

BR1202.line.mom0.fits—raster

0.000709457 km/s

# DSES Science Meeting

February 24, 2020

Dr. Rich Russel

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J2000 Declination

25".9

25".8

25".7

25".6

25".5

25".4

7°24'25".3

1

0.8

0.6

0.4

0.2

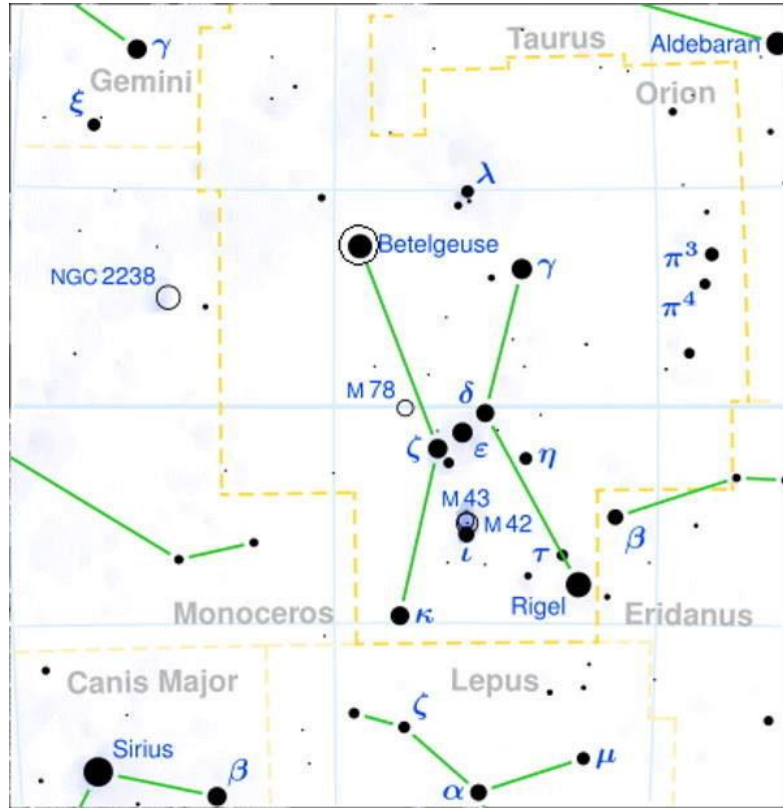
0

(Jy/beam)  $\times 10^{-4}$

# Agenda

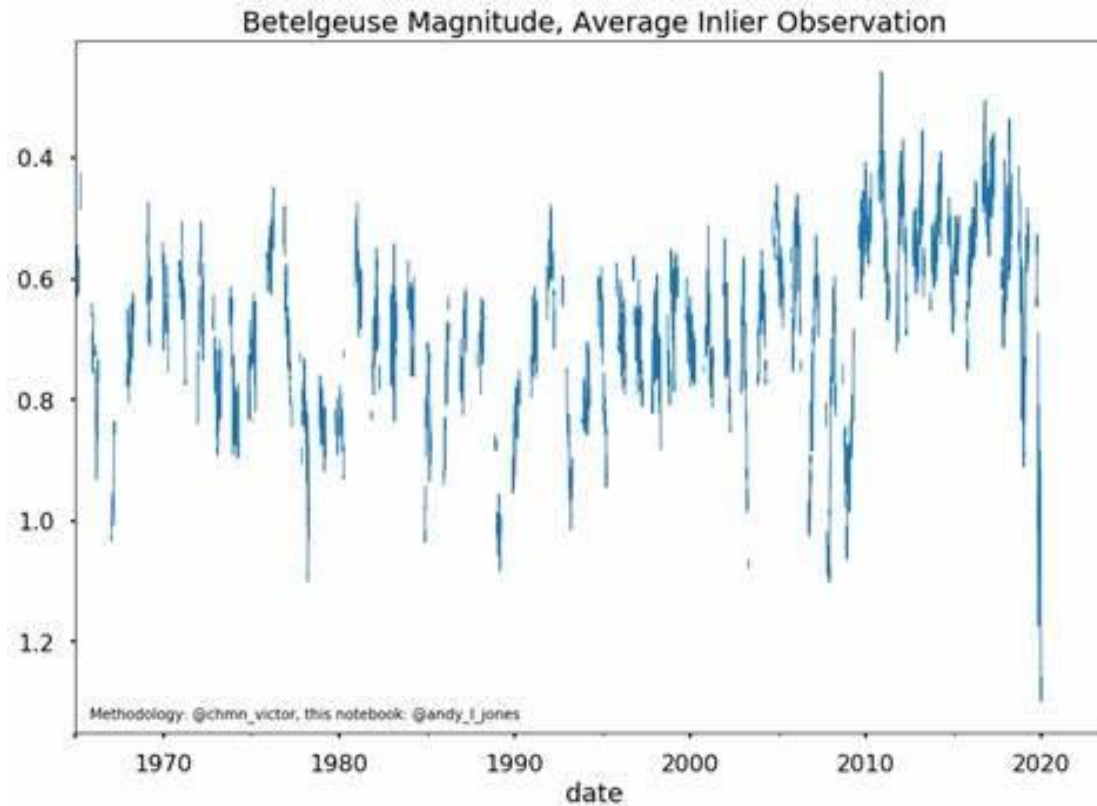
- Betelgeuse dimming experiment
- Pulsar Observation status
- Latest DSES papers and presentations

