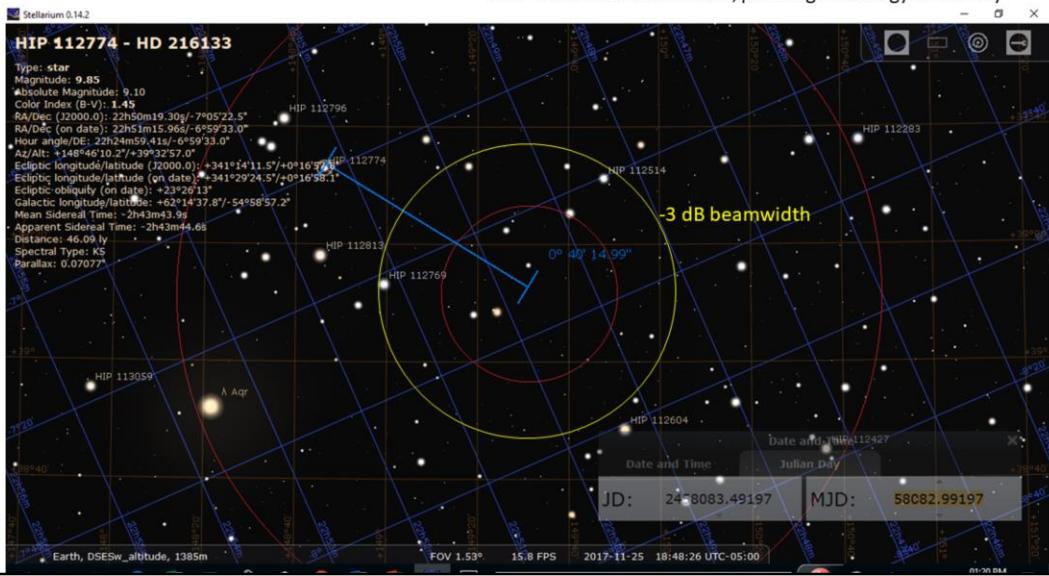
Green Bank HIP 112774 on Stellarium during closest MJD 58082 transit;  
 GBO Forty Foot pointing is 36 arc minutes south ~½ FHPBw  
 HIP 112774 is the 805<sup>th</sup> closest star of 119,615 stars in HYGV3 database.

GBO 40 Foot has  
 ≈18,000 1.5° non-overlapping  
 pointing directions



Haswell DSES field of view at 58082.99197; at occurrence of simultaneous  $\sim$ 1 Hz offset pulse

0.86° measured beamwidth; pointing metrology underway



June 7-9 , 2018 Green Bank 40 Foot standalone  
close frequency pulses

## Green Bank

April 11, 2018 (simultaneous w/ Haswell)

**58219.6160272 MJD**

**1422.554759 MHz**  $\Delta f = 1.0692 \text{ kHz}$   
**12.021 dB**

## Green Bank

June 7, 2018 (GB 40 Foot standalone)

**58276.4604225 MJD**

**1422.5536898 MHz**  
**12.571 dB**

June 7, 2018 Green Bank low  $\Delta f$  pulses, at same time as June 7, 2018 1422.5536898 MHz pulse

MJD	Frequency	composite 1s SNR	Likelihood=		
58276.46042250	1413.5711245 MHz	<b>13.863 dB</b>	Den( $\Delta$ bins)	Num( $\Delta$ bins)	Num+Den( $\Delta$ bins)
58276.46042250	1413.7422457 MHz	12.799 dB	958612	45935	1004547
58276.46042250	1413.7898698 MHz	12.229 dB	45935	12784	58719
58276.46042250	1413.8526000 MHz	12.048 dB	12784	16839	29623
58276.46042250	1413.9137693 MHz	12.763 dB	16839	16420	33259
58276.46042250	1414.0745975 MHz	12.072 dB	16420	43172	59592

**Product of Likelihoods =  $2.2 \times 10^{-8}$**

## Green Bank

April 11, 2018 (simultaneous w/ Haswell)

**58219.6160272 MJD**

**1422.554759 MHz** Δf = 22.82 kHz

**12.021 dB**

## Green Bank

June 8, 2018 (GB 40 Foot standalone)

**58277.4583912 MJD**

**1422.577580 MHz**

**12.065 dB**

June 8, 2018 Green Bank low Δf pulses, 0.75 s before June 8, 2018 1422.57758 MHz pulse