# Betelgeuse (Orion Constellation)



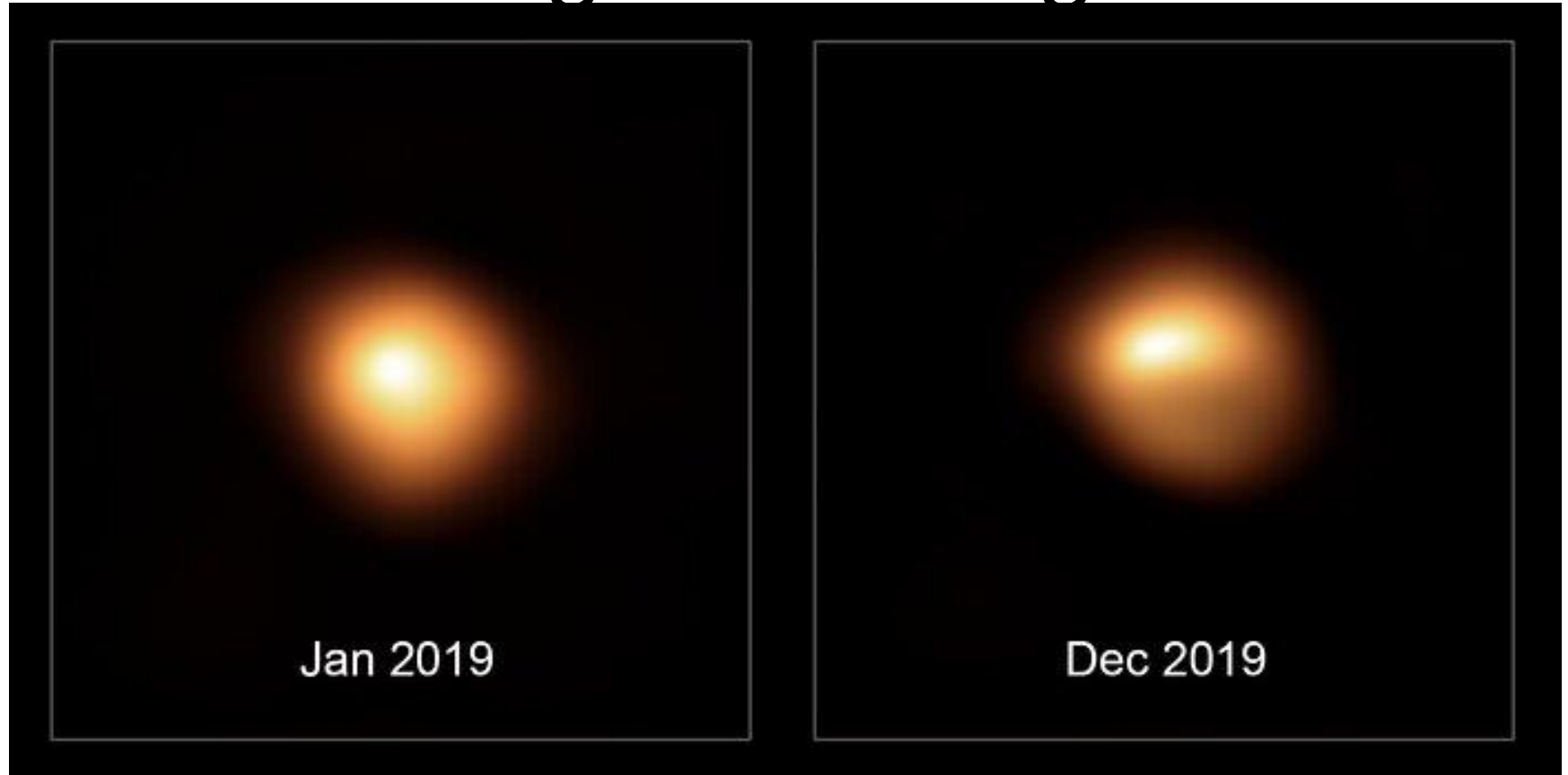
Source: Wikipedia

# Magnitude Variation



Source: <https://www.severe-weather.eu>

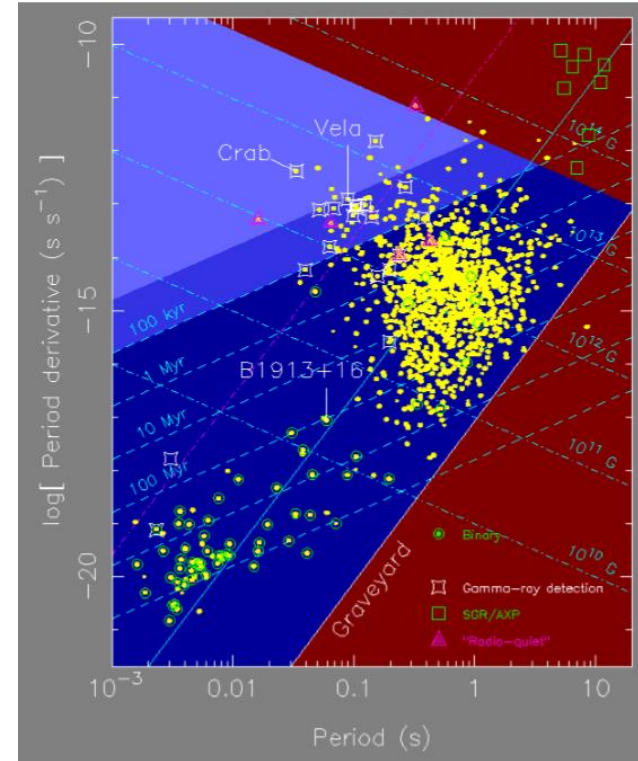
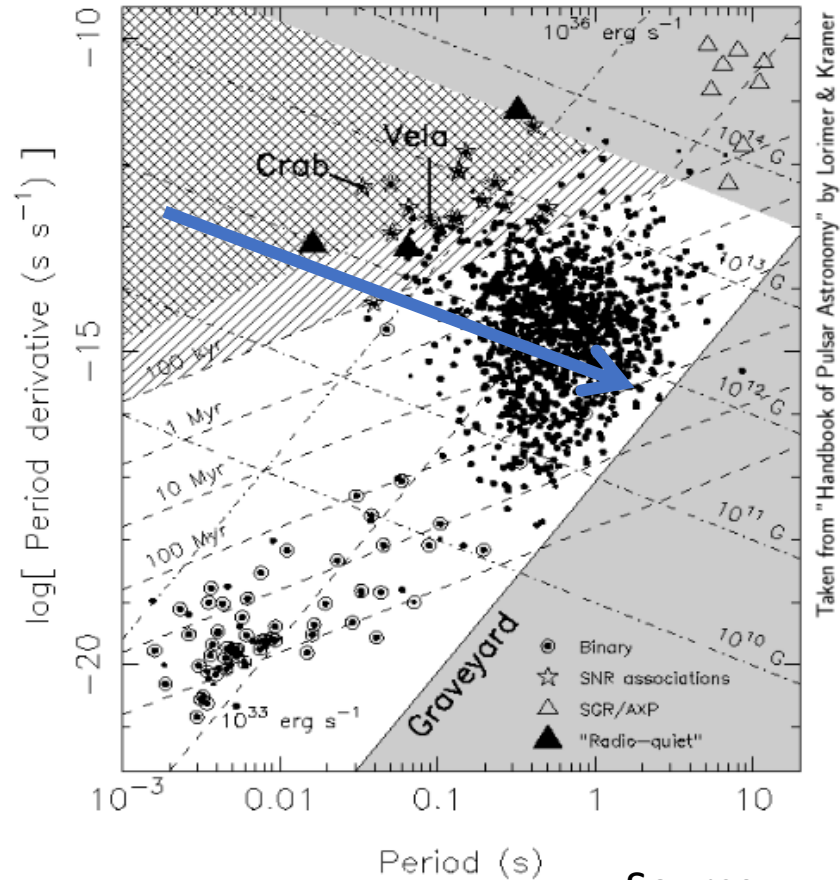
# Betelgeuse Images



Source:

<https://media.breitbart.com>

# Life Prediction

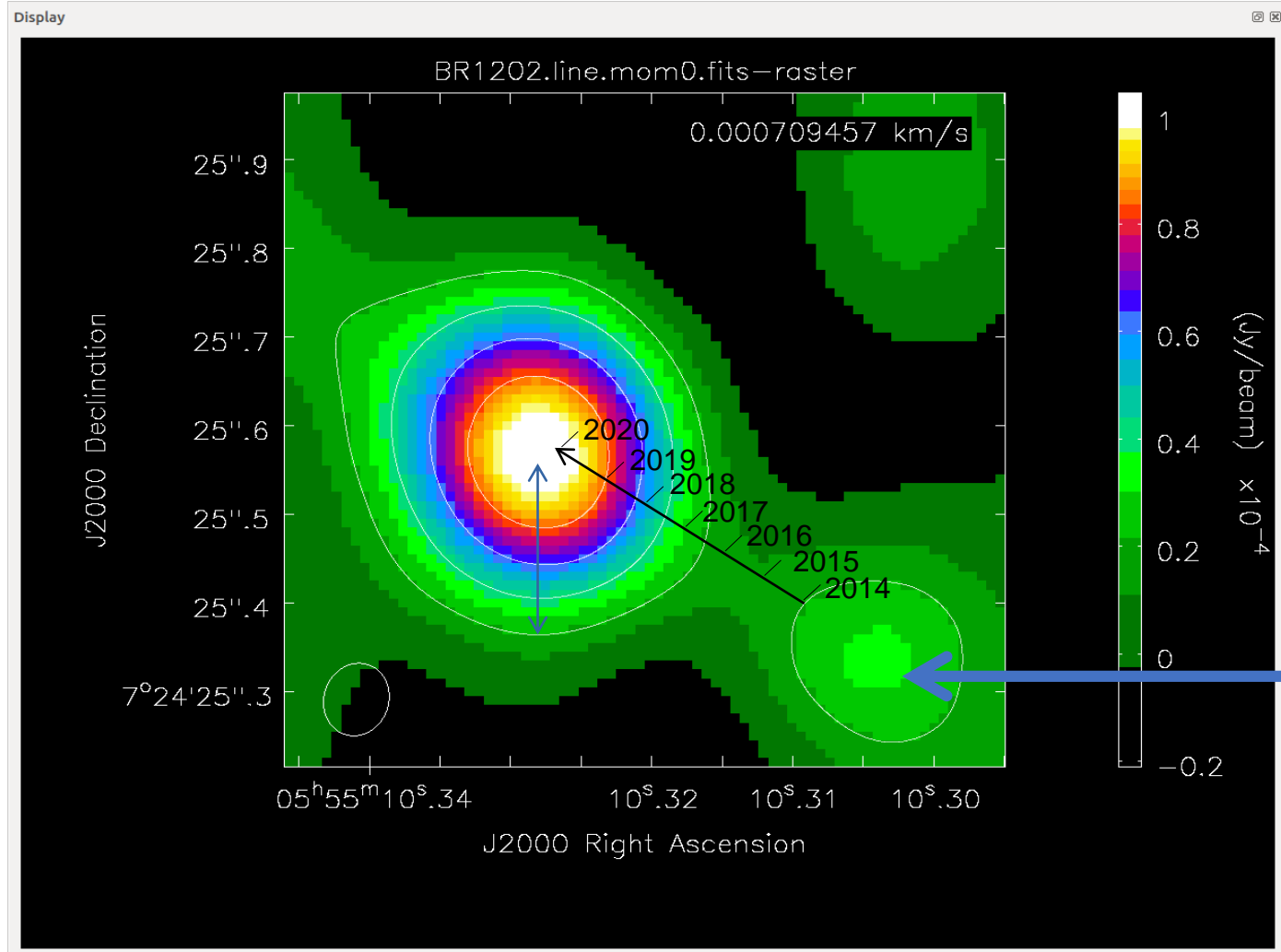


Source: [www.Wikipedia.org](http://www.Wikipedia.org)

# Quick Speed estimate

- Radius of Betelgeuse –  $R_{\text{Jupiter-Sun}} = 778,547,200\text{km}$
- Measure distance from Edge of cloud to center of Betelgeuse
  - $= 1.4 \times R_{\text{Jupiter-Sun}} = 1,089,966,080\text{ km}$
- Date of Image: 27 Feb 2014
- Estimate (t) seconds = 6years =  $6 \times 365 \times 24 \times 3600 = 189,216,000\text{ seconds}$
- Estimated speed of cloud =  $R_{\text{Jupiter-Sun}}/t = 4.115\text{ km/s}$
- Well within speed possibilities!!!!