MJD	Frequency	composite 1s SNR	Den (Δbins)	Num (Δbins)	Num+Den(Δbins)	Likelihood ≈ $\ln(2)^*\text{Num.}(\Delta \text{ bins}) / 479423$
58277.45838250	<b>1444.2654788 MHz</b>	12.168 dB				
58277.45838250	<b>1444.6393527 MHz</b>	12.695 dB	349050	100361	449411	0.145
58277.45838250	<b>1444.6530879 MHz</b>	<b>13.556 dB</b>	100361	3687	104048	0.005
58277.45838250	<b>1444.7444357 MHz</b>	12.004 dB	3687	24521	28208	0.035
58277.45838250	<b>1444.7740406 MHz</b>	12.426 dB	24521	7947	32468	0.011
58277.45838250	<b>1444.8213480 MHz</b>	<b>13.079 dB</b>	7947	12699	20646	0.018
58277.45838250	<b>1444.9525341 MHz</b>	12.736 dB	12699	35215	47914	0.050

Product of Likelihoods =  $1.7 \times 10^{-9}$

June 9, 2018 Green Bank observations

MJD	Frequency (MHz)	1s Composite SNR(dB)	Den ( $\Delta$ bins)	Num ( $\Delta$ bins)	Likelihood = $\ln(2) * \text{Num.}(\Delta \text{ bins}) / 507447$
58278.45421590	1411.2440325	12.127 dB	1558035	1460262	
58278.45421590	1411.2445056	12.437 dB	1460262	<b>127 bins</b> rank 10/64376 = 0.00016	0.00017
58278.45421590	1411.2869218	12.596 dB	127	<b>11386 bins</b> rank 974/64376 = 0.015	0.016
58278.45421590	<b>1429.8427790</b>	12.228	1601523	1239291	<b>Product of Likelihoods = <math>2.7 \times 10^{-6}</math></b>

$\Delta f = 23.455 \text{ kHz}$ ; expect  $\approx 2.5 \text{ MHz}$

November 25, 2017 Green Bank partner pulse;  $\Delta f=1590.7 \text{ Hz}$ ;  $\Delta t=8.0\text{s}$  from Green Bank & Haswell simultaneous pulse

58082.99187790	1429.536273	12.739 dB
58082.99187790	1429.671632	12.304 dB
58082.99187790	<b>1429.866234</b>	12.221 dB

*What was telescope pointing  
during the simultaneous November 25, 2017, April 11, 2018 and June 7-9 events?*

## File records of the April 11, 2018 simultaneous pulse

=====

**Green Bank** 3.1 58219.61847 **58219.6160272** -52.81719 23650 6054497 **1422.554759 MHz** -53.45 -12.056 -22.148

**12.021 dB** = 0.26 second SNR **2.204 dB**

**12.469 dB**=one second composite **17.199 dB**=three tone triplet composite 431036 1715899 2.69951 2.37878 3.11432

**ratio = 3.98087166733173099230690707968**<sub>7e+00</sub> <no 4 5 digits

2 5.3 **563**<sub>70361505369432941860001498e+01 16 1.8298728 **365**<sub>05673582862742297708e+23 4</sub> s **0.0000043 MHz**</sub>

ratio = 3.98087166**733**1730992306907079687079**5014**<sub>8015478985513970990822112306164682300318302879573863899999072</sub>  
*Pr(digits a,b ∈ {0..9} ∈ { 34 digits of ratio }) = 0.0228*

**Haswell** 3.2 58219.61732 **58219.6160272** -69.19056 23651 6054673 **1422.555415 MHz** -69.09 -1.73 -0.252

**11.315 dB** **6.563 dB**

**12.941 dB**=one second composite **17.378 dB**=three tone triplet composite 1938800 1104425 2.64559 2.75343 2.69951

**ratio = 5.69643593975655044357334433670**<sub>3e-01</sub> <no 1 2 8

**digits** 4 1.76763694269789817606**733**5706809e+00 8 5.85699657711311722342559352924e+00 5 s **0.0000043 MHz**

ratio = 5.6964359397565504435**733**44336703**1153**<sub>29069527542809985558077161130596245100061893955023726016092428306e-1</sub>  
*Pr(digits a,b,c ∈ {0..9} ∈ { 31 digits of ratio }) = 0.00189*