# Target Cloud?





# Betelgeuse (Alpha Orionis) Coordinates

## Basic data :

### \* **alf Ori** -- Red supergiant star

Other object types: \* (\*,AG,...), IR (EIC,IRAS,...), \*\* (\*\*,ADS,...), V\* (V\*,AAVSO), smm (JCMTSE,JCMSTF), s\*r (Ref), LP\* (Ref), FIR (Ref), UV (TD1)

ICRS coord. (ep=J2000) : 05 55 10.30536 +07 24 25.4304 (Optical) [ 9.04 5.72 90 ] A 2007A&A...474..653V

FK4 coord. (ep=B1950 eq=1950) : 05 52 27.79597 +07 23 57.8110 [ 9.04 5.72 90 ]

Gal coord. (ep=J2000) : 199.78723027 -08.95860566 [ 9.04 5.72 90 ]

Proper motions mas/yr : 27.54 11.30 [1.03 0.65 0] A 2007A&A...474..653V

Radial velocity / Redshift / cz : V(km/s) 21.91 [0.51] / z(-) 0.000073 [0.000002] / cz 21.91 [0.51]  
A 2005A&A...430..165F

Parallaxes (mas): 6.55 [0.83] A 2007A&A...474..653V

Spectral type: M1-M2Ia-Iab B 1989ApJS...71..245K

Fluxes (8) : U 4.38 [~] C 1966CoLPL...4...99J

B 2.27 [~] C 1966CoLPL...4...99J

V 0.42 [~] C 1966CoLPL...4...99J

R -1.17 [~] C 2002yCat.2237....00

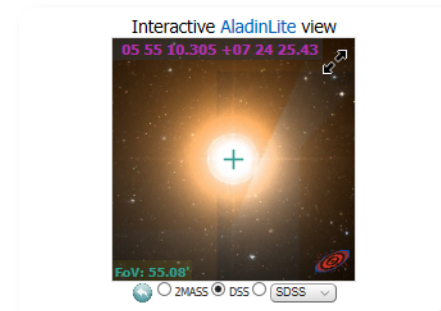
I -2.45 [~] C 2002yCat.2237....00

J -3.00 [~] C 2002yCat.2237....00

H -3.73 [~] C 2002yCat.2237....00

K -4.05 [~] C 2002yCat.2237....00

SIMBAD  with radius  arcmin



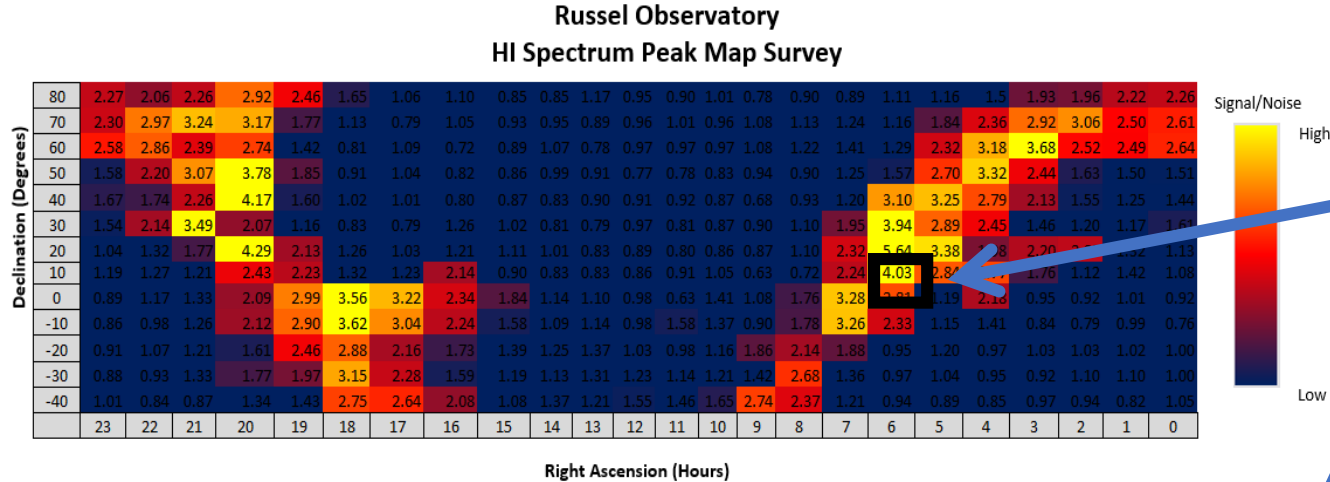
VizieR photometry viewer

Search within radius 1.0 arcmin

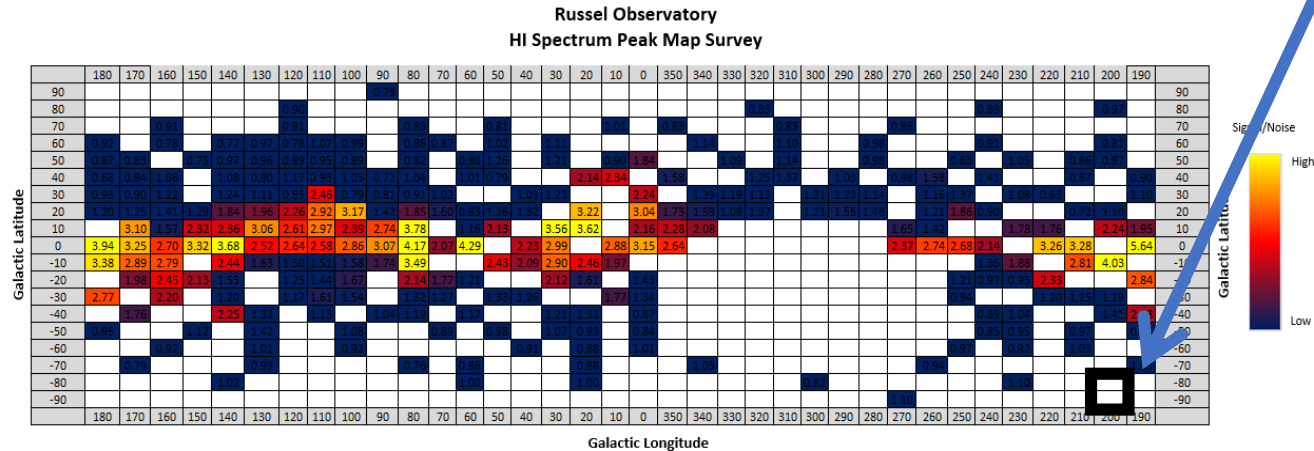
Source: <http://simbad.u-strasbg.fr/simbad/>

# HI Drift Scan using 9-ft Dish at Russel Observatory September 2019

Celestial  
Coordinates



Galactic  
Coordinates

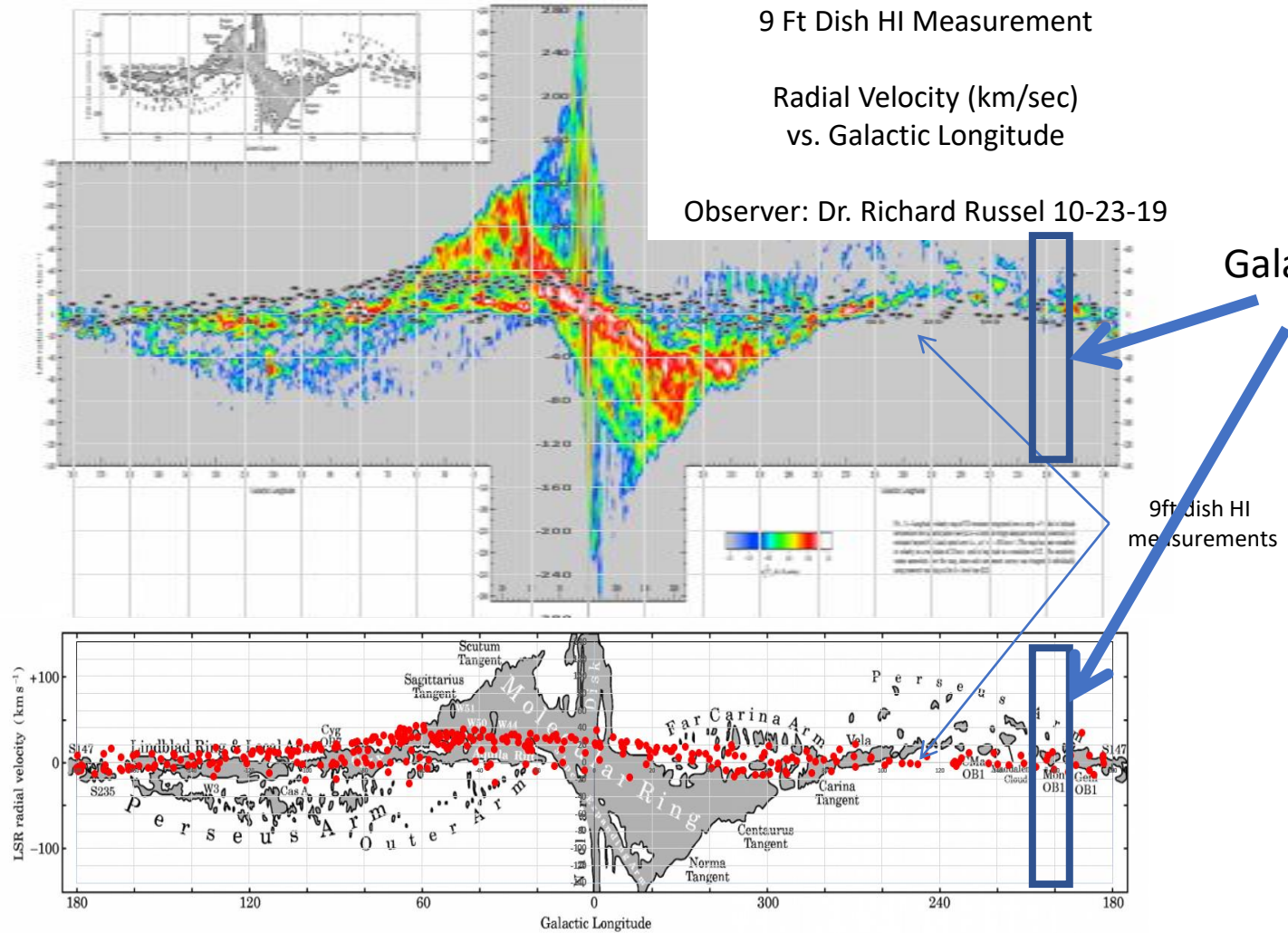


9 Ft Dish HI Measurement

Radial Velocity (km/sec)  
vs. Galactic Longitude

Observer: Dr. Richard Russel 10-23-19

Betelgeuse  
Galactic Longitude



Background images obtained at:

<https://www.britannica.com/place/Milky-Way-Galaxy/The-structure-and-dynamics-of-the-Milky-Way-Galaxy>

# Next Steps

- Find other images of Betelgeuse and see if “cloud” is visible and is moving
- Start taking H1 measurements of Betelgeuse with 9 ft and 60 ft dish – compare with baseline values from 2019

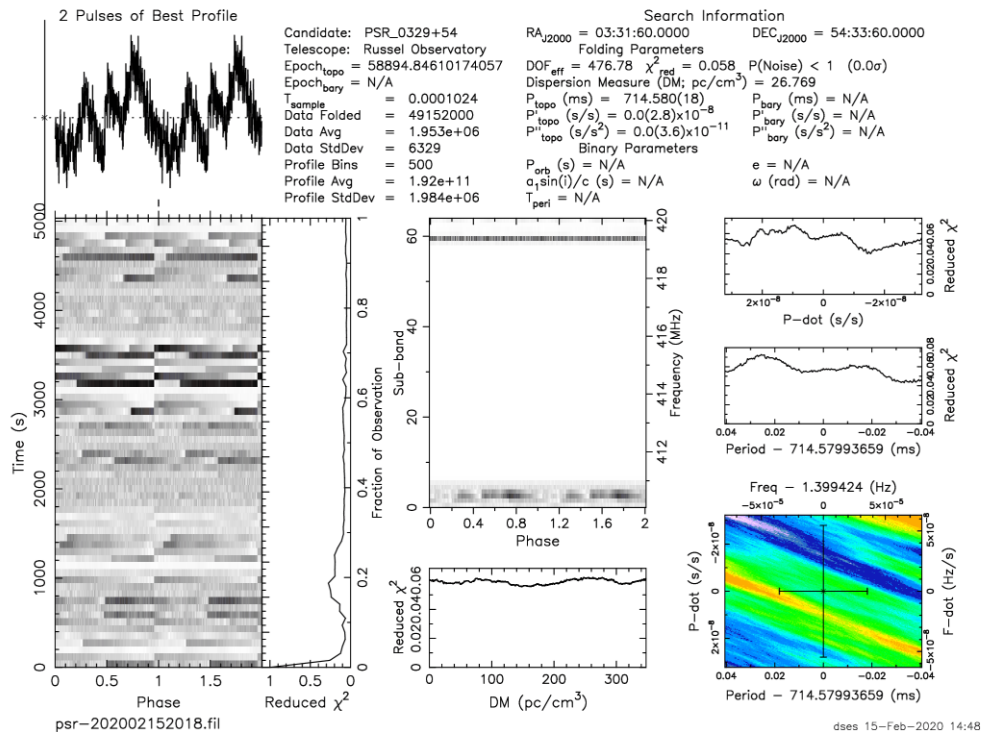
# Pulsar Observation Status

# DSES Observations

- Observed with 60 ft dish
- Results – no pulsar detected
- Possible issues:
  - Gain on antenna
  - Collection software
  - SDR settings
  - Processing approach

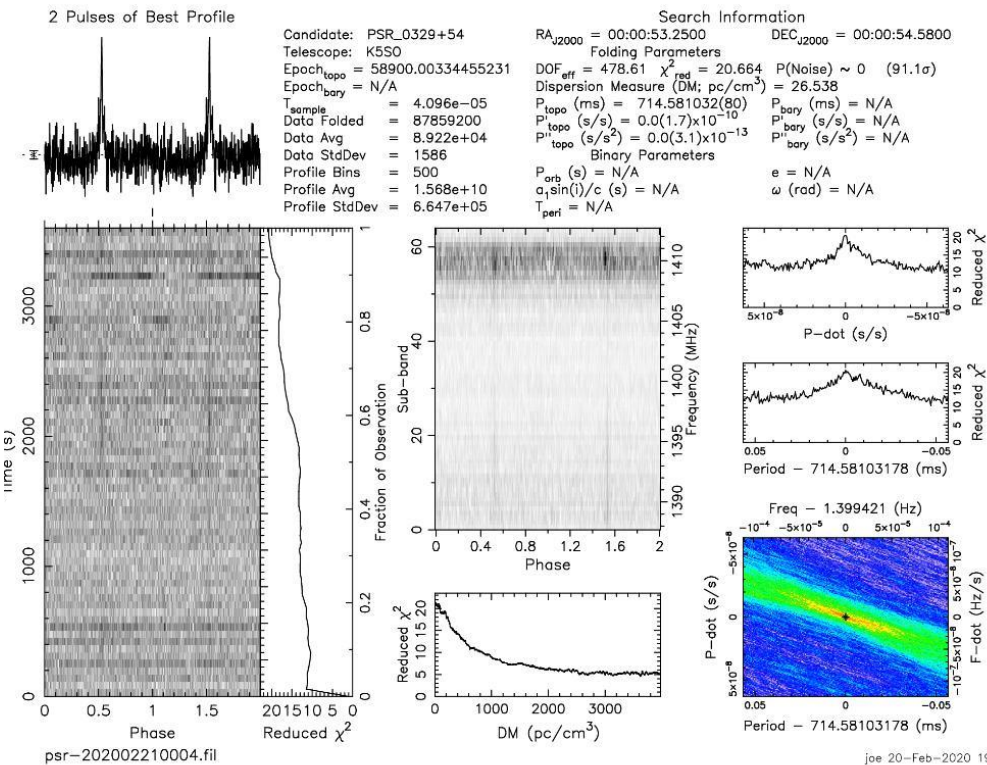
# DSES VS K5SO

## DSES



dses 15-Feb-2020 14:48

## K5SO



joe 20-Feb-2020 19:59

# Pulsar Next Steps

- Reinstall Yagi pulsar system
- Take SDR and laptop to Joe Martin's house after SARA Conference
  - Hook up to his antenna system and verify that the SDR/computer software works properly
- Haswell – bring some amplification for next try



# SARA West Papers

**Reducing Observations from the JVLA, ALMA, ATCA and EVLBI Radio-Telescope Archives**

**Dr. Richard A. Russel**

The large interferometry antenna systems maintain online archives of all of the observations conducted through their history. These observations can be processed (reduced) using CASA software. The results provide images of the astronomical objects as well as polarization, Jansky signal strength, relative velocity as well as chemical signature measurements. A number of archive files were downloaded and reduced. The analysis of the results, as well as the research on the astronomical objects, provides an excellent learning opportunity for the amateur radio astronomer.

# SARA West Papers

**Python Program for Mitigating Radio Frequency Interference Observed in SpectraCyber Receiver Drift**

**Scan Data Files**

**Jon Ayres**

**Deep Space Exploration Society**

The Colorado Springs Deep Space Exploration Society (DSES) is developing capability for collecting 21-cm HI Line drift scan data on the club's 60-foot parabolic dish antenna. Initial efforts have collected drift scans using a 9foot parabolic dish interfaced to a SpectraCyber I receiver. These initial scans show that HI Line data is being collected; however, the desired data includes corrupted samples caused by numerous Radio Frequency Interference (RFI) events and also exhibits signal level offsets thought to be due to component temperature changes. This initial progress report demonstrates and describes a method for automated removal of transient RFI events using a Python program that is part of the DSES drift scan development effort. This report also outlines future goals.

# Jan – Feb 2020 SARA Journal

**Alpha Centauri A, B & C Imaging using the ALMA Archive**

**Richard A. Russel**

**Deep Space Exploration Society**

**Abstract** The ALMA archives contain observations of the Alpha Centauri star system. Alpha Centauri is important because it is the closest star system to Earth and may have a planet that is in the habitable zone. This paper provides a basic analysis of some of these observations that were downloaded and reduced by the author.

Questions?